

FIELD, PETROGRAPHIC AND ⁴⁰Ar/³⁹Ar CONSTRAINTS ON THE TECTONIC HISTORY OF THE CENTRAL MANZANO MOUNTAINS, CENTRAL NEW MEXICO



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ABSTRACT

The Proterozoic rocks in the Capilla Peak area of the Manzano Mountains of central New Mexico preserve evidence, on all scales, for multiple deformation events. The dominant NNE trending foliation (S_2) overprints an older, nearly perpendicular foliation (S_1) . In schists and metarhyolites S_2 is a crenulation cleavage, while in quartz mylonites and amphibolites, S_2 is a mylonitic fabric. Within the amphibolites S_1 is preserved as a relict foliation subparallel to S_1 in the schist and metarhyolites.

Petrographic, microprobe and ⁴⁰Ar/³⁹Ar studies have been performed on rocks and mineral separates from the Capilla Peak area to help constrain the timing, and characterize, the deformational and metamorphic history in the Manzano Mountains. Quartz microstructures and crystallographic preferred orientations indicate that D₂ was accommodated at upper greenschist to lower amphibolite facies conditions, in a non-coaxial, general shear zone. This estimate of metamorphic grade is refined on the basis of petrographic studies and microprobe data from amphibolites. These data indicate that two distinct periods of amphibole growth occurred, resulting in an older (D₁) actinolite and a younger (D₂) hornblende. The hornblende indicates that (D₂) took place under lower amphibolite facies conditions. The younger hornblendes and biotites yield ⁴⁰Ar/³⁹Ar cooling ages of 1.4 Ga indicating that the area cooled through the hornblende and biotite closure temperatures by 1.4 Ga. Muscovites yield ⁴⁰Ar/³⁹Ar ages between 1320 and

1400 Ma indicating that the crust cooled relatively rapidly to 250-300°C by 1.35 Ga.

Data presented support a model of deformation during upper greenschist to lower amphibolite facies conditions ca. 1.45-1.4 Ga.

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INTRODUCTION

In the last decade, work in the Proterozoic of the southwestern United States has focused on understanding the extent, the character, and the driving force behind 1.4 Ga tectonism (Bowring and Karlstrom, 1990; Karlstrom and Bowring, 1991; Bowring, 1991; Shastri and Bowring, 1992; Bauer and Williams, 1993). Many of the studies have concentrated on the numerous 1.4 Ga plutons and their metamorphic aureoles, while few have tried to understand the 1.4 Ga tectonism outside the aureoles. The central Manzano Mountains is an optimal location to study the effects of 1.4 Ga tectonism because of the excellent exposure and varying lithologies, including 1.4 Ga intrusives. The study area was chosen greater than 10 km from the margins of the Priest and Ojita granodiorites in order to characterize and date the metamorphism and deformation from an area which has not experienced the thermal and deformational effects of 1.4 Ga plutonism.

Due to ongoing research and abundant unanswered questions related to 1.4 Ga tectonism, this thesis is presented in the form of two separate papers which will be published in order to make the data available for others studying or interested in the Proterozoic of New Mexico.

Part I is a description of the structural relationships observed in the Capilla Peak area of the Manzano Mountains. The emphasis is on

understanding the character of the latest deformational event through detailed analysis of microstructures and the microfabrics preserved in the rocks.

Part II is a presentation of new structural, microprobe and ⁴⁰Ar/³⁹Ar data supporting a model for deformation and metamorphism at 1.4 Ga in the Manzano Mountains. This model is somewhat controversial, but is becoming more widely accepted as the idea of "anorogenic" plutonism at 1.4 Ga is being questioned.

This thesis is a contribution to the ongoing studies of Proterozoic crustal evolution. Additional information, even within the Capilla Peak area of the Manzano Mountains, is needed in order to fully understand the middle to late Proterozoic in the southwestern United States.

Part I

FIELD AND MICROSTRUCTURAL OBSERVATIONS FROM THE CAPILLA PEAK AREA IN THE MANZANO MOUNTAINS, CENTRAL NEW MEXICO.

ABSTRACT

Proterozoic rocks in the Capilla Peak area of the Manzano Mountains preserve evidence for two distinct phases of deformation including folding of a pre-existing foliation and two phases of mineral growth. Little information about the character of the first deformation remains due to extensive overprinting. However, microstructural and crystallographic preferred orientations provide valuable information about the deformational history of the second event and place constraints on the metamorphic grade and character of this deformation. Microstructures and fabrics are typical of those forming during an upper greenschist to lower amphibolite grade metamorphic event. Quartz c-axis preferred orientations display both slightly asymmetric and symmetric patterns with respect to foliation and lineation, which is interpreted to suggest that deformation occurred under a combination of coaxial and noncoaxial strain. Asymmetric quartz c-axis distributions and microscopic and outcrop-scale kinematic indicators from quartz mylonites, metarhyolites, amphibolites and schists are interpreted to suggest a dominant east-side-up sense of shear. Metamorphic conditions inferred from crystallographic and microstructural fabrics are interpreted to be upper greenschist to lower amphibolite during deformation in the area of Capilla Peak, which is consistent with amphibole composition (Part II). The rocks in the Capilla Peak area are proposed to have accumulated high strain (inferred from microstructures and c-axis patterns) in a zone of general ductile shear.

INTRODUCTION

The Manzano Mountains are an eastward-tilted fault block composed primarily of Proterozoic metasedimentary, metavolcanic, and plutonic rocks. The mountain range is approximately 70 km long and 15 km wide and defines the eastern margin of the Rio Grande Rift between the Manzanita Mountains and Abo Pass(Fig. 1-1). The study area is in the central Manzano Mountains, located between Capilla Peak and Trigo Canyon (Fig. 1-1). The Capilla Peak area of the Manzano Mountains consists of a structurally complex sequence of multiply deformed Proterozoic quartzites, pelitic schists, metarhyolites, and amphibolites (Fig. 1-2). The Sevilleta metarhyolite is believed to be the oldest rock in the sequence; five Rb-Sr results define an isochron and yield an age of 1700 ±58 Ma (Bolton, 1976) and a U-Pb zircon age of 1680 Ma is also reported (Bowring et al., 1983).

The Monte Largo shear zone (MLSZ) has been mapped in the Manzano Mountains (Bauer, 1988; Thompson et al., 1991), and was first recognized in Monte Largo Canyon by Bauer (1988) as a 2m wide ductile fault. Structural mapping at 1:12,000 and petrologic investigations from Trigo Canyon north to Capilla Peak have shown that the MLSZ, as observed in Monte Largo Canyon, is not one discrete shear zone; rather, is just one localized high strain zone

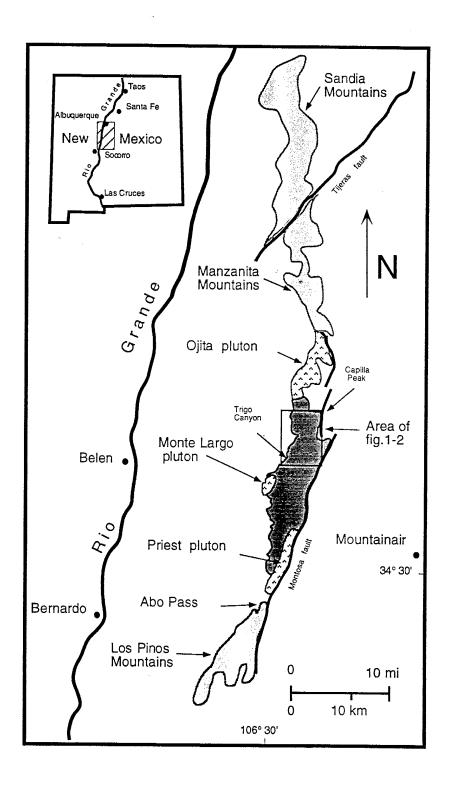


Figure 1-1. Geologic map of Proterozoic outcrops in central New Mexico showing the location of the study area.

within a larger deformation zone. Discontinuous zones of high strain appear to be lithologically controlled and have been mapped from Trigo canyon on the western edge of the range to Capilla Peak where they are overlain unconformably by Pennsylvanian limestones. The high-strain zones trend parallel, and adjacent to, the contact between the Blue Springs schist and structurally overlying Sevilleta metarhyolite (Fig 1-2). It is proposed in this paper that the Proterozoic rocks currently exposed in the Manzano Mountains were all part of a large ductile shear zone. The timing of this deformation is postulated to have occurred at 1.4 Ga based on 40 Ar/ 39 Ar geochronology of amphibole, muscovite and biotite from the Capilla Peak area (Part II).

The purpose of this paper is to describe the structural relationships and the deformational conditions of the 1.4 Ga event, in the Capilla Peak area of the Manzano Mountains. Our goal is to document these relationships and make interpretations in order to better understand the character and significance of the 1.4 Ga deformation event in New Mexico.

PREVIOUS WORK

The geology and structure of the Manzanita Mountains and the northern Manzano Mountains was first studied by Reiche (1949). In the Southern Manzano Mountains, Stark (1956) made correlations with observations of

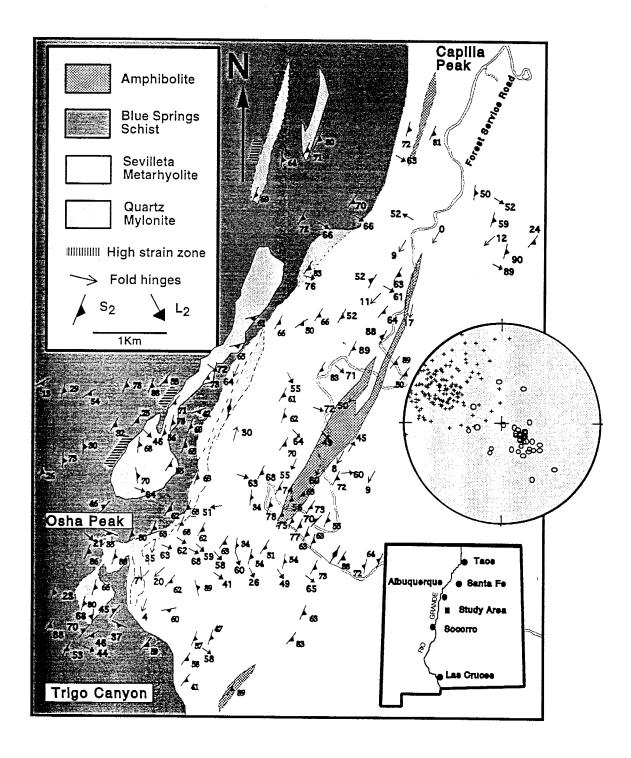


Figure 1-2. Geologic map of the study outlined in figure 1-1, showing the distribution of geologic units, fold hinges, the main foliation S_2 and the lineation L_2 . Poles to foliation (+) and mineral lineations (O) are plotted on a lower hemisphere equal area net.

Reiche's; and made a correlation between parts of the stratigraphic section in the Los Piños Mountains (Stark and Dapples, 1946) with that in the Southern Manzano Mountains. Stark (1956) suggested that the Precambrian rocks in the Manzano Mountains were only moderately metamorphosed and therefore assumed an upright stratigraphic sequence. Within this sequence, Stark (1956) interpreted the quartz layers within the Blue Springs schist as vein quartz. After 1956 little work was performed in the Manzano Mountains until Grambling (1982) produced a map of Precambrian structures in Trigo Canon. Grambling documented evidence for three phases of folding, disproving the idea that the original stratigraphic sequence was preserved. Bauer (1982) produced a detailed geologic map and structural interpretation of the southern Manzano Mountains. Like Grambling (1982), Bauer (1982) recognized three phases of folding in the Manzano Mountains: early isoclinal folding, a strong crenulation cleavage and third, open folds in quartz-rich rocks and small crenulations in shistose rocks. Bauer (1982) also observed crossbedding in the quartz layers within the Blue Springs schist disproving Stark's (1956) conclusion that the layers were veins. In Monte Largo Canyon, Bauer (1982) mapped the MLSZ as a probable fault, and Thompson et al. (1991) recognized this structure as a shear zone separating two different terranes based on contrasting metamorphic grade. She proposed that the MLSZ juxtaposed the greenschist grade Blue Springs schist from the overriding amphibolite grade Sevilleta metarhyolite. Northrup (1991) studied the geochemical fluid-rock

interaction within the MLSZ in Monte Largo Canyon and conclude that several episodes of hydrothermal veining had occurred.

The rocks present in the Capilla Peak area are similar to those described in the Los Pinos and southern Manzano Mountains by Stark and Dapples (1946) and Bauer (1982) and are summarized below.

Blue Springs schist

The Blue Springs schist of Stark and Dapples (1946) is intruded in the south by the 1656±10 Ma (Bauer et al., 1992) Monte Largo pluton and trends NNE for approximately 17 km, where it is overlain by Pennsylvanian limestones. The schist exposure is approximately 4 km wide, covered on the west by overlying pediment gravels and bound on the east by the White Ridge quartzite. Three dominant rock types exist within the Blue Springs schist: metasiltstones, phyllites, and muscovite-chlorite schists (Stark 1956; Bauer 1982).

The metasiltstones are typically folded, fine-grained pink, gray and green, finely-laminated rocks. S_1 is parallel to compositional laminae. Folds in S_1 are tight to isoclinal and slightly disharmonic. The phyllites are composed dominantly of quartz and muscovite and preserve a strong foliation S_2 . The muscovite-chlorite schists range from gray-green to reddish brown in color, and are found throughout the Blue Springs schist. Folds in S_1 in the schists include 0.5-4 cm crenulations and 1-3 cm chevron folds. Between 30 and 40% of

the muscovite-chlorite schists consist of folded and boudined quartz layers aligned parallel to the crenulated and chevron folded foliation S_1 in the schist.

Quartz mylonites

The quartz mylonites within the Manzano Mountains have previously been subdivided into the Sais and the White Ridge quartzites within the Blue Springs schist by Stark and Dapples (1946), Reiche (1949), Stark (1956) and Bauer (1982). The Sais quartzite does not outcrop in the study area. The quartz mylonites outcrop in discontinuous layers parallel to the foliation trend within the Blue Springs schist, whereas the White Ridge quartzite outcrops between the Blue Springs schist and the Sevilleta metarhyolite (Fig 1-2). Due to regional folding, compositional variations and the possibility of original cyclic deposition, it is unclear whether the quartz mylonites should be subdivided into separate formations. The White Ridge quartzite and quartz "reefs" in the field area are macroscopically and microscopically similar and will be grouped together as quartz mylonites for discussion in this paper. Quartz mylonite layers range from less than one meter to more than a kilometer in width. Compositions of the mylonites range from 90 to 99% quartz and 1-10% muscovite, with minor magnetite, biotite and epidote. Locally color ranges from pure white to purplish-gray depending on biotite and magnetite percentages. Outcrops range in character from massive to schistose.

Sevilleta metarhyolite

The Sevilleta metarhyolite of Stark and Dapples (1946), outcrops predominately on the east side of the range in the Capilla Peak area. It is a pink to brown, well foliated and lineated, metamorphosed rhyolite. In lower strain areas, the metarhyolite contains 1-2 mm equidimensional quartz and feldspar porphyroclasts. In the higher strain zones, porphyroclasts are elongate parallel to the foliation and lineation. The matrix is predominantly fine-grained quartz, feldspar and biotite with minor opaques.

Amphibolites

Amphibolite layers occur only within the Sevilleta Metarhyolite and are interpreted to be early mafic dikes (Stark and Dapples, 1946). The amphibolite layers are one to twenty meters wide and up to 400 meters long. The outcrop character ranges from highly schistose to massive. The amphibolites are petrologically complex and contain two amphibole phases, a hornblende and an actinolite, plagioclase, biotite, quartz, epidote and minor amounts of magnetite and pyrite (see Part II).

STRUCTURE

Macrostructural relationships

The dominant structure in the central Manzano Mountains is a pervasive NE-striking steeply SE dipping foliation (Fig. 1-2). The NE-striking foliation in

the schistose units is a crenulation cleavage, cutting a well defined earlier foliation (S_1) . In the non-schistose units, the main foliation (S_2) is defined by aligned minerals. S_2 is locally absent to well developed and is subparallel to lithologic contacts with the exception of areas of disharmonic folding where S_2 is variable. Where S_2 is absent to poorly developed, S_1 is the dominant foliation; this relationship is observed in several locations within the Blue Springs schist and in a few locations within the metarhyolites. The character of the macroscopic folding is integrally related to the degree of S_2 development and the rock type. Where S_2 is well developed, S_1 and S_0 are typically not observed.

The main foliation (S_2) in the Blue Springs schist exhibits the highest degree of variation, both in character and in orientation. S_2 in the Blue Springs schist is parallel to the regional foliation trend and in most places is a crenulation cleavage (Fig. 1-3). S_2 in the phyllites and the quartz-rich zones of the muscovite-chlorite schists is a well developed planar foliation, striking approximately 030° and dipping 60-80° SE. However, S_2 in the majority of the muscovite-chlorite schist is non-planar. S_2 anastamoses between the numerous folded quartz layers with

orientations ranging from 0-050°, 60-90° SE. A few areas within the schists exhibit an asymmetrical chevron fold crenulation with the limbs 1 to 3 cm in length (Fig. 1-3), and S_2 axial planar to the folds. The two limbs have the same strike and are oriented approximately 030°, dipping 85° SE and 55° NW with the main foliation subparallel to the SE dipping limbs. The metasiltstones are

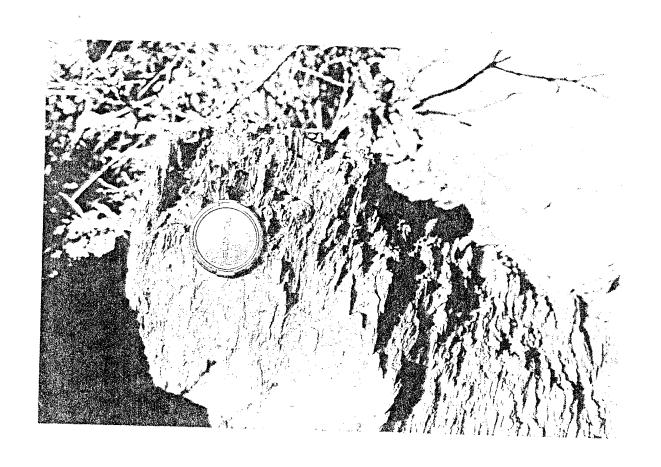


Figure 1-3. Asymmetric chevron folding within the Blue Springs schist. S_2 is nearly vertical and is a crenulation cleavage. The limbs of the crenulations are 1-3 cm in length. Long dimension of photograph is approximately 30 cm.

characterized by small disharmonic folding of the compositional layering. The main foliation in the metasiltstones is poorly developed and highly irregular in orientation. The foliation is axial planar to the disharmonic folding which may explain the large variation in orientation of S_2 .

A distinctive, narrow zone of highly deformed Blue Springs schist is seen in three locations. The most continuous outcrop is located on the west side of the contact between the ridge-forming quartz mylonite layer and the Blue Springs schist (Fig. 1-2). Two other highly deformed zones are located along strike within the schist, adjacent to one of the smaller quartz mylonite layers. The contact between the highly deformed zone and the schist is gradational. The character of the Blue Springs schist changes with increasing proximity to the zones; from containing abundant quartz-rich layers, to discontinuous and boudinaged quartz layers in the contact zone, to no quartz layers in the highly deformed zone. The highly deformed zones are less than a meter wide and defined by a well developed, highly planar foliation.

The Sevilleta metarhyolites have a moderate to well-developed NNE-striking foliation (S_2) and a locally well-developed down dip lineation. The foliation is defined by aligned biotite and lens-shaped quartz and feldspar porphyroclasts. A older foliation (S_1) is observed in a few locations where S_2 is poorly developed. No evidence is preserved for original compositional layering.

Folds in $\rm S_1$ within the metarhyolite include both small 0.5-1 m isoclinal folds and 3-5 m open folds in foliation. These folds plunge 0-50° to the SW and

are seen with a strong down dip lineation. Where crenulations are observed, S_2 is a crenulation cleavage. S_2 is axial planar to the outcrop scale folds, but is typically not observed in folded areas.

S₂ in the amphibolite layers is poorly to well-developed and is usually parallel to the foliation in the surrounding metarhyolite. One amphibolite layer contains distinctive 1-10 cm less deformed `pods' (Fig. 1-4). A well developed foliation wraps around these pods. The pods have length to width ratios ranging from 1:1 - 2:1 and are characterized by 1-3 mm, unaligned, green to black amphibole crystals surrounded by strongly foliated fine-grained amphibolite. The pods are sigmoidal in shape and asymmetric with respect to the main foliation, showing a consistent east-side-up sense of shear. Several 0.5-1 m scale asymmetric sigmoidally shaped structures within the amphiboite layers are also observed. These structures are similar to the 1-10 cm pods but contain an internal foliation parallel to, but weaker, than the surrounding foliation.

The quartz mylonites have a strong mylonitic foliation (S_2) oriented approximately 025°, dipping 60-80° SE (Fig. 1-5). In areas with a higher percentage of muscovite, the foliation is better developed.

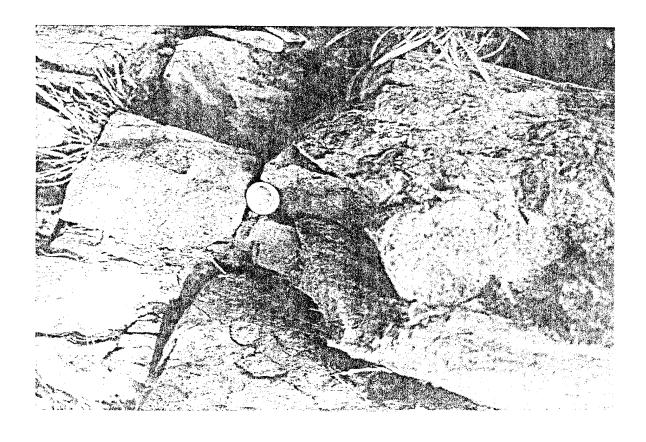


Figure 1-4. Photograph of asymmetric sigmoidal structure developed around the pods within the amphibolite layers. The pods are zones of relatively undeformed amphibolite within a strongly deformed amphibolite layer. Long dimension of photo is approximately 50 cm.



Figure 1-5. Field photograph of highly foliated quartz mylonite. The dominant foliation, S_2 , is nearly vertical.

Open to isoclinal folds in S_1 are common in the quartz mylonites, and locally the axial planar foliation S_2 is observed. Folds in the quartz mylonites appear to be preserved in the lower strain zones. The discontinuous quartz mylonite layers within the Blue Springs schist may be repeated by large-scale folding, however, no direct evidence for this has been observed in the Capilla Peak area.

Microstructural relationships

In the following section, the microstructural relationships related to the youngest deformation event are described.

Blue Springs schist microstructures

In thin section, the character of the Blue Springs schist is as variable as it is at the outcrop scale (Fig. 1-6). In general the S_2 crenulation cleavage is best developed adjacent to the quartz mylonites. With greater S_2 development the cleavage plane spacing is decreased. In zones with a higher amount of quartz, recrystallized grains preserve a grain shape preferred orientation at a 45° angle to S_2 .

Sevilleta metarhyolite microstructures

The metarhyolites are characterized by 1 to 2 mm quartz and feldspar porphyroclasts in a fine-grained (0.1-0.25 mm) quartz and feldspar matrix. The majority of the feldspar porphyroclasts are symmetric with respect to the

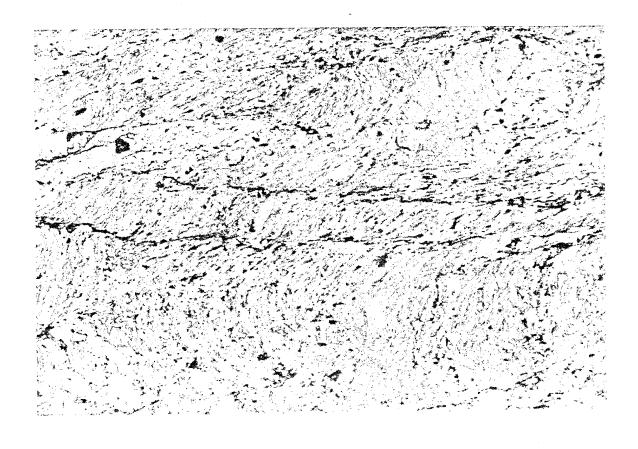


Figure 1-6. Photomicrograph from the Blue Springs. Long dimension of photo is parallel to S_2 and approximately 2mm.

foliation, however asymmetric delta-shaped porphyroclasts are present (Fig. 1-7). Asymmetries are also preserved by quartz grain-shape-preferred orientations within the matrix and porphyroclast tails, but are often ambiguous.

Amphibolite microstructures

The amphibolites exhibit the most complex microstructures in the study area and are dealt with in more detail in Part II. The amphibolites contain two chemically and structurally distinct amphibole phases, an older actinolite and a younger hornblende, plagioclase, biotite, quartz and epidote. The older anhedral actinolite is cross-cut and overgrown by the foliation-forming euhedral hornblendes (Fig 1-8). The actinolites contain inclusion trails that mark an older foliation parallel to cleavage plane traces within the actinolites and nearly perpendicular to the well defined hornblende foliation. The amphibole phases are chemically distinct (Table 2-1 Part II) and it is proposed in Part II that the younger hornblende crystals grew at lower amphibolite facies conditions, probably around 500°C.

Quartz mylonite microstructures

Quartz mylonites from the central Manzano Mountains exhibit a variety of microstructures which can be divided into three groups. Samples ML 11-3, ML 7-15 and ML 6-2 are characterized by 0.1 to 0.25 mm recrystallized quartz grains (Fig. 1-9). Quartz grains typically have a strong to moderate grain-shape preferred orientation (GSPO) which defines a foliation at an angle between 45°

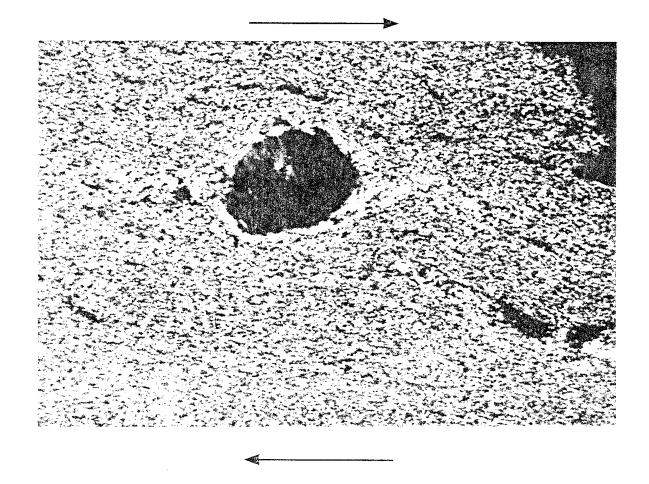


Figure 1-7. Photomicrograph from the Sevilleta Metarhyolite. Note the delta-type porphyroclasts asymmetry indicating an east-side-up sense of shear. Photomicrograph view is to the south, parallel to L_2 . Long dimension of photo is approximately 6mm.

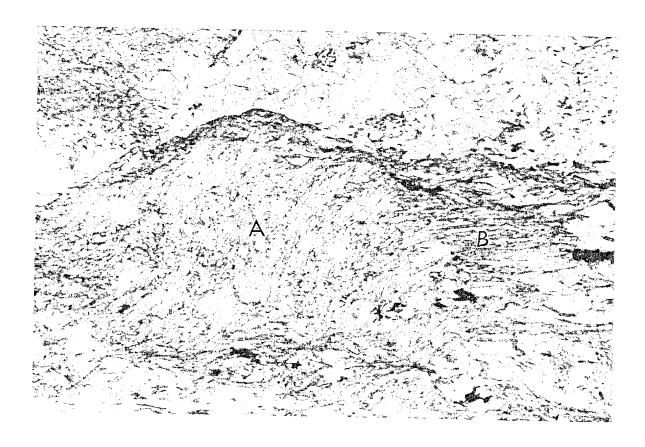


Figure 1-8. Photomicrograph from amphibolite layer showing two amphibole phases. Note the anhedral tan actinolite (A) has mineral inclusions nearly perpendicular to the main foliation defined by the green euhedral hornblendes (B). Long dimension of photo is approximately 6mm.

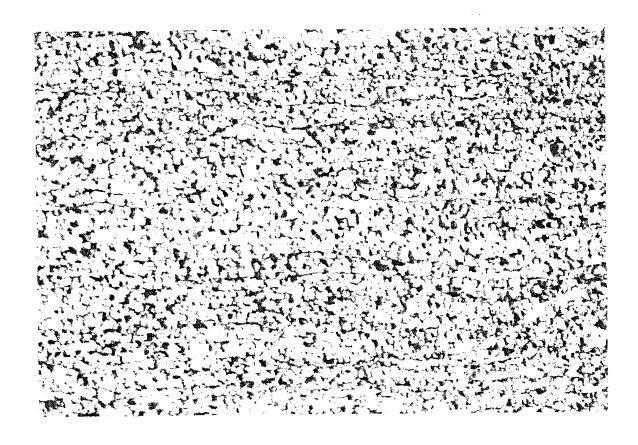


Figure 1-9. The first group of quartz microstructures is characterized by a fine-grained fabric with 0.1 to 0.25 mm equigranular quartz crystals. Note the foliation $\sim\!50^\circ$ from S_2 . View is perpendicular to S_2 and L_2 . The long dimension of the photo is approximately 6mm.

and 60° from S₂. The second group of quartz mylonites (ML 7-1 and ML 7-3) are characterized by fine-grained 0.1 to 0.25 mm recrystallized quartz grains with relict monocrystalline quartz ribbons (Fig. 1-10). Relict ribbons are 1 to 3 mm long, with a 4:1 length to width ratio. The third group (ML 1-2, Ml 5-10) is characterized by 3 to 20+ mm long monocrystalline quartz ribbons (Fig. 1-11). Ribbons exhibit undulose extinction with minor subgrain formation and the rocks locally exhibit S-C fabrics. In all mylonite samples, micas are included within the recrystallized quartz grains as well as between grains indicating a high degree of grain boundary mobility; this suggests either deformation at high temperature or low strain rate (Lister and Dornseipen, 1982).

c-axis crystallographic preferred orientations

In combination with the quartz microstructures, c-axis crystallographic preferred orientations can be a helpful tool for determining both the sense of shear in the mylonite and the conditions of deformation (cf. Hobbs, 1972). As with microstructural studies, several factors need to be taken into account when interpreting the observed c-axis crystallographic preferred orientations.

Temperature, strain rate, trace impurity content, type of strain (coaxial or non-coaxial) and OH content are factors which determine how quartz deforms therefore influencing the development of the crystallographic preferred orientation (e.g. Lister and Dornsiepen, 1982; Schmid and Casey, 1986). Kronenberg (1981) has proposed

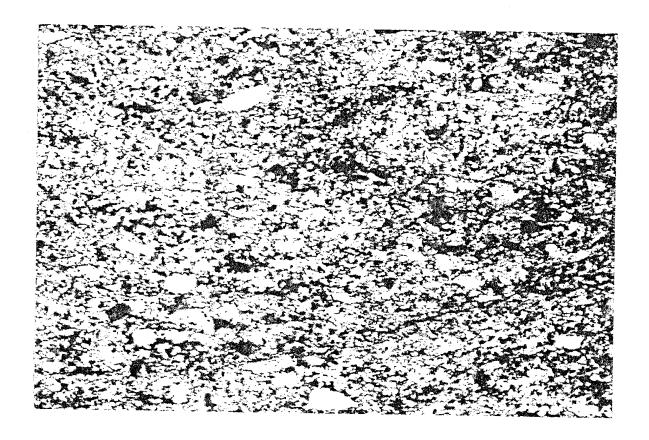


Figure 1-10. The second microstructural group is characterized by a bi-modal grain size distribution. The larger grains have sutured grain boundaries and are elongate parallel to the main foliation. The recrystallized grains are 0.1 to 0.25 mm in width. The fabric is relatively symmetric with respect to S_2 and L_2 . View is perpendicular to S_2 and L_2 . The long dimension of the photo is approximately 6mm.

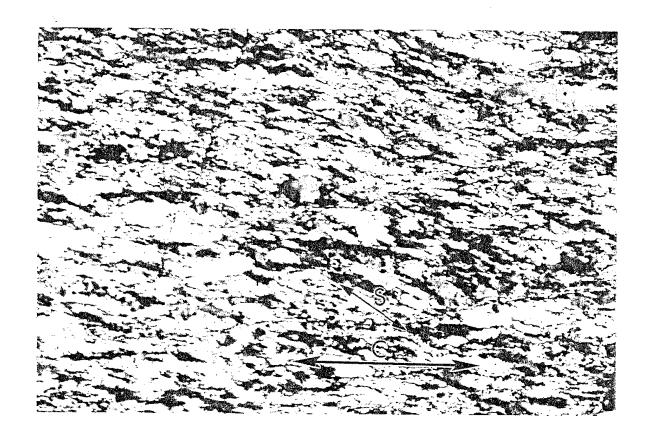


Figure 1-11. The third microstructural group is characterized by 1-10 mm slightly undulose monocrystalline quartz ribbons and <5% white mica. Ribbon boundaries are sutured and exhibit significant subgrain formation. These fabrics are typically asymmetric and exhibit S-C relationships. View is perpendicular to S_2 and L_2 . Long dimension of photo is approximately 6mm.

that the development of c-axis preferred orientation patterns in quartz is also strongly effected by the presence of phyllosilicates. In this situation slip may be preferentially accommodated by the phyllosilicates with optimal orientations. It has also been proposed by Lister et al. (1982) that increasing OH has the same effect on the fabric as does decreasing the temperature. It is very difficult to isolate the influence of each of the factors since all influence deformation (Schmid and Casey, 1986). However, the influence of temperature and strain rate on fabrics is better constrained. Experimental studies show that an increase in temperature or a decrease in strain rate corresponds to an increase in girdle angle (on the c-axis plot) and a difference in location of c-axis maxima (Tullis et al., 1973). The type of strain also has an effect on the fabric, for example small circle girdles are indicative of flattening, while cross girdles are often indicative of plane strain and similarly, single girdles indicate constrictional strain (Schmid and Casey, 1986). The above factors affecting c-axis patterns complicate identification of the active slip systems. This identification of active slip systems is necessary in order to interpret the deformation conditions. In general, TEM analysis of structures associated with slip is necessary to determine the active slip systems in naturally deformed rocks (e.g. Law, 1990; Ralser, 1990); however, similarities between experimentally and naturally deformed quartzites (Tullis et al., 1973) make it possible to use c-axis crystallographic preferred orientations to infer the relative activity of different slip systems. If the dominant slip systems can

be identified, it may be possible to work backwards and estimate temperature or strain rate during deformation (Lister et al., 1978). From c-axis fabrics it has been shown (Lister and Hobbs, 1977) that an asymmetry of the c-axis pole figures with respect to the foliation and lineation reflects the sense of shear that the sample has experienced. This asymmetry also indicates non-coaxial deformation. Caution must be used when interpreting sense of shear from asymmetric c-axis patterns in multiply deformed terranes. In such a situation, it may be difficult to distinguish whether the asymmetry results from non-coaxial deformation or from a pre-existing asymmetry (Ralser, 1990).

Description of Crystallographic Preferred Orientations

Two hundred c-axis orientations from each of the representative quartz mylonite samples were measured from thin sections cut perpendicular to the lineation and foliation (Fig. 1-12). Data were then plotted on equal area, lower hemisphere projections and rotated so that the trace of the foliation is horizontal and the lineation is horizontal and plunging to the right.

Three different patterns are observed in the quartz mylonite samples. Incomplete small circle girdles about the pole normal to the foliation are developed in three samples (ML 11-3, ML 7-3 and ML 7-15; Fig. 1-12). In each girdle, two maxima are symmetrically distributed on either side of the pole to the foliation; however, the size of the maxima on either side of pole to the foliation vary.

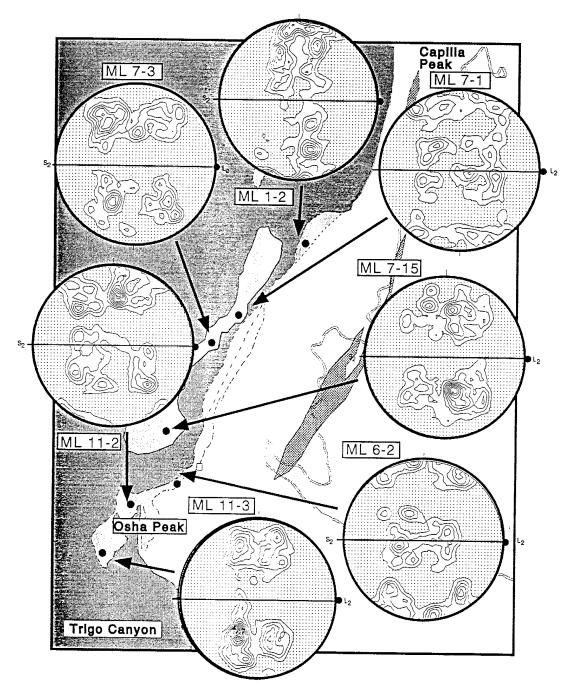


Figure 1-12. Geologic map of the study area showing the locations of quartz mylonite samples. Equal area, lower hemisphere plots of 200 quartz c- axes from each sample. Contours represent multiples of uniform distribution, from 1-10, area less than one multiple of uniform distribution is shaded (Starkey, 1977). S₂ is the trace of the main foliation and L₂ is the quartz mineral lineation on S₂. See figure 1-2 for map key.

C-axis patterns from samples ML 6-2, ML 7-1 and ML 11-2 (Fig 1-12) have a distribution of c-axis orientations that range from normal to the foliation to perpendicular to the lineation, within the foliation plane. The maxima in these three samples are relatively symmetric about the pole to the foliation, in both distribution and size. These plots are different from the first group of plots in that the maxima are located preferentially in the center and in the outer northern and southern portions of the plot.

Sample ML 1-2, (Fig. 1-12), has a different c-axis pattern than the other six samples. Maxima are located on both sides of the pole to the foliation as in the other samples, but they lie closer to the pole to the foliation and define a girdle approximately normal to the foliation.

DISCUSSION

The lithologic units in the study area each exhibit a different structural character on both the macroscopic and microscopic scales, as detailed in the previous sections. These differences in structural character are a result of different behavior during deformation. In a complexly deformed terrane with multiple lithologies, deformation will be accommodated in different ways. For example, a metarhyolite with a slightly higher percentage of phyllosilicates may appear to have accommodated more strain than a metarhyolite with fewer, even if the strain was the same. In the same scenario, a metarhyolite with a

higher percentage of phyllosilicates may be easier to deform and hence accommodate more strain. With this in mind it is important to interpret the structural character in the context of the lithology and composition.

Microstructural studies are used to help interpret the deformational history by placing sense of shear, relative temperature and/or strain rate constraints on the main deformational event.

The quartz microstructures are relatively similar, even though they have been divided into three groups. These fabrics appear to be recording the transition from a coarser-grained microstructure characterized by monocrystalline quartz ribbons to a fine-grained roughly equigranular microstructure. Sample ML 11-2, for example, appears to have the entire range preserved within one thin section. Zones of monocrystalline quartz ribbons and zones with fine-grained equigranular microstructure co-exist (Fig. 1-13). This is interpreted to suggest that the polycrystalline quartz ribbon is a recrystallized monocrystalline quartz ribbon texture. The microstructures are consistent with those described by Hamner and Passchier (1991) and Simpson, (1985) as forming during upper greenschist to lower amphibolite facies conditions with slow strain rates. Assuming that the original grain-size and mineralogy of the quartz mylonite protolith was similar everywhere, and that temperatures were likely similar, the differences in microstructure observed are likely the result of either strain variations, strain rate variations, or variations in

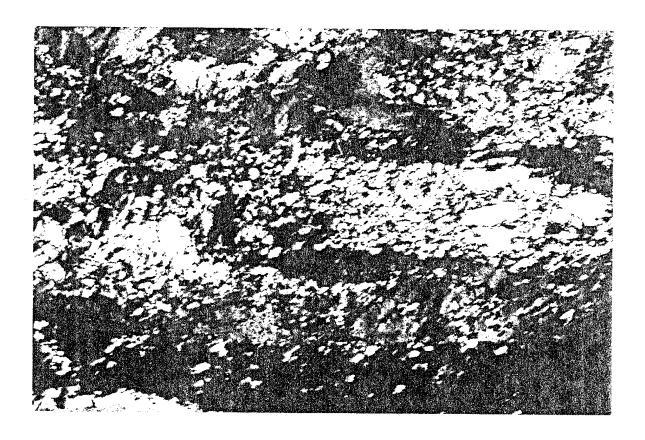


Figure 1-13. Photomicrograph of quartz mylonite sample ML 11-2, view is perpendicular to S_2 and L_2 . Grain-shape preferred orientations ~45° from the C-surfaces (S_2) are consistent with an east-side-up sense of shear. The long dimension of the photo is approximately 6mm.

composition and/or water content (Kronenberg, 1981). Such local variations could also be controlled by deformation zone width, contrasting rheology with surrounding protolith or differences in the strain (cf. White et al., 1980).

The presence of predominantly delta-type porphyroclast systems within the metarhyolites, as opposed to sigma-type porphyroclast systems has been proposed to suggest that both, the strain (rotation) rate was high relative to the recrystallization rate and that there was a component of extensional shear (Hamner and Passchier, 1991). Based on the presence of asymmetric fabrics and porphyroclast systems and the nearly orthorhombic symmetry of most quartz mylonite fabrics and c-axis patterns, we interpret that the observed deformation formed during progressive general non-coaxial flow.

The majority of the kinematic indicators in the Capilla Peak area such as GSPO, shear bands and rare asymmetric porphyroclasts in the quartz mylonites and schists and asymmetric porphyroclast systems in the amphibolites and metarhyolites indicate an east-side-up sense of shear. A few (<5%) asymmetric porphyroclast systems indicate a west-side-up sense of shear; the geologic significance of this is unknown.

Interpretations from Crystallographic Preferred Orientations

The c-axis orientations of samples ML 11-3, ML 7-3 and ML 7-15 resemble small circle girdles about the pole normal to the foliation (Fig. 1-12). If these patterns are small circle girdles or possibly a transition to type 1

crossed girdles (Lister, 1977), then they suggest predominantly flattening strain (Law et al., 1984). In comparison with the c-axis distributions presented by Schmid and Casey (1986), the c-axis patterns from samples ML 7-3, ML 11-3 and ML 7-15 could suggest that slip occurred dominantly on the rhomb planes in the <a> direction. Quartz microstructures are consistent with formation during upper greenschist to lower amphibolite conditions. Amphibolite facies conditions are unambiguously indicated by the composition of syntectonic hornblende (Part 11). However at these conditions, slip should occur dominantly on the basal and prism systems (Fig. 1-14). As rhomb slip, in the absence of prism and basal slip, would only be expected under very specific conditions, such as properly oriented crystals, this scenario is probably unlikely (Fig. 1-14). An alternative and more plausible explanation would be that these patterns represent a combination of slip on the prism and basal planes. In such a situation where slip was accommodated by both the prism and basal systems, the c-axes may plot somewhere between the edge and the middle of the plot, giving the appearance of predominantly rhomb slip. An equal activation of both the prism and basal slip systems would place the samples in a temperature range consistent with that interpreted from the quartz mylonite microstructures.

It is not clear whether the CPO's from samples ML 6-2, ML 7-1 and ML 11-2 can be classified as either type 1 or type 2 fabrics of Lister (1977), however, they do not exhibit small circle girdles. c-axis maxima perpendicular

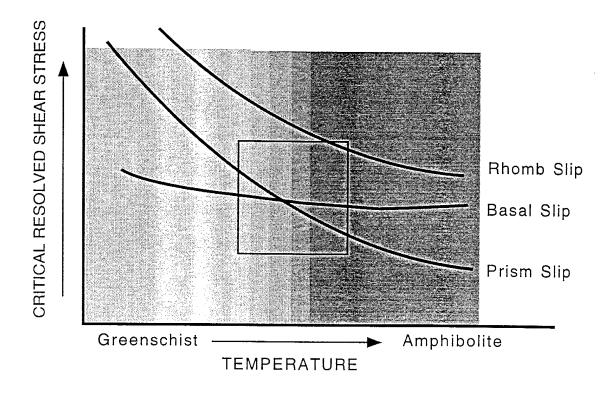


Figure 1-14. Plot of critical resolved shear stress against temperature for basal, prism and rhomb slip systems in quartz, modified from Hobbs (1985). Box represents proposed deformation conditions in the Capilla Peak area based on quartz mylonite c-axis patterns and other information (see text).

to the foliation plane in samples ML 6-2, ML 7-1 and ML 11-2, are consistent with basal slip; and c-axes sub-parallel to the foliation plane are consistent with prism slip (Schmid and Casey, 1986). These three plots lie somewhere between the flattening and plane strain field.

The crystallographic preferred orientations from the third group, sample ML 1-2, are different from those of the other groups. The c-axis distribution is relatively asymmetric with respect to S_2 and L_2 and is consistent with a larger component of non-coaxial flow than in the other samples (c.f. Schmid and Casey, 1986). This is consistent with the microstructure seen in thin section (Fig. 1-13) for example, Berthe et al. (1979) suggest that S-C fabrics are indicative of a large component of non-coaxial flow.

Slip dominantly on the basal systems is believed to be favored at low temperatures and is consistent with slip under greenschist facies conditions (Lister, 1978), while slip dominantly on the prism systems is proposed to occur at higher temperatures (Hobbs, 1972). Based on the c-axis preferred orientations, we interpret that slip occurred mainly along both the basal and prism planes compared to the other possible slip directions (Schmid and Casey, 1984). This is consistent with the inferred metamorphic grade of the deformation (c.f. Fig. 1-14).

A weak asymmetry in CPO's from ML 11-2, ML 1-2, ML 7-3, ML 7-15 and ML 6-2 is defined by the more populated NE-SW girdle suggesting a weak component of simple shear and an east-side-up sense of shear. In fabric ML

11-3, the NW-SE girdle has a higher population than the NE-SW girdle and is interpreted as west-side-up sense of shear.

CONCLUSIONS

The dominant NNE striking foliation (S_2) is observed in all lithologic units and overprints at least one older foliation. Deformation related to S_2 is preferentially preserved and partitioned into high and low strain zones. In low strain zones deformation is expressed as small scale folding and/or a poorly-defined crenulation cleavage. In high strain zones deformation is expressed as a strong planar foliation.

Microstructures such as grain-shape preferred orientations and S-C fabrics in the quartz mylonites and schists as well as asymmetric porphyroclast systems and hornblende growth in the metarhyolites and amphibolites record information about the character of deformation and metamorphism. From these microstructures the metamorphic grade, sense of shear and type of strain are inferred.

Constraints on metamorphic grade are based primarily on quartz mylonite microstructures and crystallographic fabrics and aligned mineral growth within the amphibolite layers (Part II). Quartz mylonite microstructures ranging from monocrystalline quartz ribbons to fine-grained equigranular foam textures are indicative of formation during an upper

greenschist to lower amphibolite grade deformational event (c.f. Simpson, 1985). This interpretation is consistent with that inferred from the quartz c-axis fabrics. From the c-axis patterns it is inferred that strain was accommodated by a combination of basal and prism slip, again consistent with upper greenschist to lower amphibolite grade conditions. In the amphibolite layers, the foliation-forming euhedral hornblendes grew synchronous with S_2 , the dominant NNE striking foliation observed in the Capilla Peak area, and are proposed in Part II to have grown at lower amphibolite grade conditions, probably around 500°C.

Microstructures within the Blue Springs Schist, quartz mylonites and Sevilleta Metarhyolite as well as outcrop scale structures in the amphibolite layers preserve evidence for sense of shear. In the schists, microscopic folds in the quartz layers and S-C surfaces preserve asymmetries consistent with east-side-up sense of shear. In the quartz mylonites, both east-side-up and west-side-up kinematic indicators such asymmetric porpyroclasts and asymmetric quartz crystallographic preferred orientations are present. However, five out of the seven quartz mylonite samples have microstructures and corresponding c-axis fabrics suggestive of an east-side-up sense of shear. In the amphibolite layers large sigmoidal shaped pods, and delta-type porphyroclast systems in the metarhyolites suggest east-side-up shear.

Delta-type porpyroclast sysems are dominant in the Sevilleta

Metarhyolite and are interpreted to suggest that the strain rate is relatively
high compared to the recystallization rate (c.f. bands Hammer and Passchier,
1991) and are interpreted to have formed during progressive general non-

coaxial flow. General non-coaxial flow is probably more common than the dominantly simple shear deformation documented in many other studies (c.f. White et al., 1980).

It is proposed that the Proterozoic rocks currently exposed in the Manzano Mountains were all part of large ductile shear zone and that the Monte Largo Shear zone observed in Monte Largo Canyon is just one localized high strain zone within this larger deformation zone. This deformation occurred during lower amphibolite facies metamorphic conditions at ca. 1.4 Ga.

Part 2

STRUCTURAL AND ⁴⁰Ar/³⁹Ar CONSTRAINTS ON MIDDLE PROTEROZOIC DEFORMATION AND METAMORPHISM IN THE MANZANO MOUNTAINS, CENTRAL NEW MEXICO.

ABSTRACT

Rocks in the Capilla Peak area of the Manzano Mountains, New Mexico have been deformed at ca. 1660 Ma and again at 1440 Ma. 40 Ar/ 39 Ar geochronologic analyses on hornblende, actinolite, muscovite and biotite were conducted to constrain the timing of the observed deformational events. Actinolites are overgrown and crosscut by the well developed foliation-forming hornblendes. The hornblendes define the youngest tectonic fabric in the amphibolite. The actinolites yield complex 40 Ar/ 39 Ar age spectra characterized by age gradients increasing from ~200 Ma to 1600 Ma. These samples have total gas ages ranging from ~1110-1290 Ma. The hornblendes have less complex spectra and overall older ages. In particular, one sample yields a plateau age for 80% of the gas release of 1409 \pm 10 Ma (2 σ). Based on this result, it is proposed that temperatures at ca. 1400 Ma were sufficiently high to reset actinolite with respect to argon (at least 400°C), and that the amphibolites cooled through the hornblende closure temperature at ca. 1400 Ma.

Muscovite separates show variable age spectra with age gradients ranging from ~200 Ma to 1400 Ma. They record terminal ages ranging from ~1320 to 1400 Ma. The muscovite samples were collected over a vertical section of more than 1.5 km and show an age discordance with the highest being ~60 Ma older than the lowest sample. Assuming the samples within the traverse have similar argon $T_{\rm C}$, the apparent ages have been interpreted to suggest a cooling rate of 0.3-1°C/Ma.

Biotites from similar structural levels yield plateau ages of 1276±6 Ma, 1379±4 and 1401±18 Ma respectively. This age discordance reflects a variation in biotite closure temperatures due to differences in grain-size coupled with apparent slow cooling.

Together, the geochronologic and structural data are interpreted to support a model of regional deformation, metamorphism and mineral growth at ca. 1400-1450 Ma. These data are interpreted to suggest that temperatures between 400 and 550°C existed ca. 1400 Ma and the rocks cooled to 200-300°C by at least ca. 1380 Ma, followed by a period of crustal stability and protracted cooling.

INTRODUCTION

Proterozoic rocks in the southwestern U.S. have experienced a complex deformational and thermal history related to cratonic assembly during the Yavapai and Mazatzal orogenies. Rocks in central New Mexico experienced only Mazatzal (1.65 Ga) and younger deformation. U-Pb dates on metamorphic zircons, crosscutting relationships between granitic plutons of known age and the regional foliations indicate that at least one deformational event occurred between 1.74 and 1.60 Ga (Bowring and Karlstrom, 1990; Karlstrom and Bowring, 1991; Bowring, 1991; Shastri and Bowring, 1992; Bauer and Williams, 1993). At that time, P-T conditions of 500-550°C and 3.5-4.5 kbar existed in structural levels currently exposed in northern New Mexico, with lower grade conditions in rocks exposed in central and southern New Mexico (Grambling, 1986; Bowring and Karlstrom, 1990). Abundant K-Ar, 40 Ar/39 Ar and Rb-Sr whole rock and mineral ages throughout the southwestern U.S. indicate that a period of metamorphism occurred ca. 1.4 Ga (Grambling, 1989; Bauer and Pollock, 1993). In the Manzano Mountains of New Mexico (Fig. 2-1), contact metamorphism and localized deformation were associated with emplacement of 1.4 Ga plutons (Thompson et al., 1996). The 1.4 Ga plutons were originally believed to be anorogenic since they were believed to be undeformed and additional evidence for orogenisis ca 1.4 Ga had not been documented.

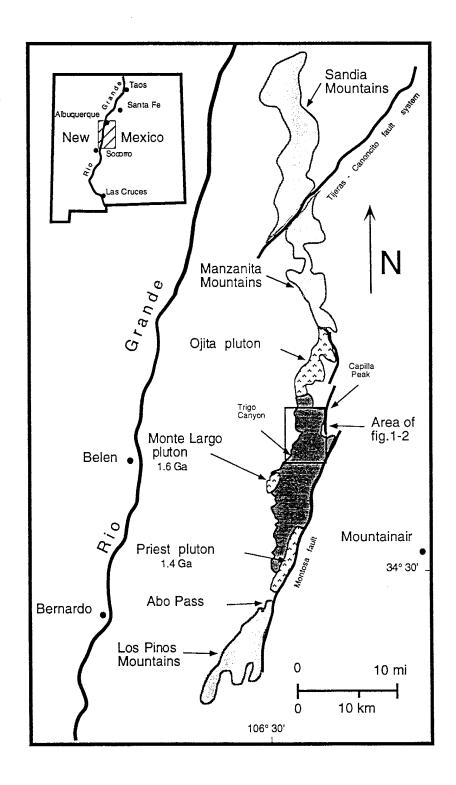


Figure 2-1. Geologic map of Proterozoic outcrops in central New Mexico showing the location of the study area and the plutonic rocks within the Manzano Mountains.

Two possible metamorphic and deformational models are proposed to explain the 1.4 Ga tectonism. The first model requires a metamorphic event at 1.4 Ga without regional deformation. This model was originally proposed to explain the absence of a pervasive tectonic foliation in 1.4 Ga plutons and the presence of 1.4 Ga mineral ages (Karlstrom and Bowring, 1993; Bauer and Williams, 1994; Karlstrom and Williams, 1995; Karlstrom et al. 1996 (in review)). The second model involves regional deformation and metamorphism at 1.4 Ga.

The purpose of this paper is to document a 1.4 Ga deformation event recorded by the Proterozoic rocks of the Manzano Mountains, and to further constrain the post-1.4 Ga thermal history of the area. We focus on petrography and chemistry of amphibolites and 40 Ar/ 39 Ar geochronology of amphiboles, biotites and muscovites located more than 10 km from the exposed margins of 1.4 Ga plutons (Fig. 2-1).

GEOLOGIC SETTING AND PREVIOUS WORK

The Manzano Mountains are an eastward tilted fault block of Proterozoic metasedimentary, metavolcanic, and plutonic rocks unconformably overlain by Paleozoic sedimentary rocks. The mountain range is approximately 60 km long and 15 km wide and defines the eastern margin of the Rio Grande rift between the Manzanita Mountains and Abo Pass (Fig. 2-1). The study area includes a steeply

dipping structural sequence of muscovite schists, quartz mylonites, metarhyolites and amphibolite layers (Fig. 2-2).

Three phases of Proterozoic deformation have been recognized in the Manzano Mountains. Two phases of folding can be bracketed between 1.6 Ga and 1.4 Ga by crosscutting relationships with the 1656±10 Ma Monte Largo Pluton and the 1427±10 Ma Priest Pluton (Bauer et al., 1993) and evidence for one post-1.4 Ga phase of deformation can be found within the Priest Pluton (Heizler et al., 1996). Regional P-T conditions at ca. 1.4 Ga were estimated at 500-540°C and 4kb based on mineral assemblages, geothermometry and geobarometry (Grambling, 1989; Williams, 1990; Thompson et al., 1991).

Geochronologic data in the Manzano Mountains and the adjacent Los Pinos Mountains include: five U-Pb dates on zircons from metarhyolites in the Los Pinos Mountains and from the Priest and Monte Largo plutons (Shastri and Bowring, 1992; Shastri, 1993; Bauer et al., 1992; Bowring et al., 1983), three Rb-Sr dates from the Priest and Ojita plutons (Bolton, 1976; White, 1979; Brookins et al., 1980), and fourteen ⁴⁰Ar/³⁹Ar age spectra analyses of muscovite and biotite from in and around the Priest and Ojita plutons (Dallamyer et al., 1990; Thompson et al., 1991, 1996; Heizler et al., 1996). The U-Pb and Rb/Sr dates help constrain the ages of the plutons and the metarhyolite while the ⁴⁰Ar/³⁹Ar ages from in and around the plutons help constrain the timing of mineral growth associated with pluton emplacement. U-Pb zircon dates from the metarhyolites and Monte Largo Pluton range from 1656 to 1680 Ma while a U-Pb zircon from the Priest Pluton yields an age of 1427±10 Ma. Rb/Sr dates from

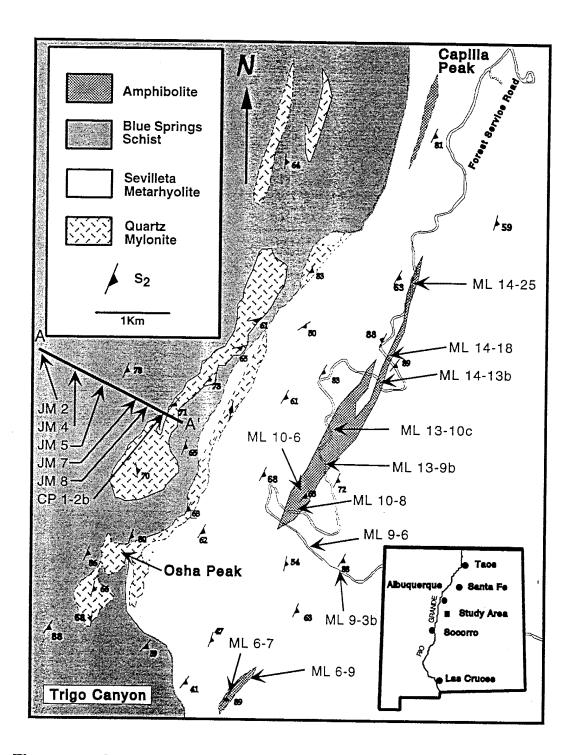


Figure 2-2. Generalized geologic map of the study area. Note the dominant NNE striking foliation and the various lithologies. A-A' marks the location of a sampling traverse across the Blue Springs schist. Amphibolite sample locations are noted.

the Manzano Mountains range from 1439 to 1569 Ma and all of the ⁴⁰Ar/³⁹Ar ages are between 1300 and 1450 Ma. The geochronologic data presented in this paper consist of ⁴⁰Ar/³⁹Ar data from the Capilla Peak area. The study area is more than 10 km from the plutonic rocks in the Manzano Mountains and ages will be used to help constrain the timing of deformation rather than pluton emplacement.

ANALYTICAL METHODS

Amphiboles and biotite from amphibolites within the Sevilleta

Metarhyolite of Stark and Dapples (1946), and muscovite concentrates from the
fine-grained Blue Springs Schist of Stark and Dapples (1946) were separated for

40 Ar/30 Ar age spectrum analysis (Fig. 2-2). All amphibolite samples contain
both actinolite and a hornblende; these phases were separated from the 200-300
mesh size fraction along with a 60 mesh coarse amphibole (bulk amphibole)
separate. Amphibole separates were purified using heavy liquids, the Franz
magnetic separator and careful hand-picking based on color. All samples were
ultrasonically cleaned for 5-10 minutes, rinsed in acetone, and dried at ~100°C.
The mineral separates were packaged in Cu foil and placed in alternating
positions around circular 6-hole irradiation trays with Fish Canyon Tuff
sanidine [27.84 Ma] and Fe-mica biotite [307.3 Ma] flux monitors. Samples
were irradiated in the L-67 position of the Ford reactor at the University of
Michigan in three irradiations (NM-19, NM-29 and NM-42).

Analyses on all samples were performed at the New Mexico Geochronology Research Laboratory. The laboratory utilizes a MAP-215-50 mass spectrometer equipped with a Faraday and electron multiplier collector. The sensitivity is $\sim 3\times 10^{-17}$ mol/pA, and has a background at mass 36 of $\sim 1\times 10^{-18}$ moles. The furnace blank is $< 5\times 10^{-15}$ moles 40 Ar at temperatures below 1300°C. J-factors were determined by CO₂ laser total-fusion analysis of flux monitors to a precision of 0.25%. Unknown samples were incrementally heated, with seven minute heating steps, in a double vacuum Mo resistance furnace with an accuracy of 5-10°C, and a precision of $\sim \pm 1$ C. The extracted gas was purified with SAES AP-10 and GP-50 getters for four minutes.

Due to the complexity of the age spectrum data, plateaus as defined by three or more contiguous steps, totaling more than 50% of the ³⁹Ar released that agree within analytical error at the 95% confidence level (Fleck et al., 1977), exist for only a few samples. Age determinations, when possible, are either integrated ages calculated for the relatively flat portion of the age spectra using ³⁹Ar as the weighting factor, or are the terminal ages defined by the high temperature steps of the age spectra.

All samples analyzed were examined petrographically. Back-scattered electron imaging, X-ray mapping and electron microprobe analyses of amphibolite samples ML 14-18 and ML 10-8 were performed on the University of New Mexico JEOL 733 electron microprobe and Arizona State University JEOL 8600 electron microprobe. Analyses were performed on doubly polished,

carbon-coated thin-sections and standard ZAF calibration techniques were used. The amphiboles were analyzed for Na, Mg, Al, Si, Ca, Cl, K, Ti, Mn and Fe (Table 2-1). The locations of 35 spot chemical analyses were individually selected from the back-scattered electron images and obtained with a 10 μ m beam. Numbers of ions are based on 23 oxygens and Fe is reported as Fe²⁺.

Locations for microprobe spot analyses were carefully selected in order to avoid inclusions and intragranular impurities. Samples are probably not representative of the total actinolite and hornblende composition. In contrast, 40 Ar/ 39 Ar analyses were performed on multiple complete crystals, hence sampling the entire crystal. As a result, some variation in microprobe compositional data and the 40 Ar/ 39 Ar compositional data is expected.

SAMPLE DESCRIPTION

Amphibolites

Amphibolite samples were collected from a series of amphibolite layers believed to be early mafic dikes (Stark and Dapples, 1946) that cut the Sevilleta Metarhyolite. The amphibolite layers are 1-20 m in width, up to 400 m in length and are located more than 10 km from the Priest and Ojita plutons (Figs. 2-1 and 2-2). The amphibolite layers are highly foliated and aligned parallel to the regional foliation trend (Fig. 2-2).

	Llowelle !		7
	Hornblende	Actinolite	
SiO ₂	$43.3 \pm .13$	53.0 ± 1.3	1
TiO_2	$0.34 \pm .05$	$0.07 \pm .04$	
Al ₂ O ₃	15.0 ± 1.5	3.8 ± 1.3	
FeO	$17.8 \pm .68$	3.8 ± 1.3	
MnO	$0.25 \pm .05$	$0.23 \pm .04$	
MgO	$\sim 8.8 \pm .90$	15.5 ± .98	
CaO	11.8 ±.16	12.2 ± .20-	127
Na ₂ O ,₀५²² K ₂ O √,∘५²	1.3 ± .11	$0.39 \pm .11$	1.005
∠	$0.4 \pm .08$	0.11± .06	- 1 .0"
<u>C1</u>	0.01 ±.01	0.01 ± .01	
Total	99.0	98.8	
-numbers of ions based on 23 oxygens			
- iron is reported as Fe ²⁺			

Table 2-1. Average oxide weight percent derived from electron microprobe point analyses of amphibolites ML 14-18 and ML 10-8, showing two distinct amphibole phases. Hornblende data are averaged from 22 point analyses; actinolite data are from 13 point analyses.

The amphibolites are petrologically complex and contain two amphibole phases (Table 2-1), plagioclase, biotite, quartz, epidote, and minor amounts of magnetite and pyrite. Johnson (1986) reported that amphiboles from the Manzano Mountains were zoned, with blue-green hornblende cores surrounded by retrograde actinolite rims. This is contradicted by backscattered electron images, which show a sharp boundary between the amphibole phases, no compositional variation from rim to core, and distinct amphibole compositions. Actinolites occur as inclusion-rich 0.25-0.75 mm anhedral, equidimensional, cores overgrown by hornblende rims and crosscut by elongate hornblende crystals. Samples ML 10-8 and ML 14-18 have been characterized by transmitted light microscopy, electron mapping and electron microprobe analysis. Electron microprobe results from samples ML 10-8 and ML 14-18 are displayed in Table 2-1. Data from the two samples are not differentiated since compositions are identical within one standard deviation. The actinolite cores are typically relatively silica-rich and magnesium-rich with respect to the higher aluminum, lower iron ferro-tschermakitic hornblende rims (c.f. Yavuz, 1996) (Fig. 2-3, Table 2-1). Quartz inclusions and traces of cleavage planes within the actinolites are aligned nearly perpendicular to the dominant hornblende foliation. These inclusion trails are interpreted as recording an older, relict foliation.

Individual ferro-tschermakitic hornblende crystals are euhedral and bladed. These amphiboles define the well-developed foliation evident in the amphibolites (Fig. 2-4). The composition of the bladed hornblende is

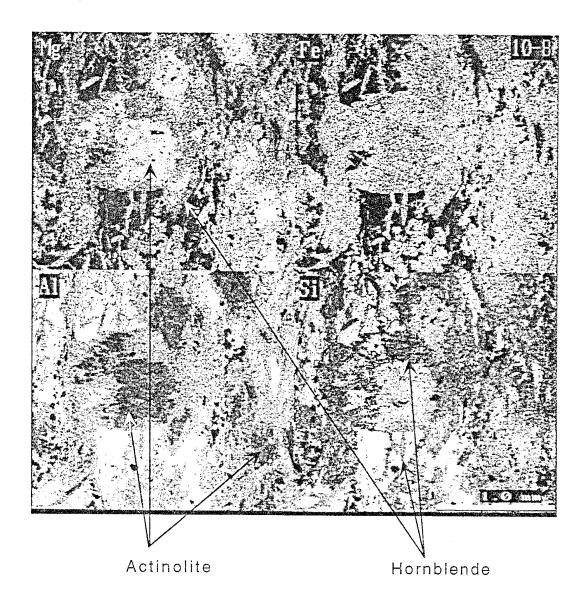


Figure 2-3. X-ray maps of ML 10-8 amphibolite showing distribution of Mg, Fe, Al, and Si. The brighter the area, the greater the amount of the labeled element. Note the high number of inclusions in the actinolite and the sharp boundary between the two distinct mineral phases.

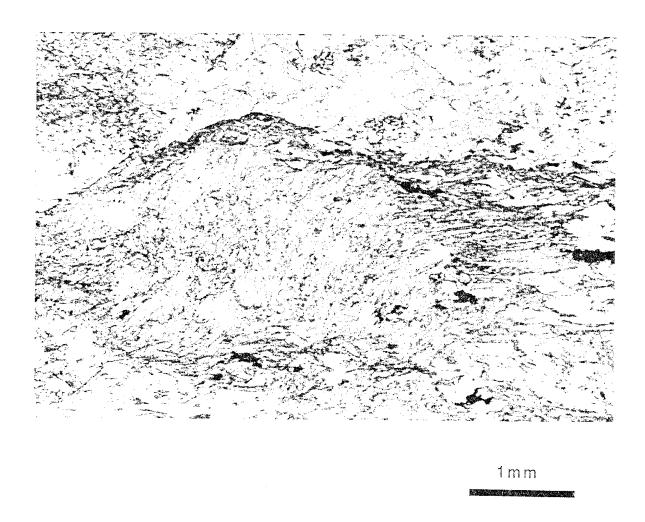


Figure 2-4. Photomicrograph of amphibolite sample ML 10-8. Long dimension of photo, ~6mm, is parallel to the dominant foliation. Note the abundant inclusion trails defining a relict foliation in the actinolite nearly perpendecular to the foliation-forming hornblende crystals.

indistinguishable from that of the hornblende rims around the actinolites (Table 2-1). Hornblende rims and individual euhedral hornblende crystals exhibit mutually cross-cutting relationships, indicating that they grew contemporaneously (Fig. 2-4).

Biotites occur locally within the amphibolite layers. The biotites are >0.25-0.75 mm in size and are aligned parallel to the dominant hornblende foliation.

Muscovite schists

The Blue Springs schist is intruded in the south by the Monte Largo pluton and trends NNE for approximately 17 km to where it is disconformably overlain by Pennsylvanian limestones in the north. The schist exposure is approximately 4 km wide, bounded on the west by overlying pediment gravels and on the east by the White Ridge quartzite (Fig. 2-2). Three dominant lithologies exist within the Blue Springs schist: metasiltstones, phyllites, and muscovite-chlorite schists (Stark 1956; Bauer 1982) (Fig. 2-2). The schist is highly foliated and crenulated and appears to be highly inhomogeneous with respect to composition and degree of deformation. Several generations of quartz veins are observed with varying thicknesses and frequency throughout the units. Also within the schist are 1-10 m-wide metasedimentary quartz mylonite layers (the "quartz reefs" of Stark and Dapples, 1946).

Six samples were collected along a traverse A-A' (Fig. 2-2) over a vertical section of greater than 1.5 km perpendicular to the regional NNE-striking foliation. The NNE-striking foliation is a crenulation cleavage in the schist (Fig. 2-5). It cuts an older foliation and becomes increasingly well developed with both increasing structural position in the traverse and increasing proximity to quartz-rich rocks. The spacing between cleavage planes decreases with increasing crenulation cleavage development. The schist contains abundant quartz and muscovite and minor amounts of biotite and chlorite. The muscovite grain size varies from 0.5 to 10 microns. The mean grain size decreases with increasing structural position and proximity to quartz-rich rocks, as the degree of crenulation cleavage development increases. The muscovites contain mineral inclusions and exhibit evidence of intra- and intergranular cataclasis (c.f. Goodwin and Wenk, 1990), particularly along crenulation cleavage planes (Fig. 2-5). In general there is a greater amount of intra- and intergranular cataclasis in the structurally higher samples than in the lower samples. In summary, the mechanisms by which the crenulation cleavage developed resulted in a decrease in muscovite grain size.

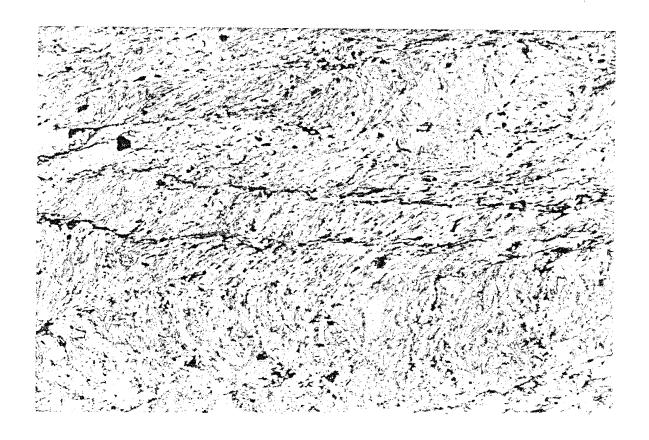


Figure 2-5. Photomicrograph of fine-grained muscovite schist from the Blue Springs schist. Long dimension of photo is 6mm. Note the older foliation is subparallel to the short dimension of the photo and is cut by a crenulation cleavage, parallel to the long dimension of the photo.

⁴⁰Ar/³⁹Ar RESULTS

General Description

Eleven bulk amphibole separates from the amphibolite layers were stepheated and show complex age spectra. Eight finer grained actinolite and seven hornblende separates from the same hand samples were separated and heated from 500°C to 1700°C in 13 steps and have age spectra showing gradients from 200 Ma to 1500 Ma (Fig. 2-6). The hornblende age spectra tend to contain fewer low age steps than the more complex actinolite and the bulk amphibole age spectra. Age determinations, when possible, are either integrated ages calculated for the relatively flat portion of the age spectra weighting ³⁹Ar, or are the terminal ages defined by the high temperature steps of the age spectra. Interpretation of the ⁴⁰Ar/³⁹Ar age spectra are quite subjective and thus actual age assignment is also subjective. ⁴⁰Ar/³⁹Ar ages assigned in this paper reflect one possible interpretation of the data.

The K/Ca ratios are plotted above the corresponding age spectra (Figs. 2-6-2-12). Often the K/Ca plot is used to evaluate the age spectrum. In the hornblende and the actinolite samples a high K/Ca ratio is associated with the abundant young age steps in the early portions of the age spectra, and a relatively constant K/Ca ratio corresponds to the relatively flat portions of the age spectra. The K/Cl ratio is also plotted above the corresponding amphibole

Figure 2-6. 40 Ar/ 39 Ar spectra from bulk amphiboles (A, B), actinolites (C, D) and hornblendes (E, F) separated from samples ML 14-18. The actinolite age spectra are highly complex, showing age gradients and abundant young ages in the early portions of the spectra. The hornblende spectra are also complex; however, a significant portion of the age spectra yields an approximate age of 1400 Ma. Sample 14-18-1 (E) does not yield a true plateau age, but the flat portion of the age spectrum yields an integrated age of 1361 \pm 8 Ma; while sample ML 14-18-2 (F) yields a true plateau age of 1409 \pm 10 Ma.

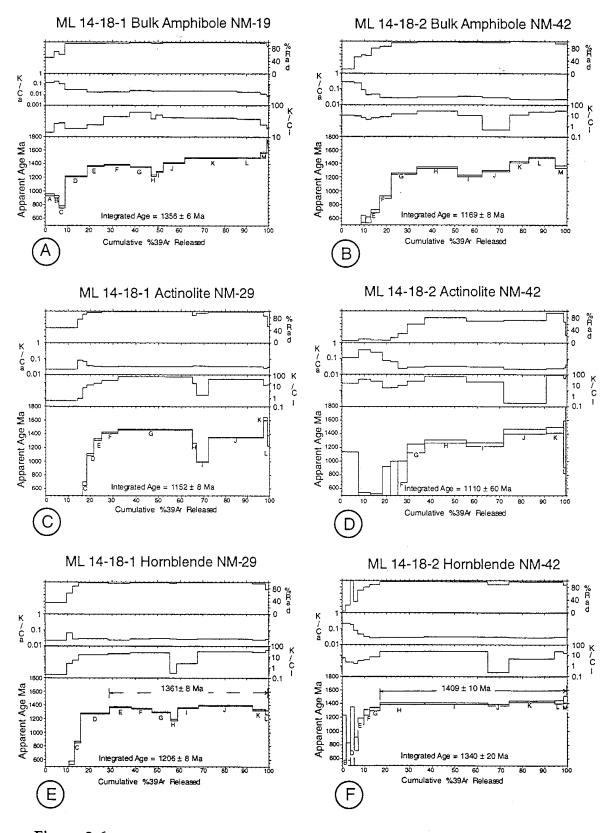


Figure 2-6.

age spectra. In contrast to K/Ca, the forms of the K/Cl plot do not appear to be directly related to the age spectra. They are included as additional information.

Detailed description of age spectra

Samples presented in this paper have been irradiated in three separate irradiations (NM-19, NM-29 and NM-42) (Table 2-2). Irradiation NM-19 included eight bulk amphiboles, separated at a 60 mesh size fraction and two biotites. Irradiation NM-29 included two biotites, five hornblendes and six actinolites separated at a 200-300 mesh size fraction from the same sample as the bulk amphiboles from irradiation NM-19. The purpose for running NM-29 samples was to help understand the complex bulk amphibole spectrum observed from the NM-19 samples. Samples irradiated in NM-42 consisted of two bulk amphibole, two hornblende, two actinolite, one biotite and 10 muscovites. Detailed descriptions of amphibole age spectra are presented below.

Amphiboles

ML 14-18: Bulk amphibole, actinolite and hornblende were separated and analyzed from sample ML 14-18. Two sets of each were separated and irradiated in three different irradiations (NM-19, NM-29 and NM-42).

The general shape and values of the K/Ca and K/Cl relative to the age spectra for ML-14-18 bulk amphibole, actinolite and hornblende are similar for

Sample	Moles ³⁹ Ar	Weight mg	J-factor	Moles ³⁹ Ar/mg	run
ML 9-3b Bulk	2.90E-15	2.6	0.01442337	7.73E-16	NM-19
ML 9-6 Bulk	3.40E-15	2.2	0.01445203	1.07E-15	NM-19
ML 13-9b Bulk	1.40E-14	2.9	0.01449799	3.33E-15	NM-19
ML 13-10c Bulk	1.80E-14	2.4	0.01450297	5.17E-15	NM-19
ML 14-18-1 Bulk	1.40E-14	3.4	0.01447063	2.85E-15	NM-19
ML 14-13b Bulk	4.00E-14	2.3	0.01444197	1.20E-14	NM-19
ML 6-7 Bulk	2.70E-14	2.8	0.01447585	6.66E-15	NM-19
ML 10-6 Bulk	2.80E-14	2.8	0.01451015	6.89E-15	NM-19
ML 10-8-1 Bulk	6.30E-14	3.8	0.01442939	1.15E-14	NM-19
			-		
ML 10-6 Act	4.10E-14	3.66	0.009526213	1.18E-14	NM-29
ML 13-9b Hbl	2.80E-14	2.08	0.009528751	1.41E-14	NM-29
ML 14-25 Act	1.60E-13	10.97	0.009518372	1.53E-14	NM-29
ML 9-3b Act	1.10E-14	1.58	0.009540874	7.30E-15	NM-29
ML 9-3b Hbl	1.00E-14	3.65	0.009543184	2.87E-15	NM-29
ML 9-6 Act	4.30E-15	1.89	0.009547693	2.38E-15	NM-29
ML 9-6 Hbl	1.00E-14	3.06	0.00954989	3.42E-15	NM-29
ML 6-7 Act	7.20E-15	1.11	0.009554171	6.79E-15	NM-29
ML 6-7 Hbl	1.50E-14	1.93	0.009556255	8.13E-15	NM-29
ML 14-18-1 Act	8.60E-15	1.99	0.009558301	4.52E-15	NM-29
ML 14-18-1 Hbl	8.80E-15	1.75	0.009560309	5.26E-15	NM-29
ML 14-18-2 Act	6.70E-16	0.51	0.01070469	1.23E-15	NM-42
ML 14-18-2 Hbl	2.10E-15	1.14	0.01073805	1.72E-15	NM-42
ML 14-18-2 Bulk	2.90E-15	1.04	0.01073225	2.60E-15	NM-42
ML 10-8-2 Act	7.80E-16	0.16	0.01071991	4.55E-15	NM-42
ML 10-8-2 Hbl	2.90E-15	0.6	0.01073843	4.50E-15	NM-42
ML 10-8-2 Bulk	1.40E-15	0.42	0.01077295	3.09E-15	NM-42

Table 2-2. Samples were irradiated in three separate irradiations (NM-19, NM-29 and NM-42). The moles ³⁹Ar/mg for each amphibole sample can be used to evaluate both the amount of gas derived from each sample, and the effectiveness of sample separation.

the majority of the bulk amphibole, actinolite and hornblende samples; exceptions will be noted. The K/Ca of the bulk amphiboles mirror the age spectra with high values corresponding to the low ages; the average K/Ca ratio around 0.02. Most resulting ⁴⁰Ar/³⁹Ar age spectra from the duplicate runs are somewhat reproducible especially for hornblende samples ML 14-18-1 and ML 14-18-2 (Fig. 2-6E,F).

The bulk amphibole age spectra are characterized by age gradients ranging from as low as 200 Ma in the first 25% of the ³⁹Ar released to 1300-1400 Ma for the intermediate temperature steps. The high temperature steps yield ages ranging from 1500-1600 Ma, with terminal ages as old as 1700 Ma (Figs. 2-6A, B). The K/Cl for actinolites and hornblendes ML 14-18-1 and ML 14-18-2 shows a direct correlation with the age spectra, including the distinctive dip in the age spectra around 1100°C, and steps H-I in ML 14-18-1 and I-J in ML 14-18-2 (Fig. 2-6); this is observed in nearly all amphibole samples analyzed.

The actinolite separates from ML 14-18 exhibit complex and non-reproducible age gradients (Fig. 2-6C, D). Terminal ages are between 1500 and 1600 Ma. The K/Ca ratios for the actinolites in ML 14-18-1 and 2 vary between 0.023 and 0.30, significantly higher than the 0.01 determined from electron microprobe analysis (Table 2-1).

The hornblende age spectra show age gradients similar to those observed for the bulk and actinolite samples for the first 20% of the ³⁹Ar release. The remainder of the hornblende age spectra are relatively flat and yield integrated ages calculated for the flat portions of the age spectra of 1361±8 and

1409±10 Ma respectively (Fig 2-6E, F). The average K/Ca for the hornblende samples is 0.04, similar to that determined from electron microprobe analyses.

ML 10-8: Two bulk amphibole splits, one actinolite and one hornblende were analyzed from this sample. 40Ar/39Ar age spectra from bulk amphiboles are reproducible in age and overall form (Fig. 2-7A, B). Bulk amphibole age spectra are characterized by age gradients from 800 to 1600 Ma; however, 50-70% of the cumulative 39 Ar released yields ages between 1300 and 1350 Ma. The actinolite age spectrum consists of an age gradient over 90% of the age spectrum with very old (>1800 Ma) initial and final steps (Fig. 2-7C). Steps H and I yield the oldest apparent ages of around 1240 Ma. The K/Ca ratios vary considerably for this sample and yield an undulatory pattern. The hornblende shows a complex age spectrum stepping from 300 Ma to 1340 Ma in the first 30% of the cumulative 39 Ar released. The final 60% of the age spectrum is relatively flat and yields an integrated age of 1329±10 Ma for steps G-K (Fig. 2-7D). Duplicate biotite analyses were also run from sample ML 10-8 biotite. The general shapes and ages of the flat portion of the two age spectra are similar. Biotite ML 10-8-1 has a large (\sim 15% of the total gas) first step with an age of 220 Ma (Table 2-2). The sample yields an age gradient from 1100 to 1250 Ma (Fig. 2-7E). Biotite ML 10-8-2 has three young steps totaling less than 5% of the total ³⁹Ar released and a flat age spectrum from step E to K yielding an integrated age of 1276±6 Ma (Fig. 2-7 F). The forms of the two biotite age spectra are similar, but the total gas ages are very different.

Figure 2-7. ⁴⁰Ar/³⁹Ar spectra from bulk amphiboles (A, B), actinolites (C), hornblende (D) and biotites (E, F) separated from sample ML 10-8. The bulk amphiboles show age gradients from 800 to 1800 Ma. A significant portion of the age spectra from both ML 10-8-1 and ML 10-8-2 yield ages of 1.4 Ga. The actinolite age spectrum (C) is highly complex, showing an age gradient and a very old first step. The hornblende spectrum is also complex, however, a significant portion of the age spectra yields an approximate age of 1350 Ma. Biotite 10-8-1 (E) has a very large young (250 Ma) first step. Neither of the biotites yield plateau ages; however, biotite ML 10-8-2 (F) has fewer young ages and yields an integrated age from the flat portion of the age spectrum of 1276 Ma; older than that from ML 10-8-1 biotite.

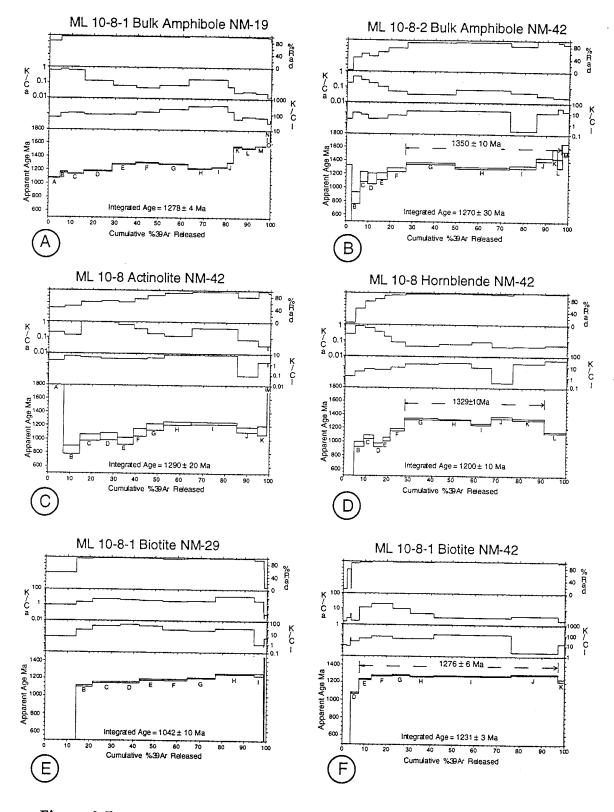


Figure 2-7.

ML 6-7: A bulk amphibole, actinolite and hornblende were separated and analyzed from sample ML 6-7. 40 Ar/ 39 Ar age spectra from ML 6-7 bulk amphibole, actinolite and hornblende (Fig. 2-8A, C, E) show age gradients from 250 to 1400 Ma. The general shape of the three age spectra is the same with a large age gradient for the first 40% of the cumulative 39 Ar released and a relatively flat portion for the remainder of the spectra. The young ages in the early portions of the age spectra are associated with relatively high K/Ca. Integrated ages calculated over the flat portions of the bulk amphibole, actinolite and hornblende age spectra are 1411 ± 4 , 1405 ± 6 and 1397 ± 6 respectively.

ML 9-6: A bulk amphibole, actinolite and hornblende were separated and analyzed from sample ML 9-6. The ⁴⁰Ar/³⁹Ar age spectrum from bulk amphibole (Fig. 2-8B) shows an age gradient from 600 to 1450 Ma. The actinolite has a complex age spectrum with 45% of the spectrum yielding an age of less than 500 Ma and a terminal age of 1450 Ma (Fig. 2-8). ML 9-6 hornblende also shows an age gradient but has a relatively flat portion between steps F and L, and an integrated age calculated from steps J and K of 1452±4 Ma (Fig. 2-8).

ML 9-3b: A bulk amphibole, actinolite and hornblende were separated and analyzed from sample ML 9-3b. In the case of sample bulk amphibole ML 9-3b, the K/Ca plot bears little resemblance to the age spectrum plot, and can not be used to explain the age variation. The bulk amphibole exhibits a highly complex age spectrum with the majority of steps having ages between 1250 and

Figure 2-8. ⁴⁰Ar/³⁹Ar spectra from ML 6-7 bulk amphiboles (A), actinolite (C) and hornblende (E). The bulk amphibole shows an age gradient from 700 to 1600 Ma. A significant portion of the age spectrum yields an age of 1.4 Ga. The actinolite age spectrum (C) and the hornblende spectrum (E) show more pronounced age gradients than the bulk amphibole. Samples ML 6-7 do not yield plateau ages; however, all three yield integrated ages, calculated from the flat portions of the age spectra, of 1.4 Ga. ⁴⁰Ar/³⁹Ar spectra from ML 9-6 bulk amphibole (B), actinolite (D) and hornblende (F) show highly complex age gradients. The hornblende yields a terminal age around 1450 Ma.

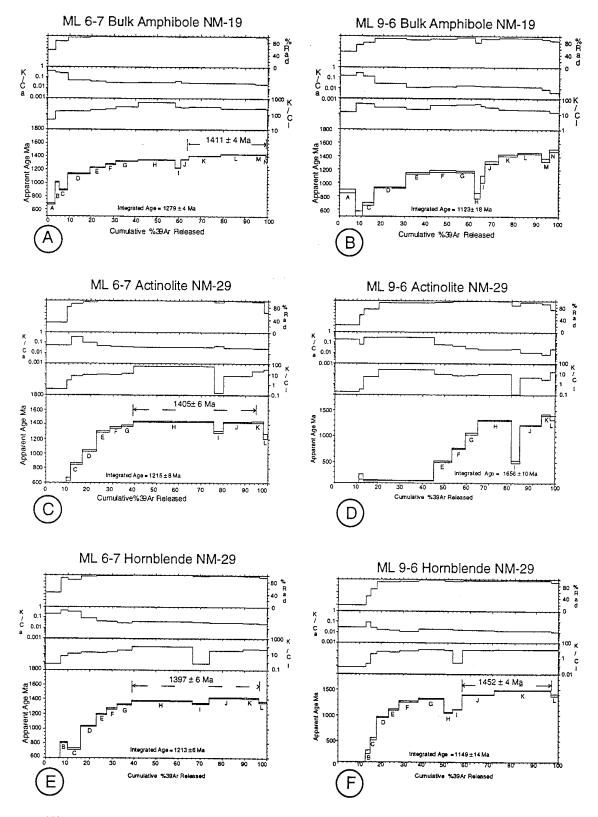


Figure 2-8.

1600 Ma (Fig. 2-9a). Both the K/Ca and K/Cl ratios are relatively flat for about 80% of the age spectrum and do not show the complexity of the age spectrum. ML 9-3b actinolite and hornblende age spectra are very different from the bulk amphibole. Both display an age gradient and maximum ages around 1400 Ma (Fig. 2-9C, E). Unlike the bulk amphibole, the shapes of the actinolite and hornblende age spectra appear to correspond to the shapes of the K/Ca and K/Cl plots.

ML 10-6: A bulk amphibole and actinolite were separated and analyzed from sample ML 10-6. ⁴⁰Ar/³⁹Ar age spectra from the bulk amphibole (Fig. 2-9b) show an age gradient from 1100 Ma to 1700 Ma. The actinolite separate yields a similar age spectrum form; however, the total gas ages differ by ~200 Ma, the terminal and initial ages are lower and the dip in the age spectrum at 1100°C is more pronounced(Fig. 2-9d).

ML 14-13b: A bulk amphibole was separated and analyzed from sample ML 14-13b. The ⁴⁰Ar/³⁹Ar age spectrum from the bulk amphibole (Fig. 2-9F) differs from the other bulk amphiboles in that 80% of the age spectrum is relatively flat at 1350 Ma.

ML 13-9b: A bulk amphibole and hornblende were separated and analyzed from sample ML 13-9b. ⁴⁰Ar/³⁹Ar age spectra from the bulk amphibole and hornblende have nearly identical age spectra and integrated ages (Fig. 2-10A, B). The age spectra are characterized by age gradients from 500-600 Ma to 1500 Ma.

Figure 2-9. ⁴⁰Ar/³⁹Ar spectra from ML 9-3b bulk amphibole (A), actinolite (C) and hornblende (E). The bulk amphibole shows a complex age spectrum with an old integrated age of 1.6 Ga, about 200 Ma older than the other amphibole integrated ages. The actinolite age spectrum (C) and the hornblende spectrum (E) are equally complex, and show more pronounced age gradients than the bulk amphibole. ⁴⁰Ar/³⁹Ar spectra from ML 10-6 bulk amphibole (B) and actinolite (D) show age gradients from 700 to 1600 Ma. ⁴⁰Ar/³⁹Ar spectra from ML 14-13b bulk amphibole (F) shows an age gradient from 1100 to 1700 Ma. The majority of the age spectrum yields an age of 1.35 Ga.

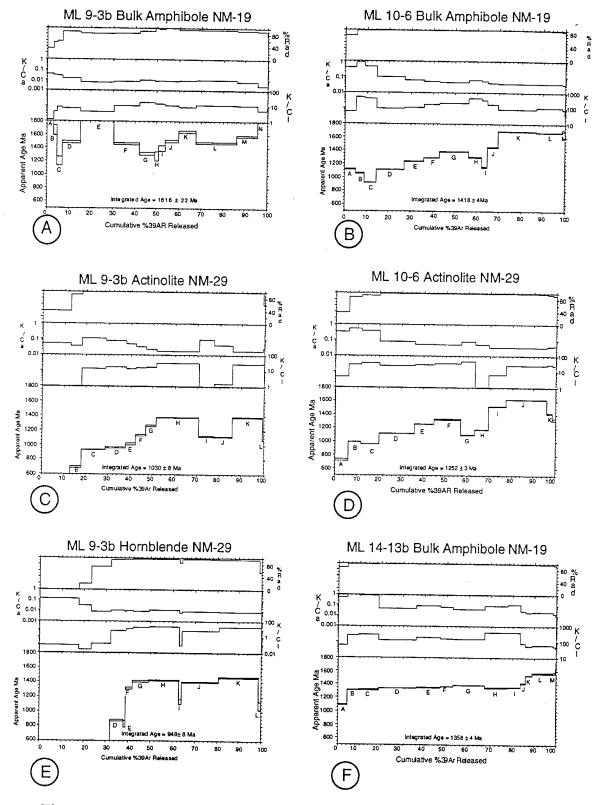


Figure 2-9.

Figure 2-10. ⁴⁰Ar/³⁹Ar spectra from ML 13-9b bulk amphibole (A) and hornblende (B). The bulk amphibole and the hornblende age spectra show age gradients from 500 to 1500 Ma, and an integrated age of 1.0 Ga. The ⁴⁰Ar/³⁹Ar spectra from ML 14-25 actinolite (C) shows an age gradient from 1100 to 1400 Ma, with the majority of the spectrum yielding an integrated age of 1350 Ma. ML 13-10c bulk amphibole (D) shows a gentle age gradient with the majority of the age spectrum yielding an integrated age of 1350 Ma, and a terminal age of 1500 Ma. ML 6-10 biotite (E) and ML 10-2 biotite (F) show relatively flat age spectra, and yield ages of 1.4 Ga.

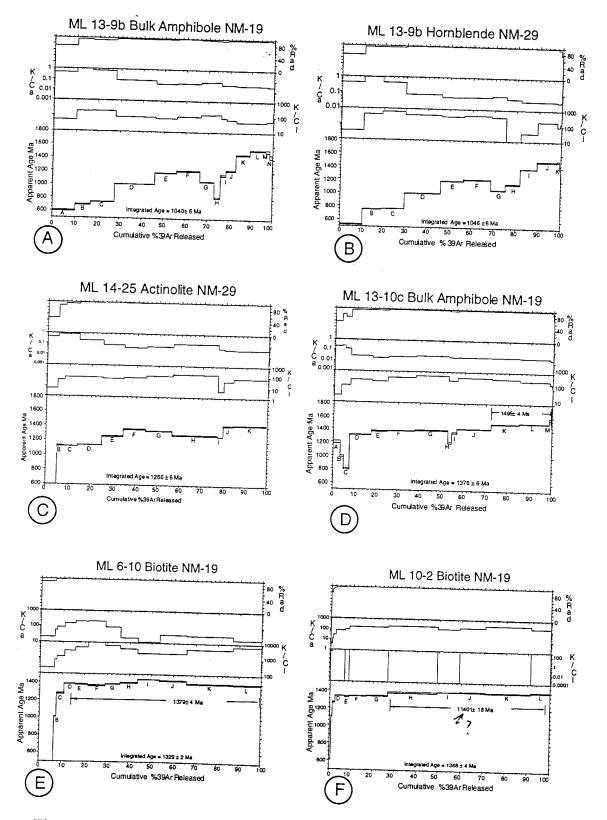


Figure 2-10.

ML 14-25: An actinolite was separated and analyzed from sample 14-25. The ⁴⁰Ar/³⁹Ar age spectrum from the actinolite yields an age gradient from 1100 to 1400 Ma (Fig. 2-10C). About 60% of the age spectrum yields ages between 1350 and 1400 Ma with relatively flat corresponding K/Ca ratios. The sample yields a total gas age of 1266±6 Ma.

ML 13-10c: A bulk amphibole was separated and analyzed from sample 13-10c. The ⁴⁰Ar/³⁹Ar age spectrum from the bulk amphibole shows a gradual age gradient over 90% of the age spectrum from 1306 to 1568 Ma with a relatively flat corresponding K/Ca plot (Fig. 2-10D). An integrated age of 1495 Ma is observed for steps K-M (Fig. 2-10D).

ML 6-10 and ML 10-2 Biotite: A biotite was separated and analyzed from both samples ML 6-10 and ML 10-2. ⁴⁰Ar/³⁹Ar age spectra from the biotites yield relatively flat spectra for at least 90% of the ³⁹Ar released and total integrated ages of 1329±4 and 1368±2 Ma respectively (Fig. 2-10E). The K/Ca for ML 10-2 is nearly an order of magnitude larger and the K/Cl is much lower than that of ML 6-10 (Fig. 2-10E, F, Appendix A).

Muscovites

The muscovite separates collected from within the Blue Springs schist show variable age spectra with age gradients ranging from ~200 Ma to 1400 Ma. The total gas ages from the six samples are all different; however, ages derived from the portion of the age spectra between 20 and 90% of the argon released, yield similar ages. These micas record terminal ages ranging from

~1320 to 1400 Ma (Fig. 2-11, 2-12). Muscovite samples from the structurally highest samples generally yield an older age than the structurally lowest samples. The total age discordance between the lowest and highest sample is approximately 60 Ma (Fig. 2-12D).

JM 2, JM 4:a, JM 5, JM 7:a, JM 8:a and CP 1-2b:a: Six muscovites and duplicate splits JM 4:b, JM 7:b, JM 8:b and CP 1-2b:b yield relatively similar ⁴⁰Ar/³⁹Ar age spectra and K/Ca ratios. All muscovite samples exhibit age gradients and have been plotted relative to JM 2 muscovite (Figs. 2-11, 2-12). JM 2 is the structurally lowest sample and CP 1-2b is the structurally highest. It is difficult to assign an individual age to the muscovite samples since there are no plateaus and since the terminal ages are not always the oldest age on the age spectra. However, a general increase in age is observed from the lowest to the highest sample. Sample JM 2 and the adjacent JM 4 samples yield essentially identical age spectra (Fig. 2-11C, D). JM 2 plots increasingly further below JM 2 in each of the samples along the traverse from the structurally lowest to structurally highest position.

DISCUSSION

As discussed earlier, amphibolites exhibit complex microstructures (Fig. 2-13). Older anhedral actinolite grains display inclusion trails at a high angle to the main foliation. The main foliation is defined by euhedral hornblende

Figure 2-11. ⁴⁰Ar/³⁹Ar age spectra from muscovites JM 2 (A), JM 5 (C) and splits of JM 4 (B, D) and JM 7 (E, F). JM 2 is the structurally lowest sample and is plotted as a reference on B-F. Age spectra A-F show gradients ranging from 200 Ma to 1400 Ma. These micas record terminal ages ranging from 1320 to 1400 Ma. Muscovite samples from the structurally highest samples generally yield older ages than the structurally lowest samples.

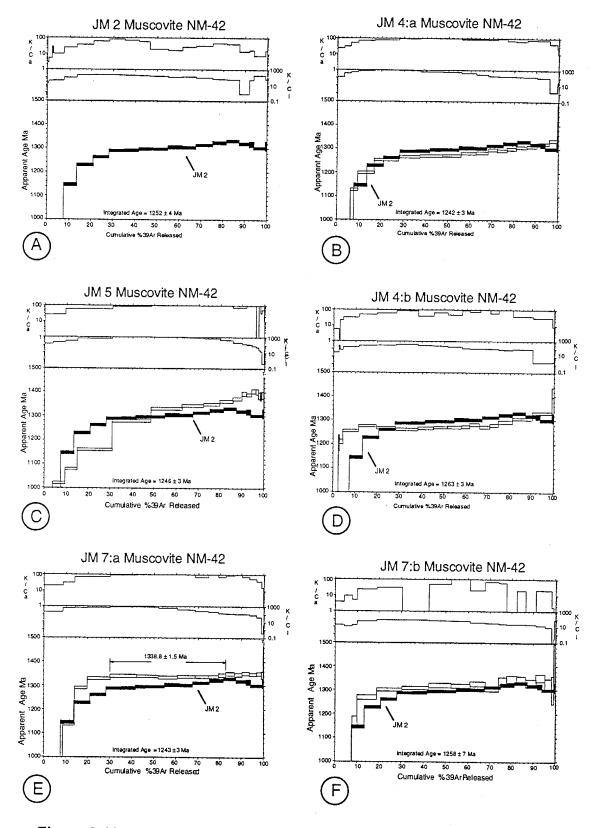


Figure 2-11.

Figure 2-12. ⁴⁰Ar/³⁹Ar age spectra from splits of muscovites JM 8 (A, B) and CP 1-2b (C, D). JM 2 is the structurally lowest sample and is plotted as a reference on B-F. Age spectra A-D show gradients ranging from 200 Ma to 1400 Ma. These micas record terminal ages ranging from 1320 to 1400 Ma. JM 8 shows a plateau age of 1333 ±1.9. Muscovite samples from the structurally highest samples generally yield older ages than the structurally lowest samples. CP 1-2b (D) is the structurally highest sample. The age discordance between JM 2 and CP 1-2b is about 60 my and is interpreted to have resulted from slow cooling along an average geothermal gradient.

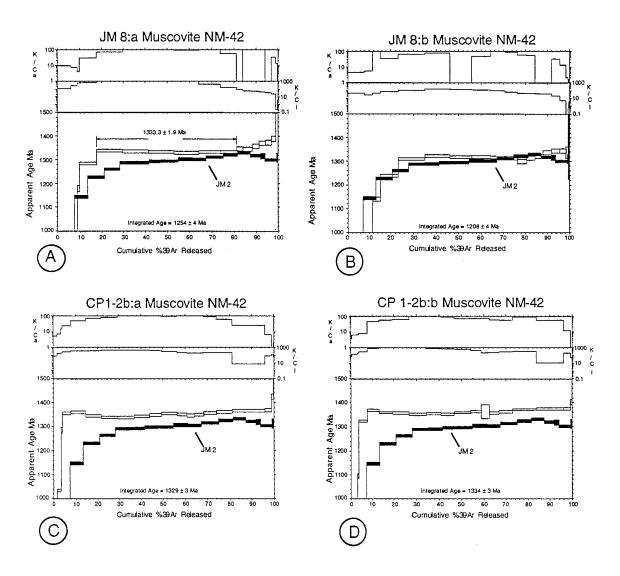


Figure 2-12.

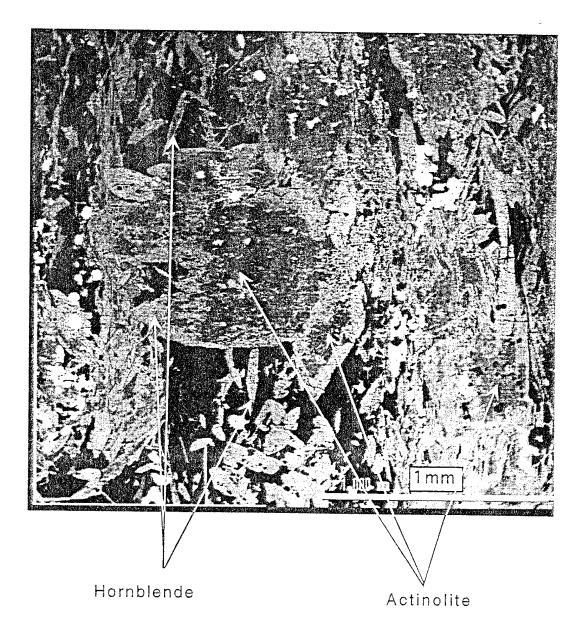


Figure 2-13. Backscattered electron image of ML 10-8 amphibolite showing the relationship between actinolite (dark) and hornblende (light). Note the high number of inclusions in the actinolite, the sharp boundary between the actinolite cores and the hornblende rims and the mutually cross-cutting relationship between the hornblende rims and bladed crystals.

grains. Hornblende also rims the actinolite grains. Hornblende rims and euhedral hornblende grains exhibit mutually crosscutting relationships, indicating contemporaneous growth (Fig. 2-13). Two different foliation orientations defined by 1) inclusion trails in the actinolites and 2) the nearly perpendicular hornblende crystals suggest that mineral growth occurred during two separate tectonic events. A miscibility gap is inconsistent with the structural and chemical data (c.f. Klein, 1968) and can not explain the sharp boundary between actinolite and hornblende of this particular composition. The chemical compositions of the ferro tschermakitic hornblendes and actinolites analyzed are inconsistent with those expected to co-exist and are not likely to be equilibrium pairs (Klein, 1968). Spear (1981) demonstrates that amphibole composition varies with temperature: typically amphiboles that form during a higher-grade metamorphic event have a higher Al and lower Si composition than observed in amphiboles formed during lower-grade metamorphic events. The hornblende studied here contains a significantly higher amount of Al and a lesser amount of Si then does the actinolite (Table 2-1). The hornblende compositions indicate formation under amphibolite facies conditions, while the actinolites record greenschist facies conditions (c.f. Spear, 1981). The structural and chemical relationships therefore indicate that the hornblende is younger than the actinolite and suggest two distinct periods of growth rather than simple prograde amphibole growth.

These data place constraints on the relative age of, and the metamorphic conditions during, amphibole growth. The ⁴⁰Ar/³⁹Ar data is used to further

constrain the timing of mineral growth. The K/Ca ratio can be obtained from both microprobe analyses and from the 37Ar/39Ar ratio measured during the $^{40}\mathrm{Ar}/^{39}\mathrm{Ar}$ step-heating experiment. Comparing these values can give an indication of how meaningful the $^{40}\mathrm{Ar}/^{39}\mathrm{Ar}$ ages are. The microprobe and argon results yield nearly identical K/Ca ratios for the hornblende samples. Amphibole separates have elevated K/Ca ratios (calculated from the ³⁷Ar/ ³⁹Ar ratio) for the first few low temperature steps (Appendix 1). These high K/Caratios are common in amphiboles and may be related to exsolution features (Harrison and Fitzgerald, 1986; McDougall and Harrison, 1988) or degassing of fine-grained alteration phases (Onstott and Peacock, 1987). The average hornblende K/Ca derived from microprobe data is 0.036 and that derived from averaging the K/Ca from the flat portions of the hornblende age spectra is 0.04. This suggests that the gas derived during these steps is probably not contaminated by fine-grained K-rich phases and therefore may yield meaningful apparent ages.

In contrast, the actinolite K/Ca ratio from the microprobe data is 0.01 and from the age spectra varies from approximately 0.02 to 0.23 (Appendix 1). The high K/Ca ratios in the actinolite age spectra are interpreted to suggest that the actinolite separates are contaminated by varying amounts of fine-grained K-rich phases throughout the age spectra. The low temperature steps with young apparent ages and high K/Ca ratios are likely a result of degassing of fine-grained K-rich phases within the amphiboles. Several samples exhibit low K/Ca in the first 5 to 15% of the age spectra and corresponding young age

steps, (e.g. ML 14-18-1 actinolite and hornblende, Fig. 2-10). This behavior is not well understood and is clearly not a result of a fine-grained K-bearing phases; it may be related to argon loss by chemical re-equlibrium as observed in complexly zoned amphiboles (Wartho, 1995). Samples with low K/Ca ratios in the early portions of the age spectra, as mentioned earlier, are all samples irradiated in irradiation NM-29. This problem is addressed below.

It is possible to evaluate both the amount of gas derived from each sample and the effectiveness of sample separation by examining the moles ³⁹Ar/mg for each amphibole sample with respect to irradiation number (Table 2-2). Samples from NM-29 have abundant young ages in the early portions of the age spectrum, without the corresponding rise in K/Ca seen in samples from NM-19 and NM-42. A zero age high-potassium contaminant could explain the low ages and low K/Ca ratios in the early portion of the age spectra. In this scenario, contamination must have occurred after separation, since only NM-29 samples are contaminated.

Fine-grained (300 mesh-size) actinolite and hornblende crystals were hand-picked under a binoculor low-power microscope. Considering the microstructural complexity of the two amphiboles, it is nearly impossible to obtain pure separates of each phase. The effectiveness of sample separation can be evaluated by comparing the hornblende: actinolite ratio of K_2O contents from the probe data, with the hornblende: actinolite ratio of moles $^{39}Ar/mg$ derived from $^{40}Ar/^{39}Ar$ analyses from the same sample (Table 2-2). A ratio of 3.6 is calculated from the microprobe data. Ratios between 0.3 and 1.2 are

obtained for amphibole pairs analyzed from NM-29 and NM-42. For example, samples ML 9-6 and ML 6-7 have ratios of 1.4 and 1.2 respectively, indicating that the actinolite and the hornblende were somewhat concentrated but are not pure separates. Sample ML 9-3b has a ratio less than one probably indicating that the separates are not pure, and possibly that ML 9-3b actinolite is as much as 50% hornblende and vice-versa. The effect of the samples being concentrates rather than pure separates is a more complex age spectra. For example, the young ages observed early in the hornblende age spectra could be explained by a small percentage of actinolite in the hornblende concentrate since in general the actinolite age spectra show more young apparent ages than the hornblende spectra. As a result, age spectra interpretations are somewhat speculative since the relative contribution of each different age mineral is unknown.

A much higher percentage of ³⁹Ar released yields young ages in the actinolite spectra compared to the higher ages in the hornblende age spectra. This is likely a result of mineral inclusions, possibly feldspar and biotite (Fig. 2-13). These inclusions are probably the source for the excess argon released in the first steps of the actinolite age spectra in ML 10-8, thus explaining the anonymously old ages. However, in most cases the first 10-60% of the actinolite age spectra show young ages which correspond to a two order of magnitude smaller radiogenic yield and an elevated K/Ca. Due to the low radiogenic yields, high K/Ca ratios and abundant inclusions, limited geologic information can be derived from the actinolite age spectra. Since the bulk amphiboles consist of an unknown percentage of actinolite, hornblende and

other minerals, both the bulk amphibole and the actinolite age spectra will not be used to constrain the thermochronology. The hornblende age spectra are also complex and probably related to the hornblendes being concentrates rather than pure separates; however, the majority have relatively flat portions of the age spectra around 1400 Ma with a flat corresponding K/Ca plot. For example, approximately 80%, 65% and 60% of the ³⁹Ar released for hornblende separated from ML 14-18, ML 10-8 and ML 6-7 respectively, yield apparent cooling ages of 1409 ± 10 Ma, 1329 ± 10 and 1397 ± 6 Ma with a constant K/Ca ratio for each step (Figs. 2-6E, F; 2-7D; 2-8C). However, several samples have a highly complex structure with a flat corresponding K/Ca (e.g., ML 9-3b, Fig. 2-9A). The fact that several samples have complex forms for the entire spectra suggests that the flat portions of the age spectra observed in the majority of samples are not a result of a simple homogenization of the ⁴⁰Ar/³⁹Ar ratio. Much of the hornblende age spectra complexity is probably related to the variable amounts of actinolite within each hornblende concentrate. Despite the overall age spectra complexity, the frequency of this recurring age between 1350 and 1410 Ma is more than coincidence, and at least suggests that this 1350-1400 Ma is a minimum time of closure for the hornblendes.

The flat portions of the hornblende age spectra clearly do not define rigorous plateau ages; however between 50 and 70% of the cumulative 39 Ar released consistently yields cooling ages of between 1330 and 1450 Ma (Figs. 2-6E, F; 7D, E, F). These relatively flat portions of the 40 Ar/ 39 Ar age spectra

suggest that the foliation-forming hornblende cooled through its closure temperature ca. 1400 Ma; in particular, ML 14-18-2 hornblende exhibits a true plateau of 1409 ± 10 Ma. The last 5-10% of the gas released yields terminal ages up to ~1450 Ma and is may be the best estimate for the age of hornblende growth; this conclusion is consistent with that of Thompson et al., (1991). The variation in ages recorded by the flat portion of the age spectra is likely a result of differences in closure temperature of the hornblendes, which could be caused by either fine-grained intergrowths (Onstott and Peacock, 1987) or a variation in Mg and Fe content (O'Nions et al., 1969). The first of these options is the most likely, since fine-grained intergrowths have been postulated to have a greater affect on closure temperatures in high-grade metamorphic hornblendes (Onstott and Peacock, 1987). The probe data show little Mg and Fe variation between samples suggesting that composition is not the main variable controlling closure temperature, consistent with the studies of Harrison, (1981). If regional temperatures were high enough (450-550°C) to grow the tschermakitic hornblende prior to 1400 Ma, than temperatures should be high enough to reset the older actinolites with respect to argon. If so, one would not expect to see through the amphibolite facies event, and the older actinolites should yield younger or similar ⁴⁰Ar/³⁹Ar apparent ages.

The hornblende age spectra therefore record a period of cooling through the variable hornblende closure temperatures at ca. 1400 Ma, as well as a minimum age of mineral growth. Petrographic observations of amphibolites suggest syn-tectonic hornblende growth. The age of the hornblende growth

can be further constrained by U-Pb ages from the 1.4 Ga Priest Pluton. Since the Priest Pluton (Fig. 2-1) is not highly deformed, syn-tectonic hornblende growth most likely occurred prior to emplacement, dated at 1427± 10 Ma (U-Pb zircon) (Bauer et al., 1993). The Priest Pluton is postulated to have been emplaced in the late stages of regional deformation (this manuscript) or in the absence of deformation (Karlstrom et al. 1996; Thompson et al., 1996; Karlstrom and Williams, 1995; Bauer and Williams, 1994; Karlstrom and Bowring, 1993). Hence, the age of hornblende growth is best constrained to 1440-1450 Ma (Fig. 2-8F), pre-dating emplacement of the Priest Pluton and corresponding to the terminal ⁴⁰Ar/³⁹Ar hornblende ages.

Two scenarios can explain the microstructural and 40 Ar/ 39 Ar data in the context of the regional tectonic history (Fig. 2-14). In the first model, the data are explained in the context of a non-deformational heating event ca. 1400 Ma; the second model involves regional deformation and metamorphism ca. 1400 Ma.

Model one requires no deformation at 1.4 Ga. Hornblende would grow syntectonically at ca. 1650 Ma (Mazatzal orogeny), and possibly undergo enlargement by overgrowth of hornblende with an identical composition during metamorphism at ca. 1400 Ma. If this is the case, an earlier deformation and metamorphic event (pre-1650 Ma) would have to be called upon to produce the older actinolites. Additionally, a younger deformational event, post-1400 Ma, would also be needed.

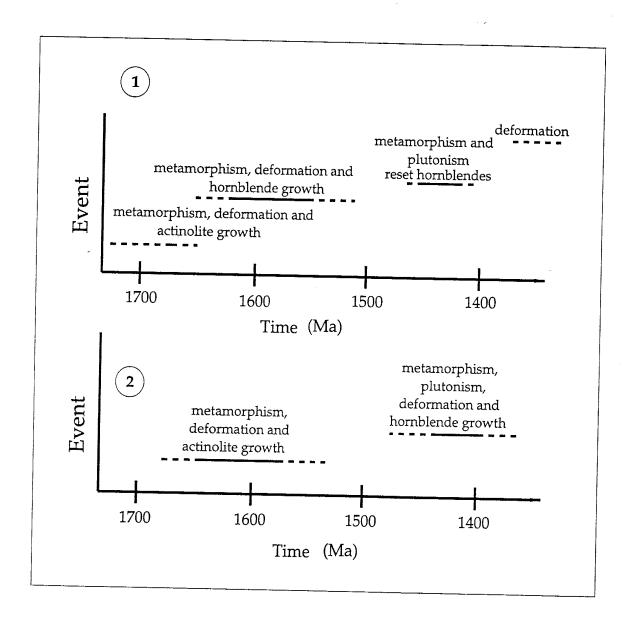


Figure 2-14. Schematic plots illustrating approximate timing of deformation and metamorphism using models 1 and 2. Model 1 requires three distinct events to explain the observed data in the context of 1.4 Ga static heating, whereas model 2 requires two metamorphic and deformational events at 1.6 and 1.4 Ga.

If metamorphism was not accompanied by deformation ca. 1400 Ma, then later deformation must be involved to produce the observed microstructures. The available microstructural data do not support a post 1400 Ma event as needed in model one; for example, microstructures in the Blue Springs schist and the adjacent quartz mylonites (see Part I) show no evidence of annealing. Monocrystalline quartz ribbons (Fig. 1-19), fine-grained muscovite and intracrystalline cataclasite, would be expected to recrystallize during the 500° C static reheating event that previous workers (Thompson et al., 1991) have suggested occurred at 1400 Ma. Although possible, there is no evidence to support a pre-1600 Ma regional deformational event or a post-1400 Ma event in the Capilla Peak area. The absence of a well-developed tectonic foliation in the Priest Pluton is therefore the result of emplacement during the late stages of deformation.

The second scenario is more consistent with the observed data: 40 Ar/ 39 Ar ages from foliation-forming hornblende and micas whose microstructures suggest that they grew syn-tectonically, probably at ca. 1440-1450 Ma, and cooled through their closure temperatures between 1330 and 1400 Ma.

THERMAL HISTORY

The thermal history for the Manzano Mountains from 1.65 Ga to the present can be evaluated from a combination of microstructural, electron microprobe and ⁴⁰Ar/³⁹Ar age spectra data. Figure 2-15 is a schematic diagram showing two different thermal histories. Number 1 is consistent with model 1 (discussed earlier), and number 2 is consistent with model 2, the metamorphic and deformational history of the Capilla Peak area proposed here.

Thermal history number 1 includes amphibolite grade metamorphic conditions around 1.6 Ga (Grambling, 1988). Data constraining regional temperatures between 1600 and 1400 Ma are limited. However, Karlstrom et al. (1996) proposed that amphibolite facies conditions existed around 1600 Ma, and a period of slow cooling followed. Abundant ⁴⁰Ar/³⁹Ar, Rb-Sr and K-Ar ages indicate that a reheating event occurred around 1400-1450 Ma. Thompson et al. (1996) suggest that temperatures adjacent to the Priest pluton of 500-550°C existed at 1438 Ma, and Karlstrom et al. (1996) suggest that crustal temperatures remained relatively constant between 1400 and 1300 Ma.

Data presented in this paper are consistent with thermal history number 2 (Fig. 2-15).

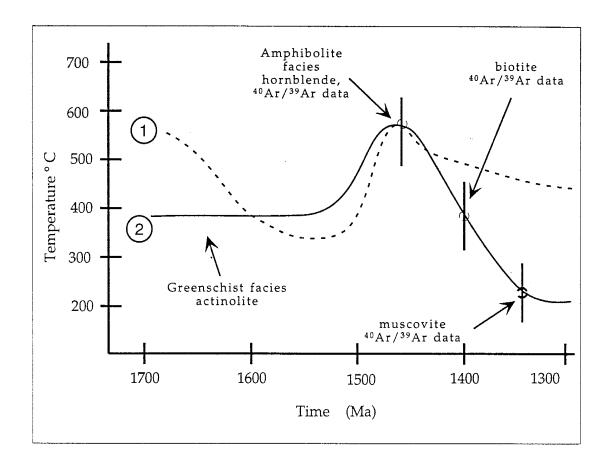


Figure 2-15. Schematic temperature-time plots for models 1 and 2. Model 1 is based on data from other workers, from around the Priest Pluton. Model 2 is based on geochronology and probe data from the Capilla Peak area. Points represent locations on the curves that are constrained by geochronology, with error bars representing approximate proposed temperature range.

Amphiboles

As mentioned earlier, evidence for the age of the actinolites is weak. However, based on 1.6 Ga terminal ages from bulk amphibole and actinolites (e.g. ML 10-6, Fig. 2-9 B and D), and the fact that the actinolites are older than

the 1.4 Ga hornblendes, it is proposed that actinolites formed at 1.6 Ga. Contrary to the conclusions of Marcoline et al. (1995), Bauer and Williams (1994) and Grambling (1988), syntectonic growth of pre-1450 Ma actinolites (probably 1600-1650 Ma) suggest that local conditions were greenschist facies, with temperatures around 300-450°C, instead of amphibolite facies conditions. Thompson (1996) and Bowring and Karlstrom (1990), suggest that crustal temperatures in the Manzano Mountains remained constant at around 300°C until 1450 Ma. Amphibole mineral stabilities (Spear, 1981) and closure temperatures suggest that metamorphic conditions at 1.45 Ga were amphibolite facies.

Biotites

Concordant ca. 1400 Ma hornblende and biotite age spectra are interpreted to suggest that regional temperatures dropped rapidly (>10° C/Ma) from ca. 500-550° C (hornblende closure temperature) through 300-350° C (biotite closure temperature) around or shortly after 1400 Ma (Fig. 2-10E, F). However a third biotite (ML 10-8) from the same area as this hornblende-biotite pair yields an age of 1276 \pm 6 Ma (Fig. 2-7F), ~130 Ma younger than ML 10-2

and ML 6-10 biotites. This younger biotite has a much smaller grain-size than the other two biotites, and in turn would be expected to have a lower closure temperature. The result of ML 10-8 biotite having a lower closure temperature than ML 10-2 and ML 6-10 biotites would be an amplification of any age discordance.

Muscovites

Inboard thermal effects of the Grenville Orogeny have been recorded in muscovites from discrete shear zones within the Priest Pluton (Heizler et al., 1996) and in northern New Mexico (Karlstrom et al., 1996). The muscovite age spectra from the Blue Springs schist may have experienced inboard thermal effects of the Grenville Orogeny however, do not preserve an 1100 Ma signature. The muscovites do, however, appear to be affected by later thermal events contributing to the abundant young ages early in the age spectra, probably related to late Paleozoic burial.

As indicated previously, muscovites were sampled from the base to the ridge of the Manzano Mountains within the Blue Springs schist (Fig. 2-2) in order to investigate the spatial variability of ages. The muscovites yield age gradients and terminal ages between 1320 and 1400 Ma (Fig. 2-11, 2-12). The muscovites have grain sizes ranging from 0.5 to 10 microns. These muscovite age spectra are similar in form and slightly younger in age to those observed in deformed muscovites near the Priest Pluton (Heizler et al., 1996). Different mechanisms proposed to explain well defined age gradients in muscovite are

summarized in Heizler et al. (1996). However, in this case the young ages are most likely related to late Paleozoic burial, and the age gradients to a mixed grain size population.

Grain size reduction is observed in the muscovites in the form of intragranular cataclasite and recrystallization. Even though this deformation has reduced the grain size to ~0.5 microns within the cataclasite zones (c.f. Goodwin and Wenk, 1990), and essentially reduced the effective diffusion radii, associated argon loss appears to be negligible. If grain-size reduction were responsible for the observed age gradients we would expect to see more pronounced age gradients in the structurally higher positions where, in general, there is a greater amount of intra- and intergranular cataclasis.

The muscovite ages show a generally increasing age trend of 60 m.y. from the structurally lowest sample (JM-2) to the structurally highest sample (CP 1-2b). The observed age discordance is not related to grain size, either by different muscovite closure temperatures or by recoil. If the observed age discordance were related to different closure temperatures, assuming the submicron to 10 micron grain size is the controlling factor for muscovite closure with respect to argon, then we would expect the oldest sample to have the coarsest grain size, and the youngest the finest grain size. However, the structurally highest samples have about an order of magnitude smaller grain size than the structurally lowest samples. For example, Table 2-3 shows the relationship between grain-size and closure temperature; the closure temperature of a 0.5 micron muscovite is 217° C while a 10 micron muscovite

grain-size	T_c	T_C
μm	Α	В
0.05	150	170
0.05	159	170
0.5	203	217
1.0	219	233
1.5	228	244
2.0	236	250
2.5	241	257
3.0	246	262
3.5	250	265
4.0	253	269
4.5	256	273
5.0	259	276
5.5	262	278
6.0	264	281
6.5	266	283
7.0	268	285
7.5	270	287
8.0	272	289
8.5	274	291
9.0	275	293
9.5	277	294
10.0	278	296
20.0	299	317
50.0	328	349

Table 2-3. Chart showing the relationship between grain-size in muscovite and closure temperature (°C). Column A gives the Tc calculated using a cooling rate of 0.3° C/Ma and B is calculated using 1° C/Ma; all calculations used an E of 40 Kcal/mole and a D_o of 6 X 10^{-7} cm²/s.

would be expected to have a closure temperature of 296° C. In this situation we would expect that the age of the fine-grained muscovites from the structurally highest position would be on the order of 30 Ma younger than the age of a muscovite from the structurally lowest position (assuming 1°C/Ma cooling), and 75 Ma younger if the samples were at the same structural level. Since the entire range of grain-sizes are observed in all samples (in varying amounts) and since the grain-size model predicts the opposite age trend of that observed, we interpret that this grain size distribution is not responsible for the observed age discordance; and actually would have the effect of minimizing the observed age discordance. However the grain-size probably is responsible for the age gradients observed in the muscovite age spectra since the fine-grained material would be expected to degas first (West et al., 1993).

The age spectra do not resemble those interpreted to have been affected by recoil (Harrison and Fitzgerald, 1986), hence, the phenomenon of older ages related to the recoil of ³⁹Ar into cataclasite (as observed by Goodwin and Renne, 1991) is unlikely; rather the data support a simple cooling model. Assuming an average geothermal gradient (25-30°C/m.y.) slow cooling on the order of 0.3 - 1°C/Ma would be consistent with the area undergoing uplift and erosion, explaining the observed muscovite age trend. Muscovite data, in conjunction with hornblende and biotite data, therefore are interpreted to suggest that a period of early (1450-1350 Ma) rapid and late (1350-?) slow cooling occurred.

CONCLUSIONS

The combination of ⁴⁰Ar/³⁹Ar age spectra, field observations and microstructures indicate that at least two episodes of regional deformation, metamorphism and mineral growth occurred in the Manzano Mountains.

Deformation and mineral growth accompanied both amphibolite grade metamorphism at 1400-1450 Ma, and an older period of greenschist-facies metamorphism. There is no direct way to date the early event; however, the presence of inclusion trails in actinolites that are nearly perpendicular to the hornblende tectonic foliation is interpreted to suggest that the actinolite formed syntectonically during a prior lower grade metamorphic and deformational event, probably at 1650 Ma.

Amphibole compositional data suggest that regional temperatures reached 550-650°C (cf. Spear, 1981) at 1450 Ma and cooled to between 400 and 550°C by 1400 Ma. Muscovite and biotite ⁴⁰Ar/³⁹Ar data are interpreted to suggest that regional temperatures cooled rapidly to 200-300°C by about 1350 Ma. Discordant muscovite ⁴⁰Ar/³⁹Ar age spectra from different structural levels indicate that a period of protracted slow cooling occurred after 1350 Ma.

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Appendix 2-1. Appendix 2-1 is a table containing all of the ⁴⁰Ar/³⁹Ar data organized by irradiations NM-19, NM-29 and NM-42. The run ID # is the lab sample number with the adjacent capital letter representing each successive heating step. For each heating step at the given temperature a value of ⁴⁰Ar/³⁹Ar, ³⁷Ar/³⁹Ar, ³⁶Ar, ³⁹Ar moles, K/Ca, Cl/K, % ⁴⁰Ar*, % ³⁹Ar and age is given.

Run ID#	Temp	⁴⁰ Ar/ ³⁹ Ar	³⁷ Ar/ ³⁹ Ar	³⁶ Ar/ ³⁹ Ar	³⁹ Ar moles	K/Ca	CI/K	% ⁴⁰ Ar*	% ³⁹ Ar	Age	± Err
MH-9-35 Bt	ılk Amph	ibole, JM, J=	0.01442337	N	IM-19	2.	6 mg				
2216-01A	700	809.484	14.900	1.8064	8.9E-17	0.034	0.7953	34.1	3.11	2895	56
2216-01B	800	195.392	17.879	0.3127	4.9E-17	0.029	0.2511	52.7	4.82	1642	37
2216-01C	900	113.061	22.331	0.1659	7.2E-17	0.023	0.1246	56.6	7.35	1180	33
				0.0387		0.014	0.1564	88.4	15.67	1472.1	7.3
2216-01D	1000	98.916	36.844		2.4E-16						
2216-01E	1040	170.375	88.551	0.1018	4.3E-16	0.006	0.2705	82.3	30.65	1995.8	7.3
2216-01F	1070	101.787	73.743	0.0552	3.3E-16	0.007	0.1098	83.9	42.24	1448.8	7.5
2216-01G	1100	81.992	50.294	0.0293	2.0E-16	0.010	0.0626	89.4	49.24	1302	10
2216-01H	1130	65.229	67.298	-0.0184	4.8E-17	0.008	0.0746	108.3	50.93	1267	33
		81.823	59.032	0.0065	8.7E-17	0.009	0.0774	97.6	53.98	. 1383	19
2216-011	1160								59.93	1491.4	9.2
2216-01J	1190	88.802	66.073	-0.0012	1.7E-16	0.008	0.0971	100.4			
2216-01K	1220	107.104	59.485	0.0186	2.3E-16	0.009	0.1258	94.9	67.85	1627.9	7.6
2216-01L	1250	92.056	50.923	0.0172	5.4E-16	0.010	0.0907	94.4	86.75	1466.2	4.4
2216-01M	1280	103.370	76.700	0.0249	2.6E-16	0.007	0.1005	92.9	95.91	1567.7	6.1
				0.0500	1.1E-16	0.003	0.2295	90.5	99.81	1996	11
2216-01N	1350	155.104	193.623								162
2216-010	1700	307.758	196.068	0.2585	5.4E-18	0.003	0.1529	75.2	100.00	2647	
	t	otal gas age	n:	=15	2.9E-15	0.010	0.0094			1613	11
MILO C D.		hala 194 1-6	01445000		NM-19	2	.2 mg				
			0.01445203					40.7	7 00	0.50	1.4
2217-01A	700	85.519	3.352	0.1456	2.5E-16	0.152	0.0760	49.7	7.32	863	14
2217-01B	800	32.986	1.828	0.0317	1.2E-16	0.279	0.0216	71.5	10.81	529	19
2217-01C	900	39.257	3.977	0.0266	1.8E-16	0.128	0.0236	79.9	16.22	674	10
			19.220	0.0110	5.0E-16	0.027	0.0372	93.3	30.84	916.0	4.0
2217-01D	1000	49.048									5.9
2217-01E	1040	68.309	43.930	0.0267	3.7E-16	0.012	0.0542	88.4	41.60	1131.9	
2217-01F	1070	68.700	40.851	0.0195	4.0E-16	0.012	0.0307	91.6	53.42	1166.6	5.7
2217-01G	1100	65.876	28.754	0.0110	2.9E-16	0.018	0.0258	95.0	61.88	1162.3	6.4
2217-01H	1130	48.011	32.115	0.0319	9.1E-17	0.016	0.0246	80.3	64.58	799	20
					6.9E-17	0.014	0.0314	93.0	66.62	1049	26
2217-011	1160	58.700	36.446	0.0138							8.5
2217-01J	1190	79.594	34.364	0.0224	2.0E-16	0.015	0.0377	91.7	72.61	1298.9	
2217-01K	1220	86.634	35.441	0.0167	3.0E-16	0.014	0.0544	94.3	81.40	1406.4	5.9
2217-01L	1250	89.469	39.657	0.0166	3.7E-16	0.013	0.0575	94.5	92.40	1440.2	4.7
		84.558	63.269	0.0274	1.3E-16	0.008	0.0583	90.4	96.14	1342	11
2217-01M	1280							86.0	99.85	1483	11
2217-01N	1350	102.653	131,754	0.0487	1.3E-16	0.004	0.0842				
2217-010	1700	238.371	214.083	0.5818	5.2E-18	0.002	0.1395	27.9	100.00	1214	318
	t	otal gas age	n	=15	3.4 E-1 5	0.041	0.0784			1120.0	8.6
MH-13-9b	Bulk Ams	phibole, JM,	J=0.0144979	9 1	NM-19	2	.9 mg				
2218-01A	700	36.405	1.216	0.0322	1.4E-15	0.419	0.0190	73.8	10.36	593.8	2.3
				0.0063	9.6E-16	1.075	0.0052	94.4	17.23	688.6	2.3
2218-01B	800	33.939	0.475								
2218-01C	900	36.876	0.778	0.0080	1.5E-15	0.656	0.0047	93.5	27.81	731.6	1.8
2218-01D	1000	52.443	5.868	0.0044	2.6E-15	0.087	0.0112	97.5	46.25	1000.6	1.8
2218-01E	1040	64.694	11.537	0.0061	1.4E-15	0.044	0.0145	97.2	56.11	1168.8	2.6
2218-01F	1070	66.668	10.983	0.0044	1.5E-15	0.046	0.0112	98.0	66.69	1202.5	2.3
						0.061	0.0078	97.5	73.27	1047.9	2.4
2218-01G	1100	55.734	8.304	0.0047	9.2E-16						
2218-01H	1130	41.519	7.301	0.0063	4.1E-16	0.070	0.0074	95.5	76.21	819.4	5.1
2218-011	1160	65.399	12.902	0.0133	3.5E-16	0.040	0.0125	93.9	78.74	1149.1	4.8
2218-01J	1190	67.882	13.305	0.0058	6.2E-16	0.038	0.0166	97.5	83.16	1213.3	3.6
				0.0074	9.1E-16	0.032	0.0227		89.63	1451.4	3.3
2218-01K		87.425	16.016						96.95	1525.5	3.0
2218-01L	1250	93.942	16.653	0.0075	1.0E-15	0.031	0.0229	97.6			
2218-01M	1280	92.241	17.942	0.0034	2.4E-16	0.028	0.0184		98.64	1520.1	9.0
2218-01N		87.684	17.739	0.0129	1.6E-16	0.029	0.0180	95.6	99.75	1435	14
2218-010		83.577	20.882	0.0579	3.5E-17	0.024	0.0054	79.5	100.00	1217	64
2210 010		total gas age		=15	1.4E-14	0.226	0.3059			1039.2	2.9
ML-13-10c	Bulk An	nphibole, JM,	J=0.014502		NM-19		2.4 mg				
2219-01A		109.045	2.477	0.1511	5.7E-16	0.206	0.0364	59.0	3.20	1189.5	7.7
2219-01B		60.748	2.047	0.0373	2.9E-16	0.249	0.0086		4.80	979.2	6.3
					4.7E-16	0.131	0.0091		7.44	795.3	6.0
2219-01C		52.075	3.883	0.0469							
2219-01D	1000	76.850	18.522	0.0120	1.8E-15	0.028	0.0042		17.59	1306.5	2.4
2219-01E	1040	81.610	25.759	0.0137	1.3E-15	0.020	0.0052	95.0	25.02	1359.5	3.5
2219-01F		80.758	18.784	0.0077	2.3E-15	0.027	0.0029	97.1	37.78	1370.7	2.2
				0.0063	2.7E-15	0.031	0.0022		53.07	1380.8	2.2
2219-01G		81.164	16.555								
2219-01H	1130	68.643	19.251	0.0112	2.7E-16	0.027	0.0038		54.57	1202.3	6.3
2219-011	1160	79.749	20.744	0.0097	4.5E-16	0.025	0.0041	96.4	57.09	1351.0	4.1
2219-01J		83.978	19.126	0.0071	2.7E-15	0.027	0.0028		72.39	1412.0	2.0
									85.29	1488.1	2.1
2219-01K		90.426	20.066	0.0069	2.3E-15	0.025	0.0031				
2219-01L	1250	91.978	21.577	0.0081	2.4E-15	0.024	0.0038	97.4	98.76	1501.7	2.2
2219-01M		100.484	30.605	0.0167	1.3E-16	0.017	0.0058	95.1	99.52	1568.5	9.3
2219-01N		111.096	31.540	0.0550		0.016	0.0065		99.81	1561	26
									100.00	1530	41
2219-010		126.476	33.249	0.1164		0.015	0.0084		100.00		2.9
		total gas age	1	n=15	1.8E-14	0.038	0.0749	,		1376.1	2.9
ML-14-18-	1 Bulk A	mphibole, JM	, J =0.01447		NM-19		3.4 mg				
2220-01A	700	96.331	3.808	0.1656	5.6E-16	0.134	0.0719	49.2	4.12	941.7	8.0

Run I	D# -	Temp	40Ar/ ³⁹ /	A - 3720									
2220	-01B	800	66.93				oles	K/Ca	CI/K	% ⁴⁰ Ar		∕6 ³⁹ Ar	
2220	-01C	900	60.23			578 3.1E	-16 0	.183	0.0384				Age ± Err
2220		1000	70.76					.113	0.0360	60.	_	~ ~ -	34.0 7.3
2220-		1040	82.73					.028	0.0568	93.6	_		68.4 8.2
2220-		1070	82.35					017	0.0437	93.			12.7 2.6
2220-		1100	79.19	-0.0					0.0236	96.2			57.2 3.1
2220-		1130	70.22		,				0.0170	97.1			77.9 2.6
2220-	011	1160	74.24						0.0274	95.7	_ ''		9.2 2.4
2220-	01J	1190	84.502	-0.00					0.0208	96.4			5.3 5.3
2220-		1220	90.769						0.0253	97.0			
2220-		1250	91.472	27.43				023 (0.0262	97.3			
2220-0	21 M	1280	100.277	41.27				019 (0.0274	96.8			
2220-0		1350	127.848	64.56)12 (.0435	95.4			
2220-0	10	1700	126.349	63.08					.0596	88.5			
		to	otal gas ag	e	n≃15			008 0	.0548	85.6			750 35
841 44	401					1.4E-1	0.0	33 0	.0534			1355	99 19
2221 0	135 B	ulk Amp	ohibole, Ji	M, J=0.0144	4197	NM-19						100	3.3
2221-0 2221-0		700	64.848	0.890		_		2.3 m					
2221-0		800	74.885	0.326				-	.0175	87.7	4.	32 1081	2
2221-0	10	900	75.256	0.605			_		.0039	97.6	9.		
2221-0		1000	76.667	9.797					0035	97.7	18.8		
2221-0		1040	77.035	5.548			_		0079	98.1	36.0		
2221-01		1070	78.249	7.484	0.003		_		0052	98.8	46.9		
2221-01		1100	80.400	12.443					0066	98.8	52.9		
2221-01		1130	76.820	5.480					0072	98.2	67.2		
2221-01		160	77.786	4.025	0.0035				0027	99.1	77.7		_
2221-01		190 220	83.785	18.038	0.0074				0026	98.7	83.6		
2221-01			93.867	23.174	0.0077	1.2E-15			0124	97.4	86.0	4 1404.	
2221-01		250 280	97.149	22.429	0.0079			_	0185	97.6	89.1		
2221-01		350	99.586	32.487	0.0155	4.4E-16		_	170	97.6	98.4		
2221-01		700	104.863	48.527	0.0250				206	95.4	99.5		
	• ,		122.183	48.192	0.0550	8.4E-17	0.01			92.9	99.7	9 158	
		1012	al gas age		n=15	4.0E-14	0.23			86.7	100.00	1674	
ML-6-7 E	Bulk A	mnhibal		0.01447585			0.20	0.4	377			1357.5	
2222-01/	A 7	700	6, JM, J=			NM-19		2.8 mg					
2222-018	_	300	57.573	1.281	0.0881	1.0E-15	0.398		004				
2222-010		00	61.181	1.714	0.0333	4.9E-16	0.298		.	54.8	3.74		4.7
2222-010		000	53.092	2.092	0.0312	1.0E-15	0.244			33.9	5.53	1002.4	
2222-01E		40	62.316 68.781	10.163	0.0076	2.8E-15	0.050			32.6	9.26	~~~.	
2222-01F		70	72.375	11.717	0.0061	1.9E-15	0.044			6.4	19.49		
2222-01G	11		76.031	16.077	0.0068	1.3E-15	0.032			7.3	26.26		2.3
2222-01H	11		77.399	16.269	0.0062	2.8E-15	0.031	0.00		7.2	31.04	1266.8	2.7
2222-011	11		68.468	14.424	0.0046	4.6E-15	0.035	0.00		7.6	41.35	1316.1	2.1
2222-01J	11:		78.583	11.275	0.0065	7.3E-16	0.045	0.00		8.2 7 <i>.</i> 2	58.27	1339.0	1.9
2222-01K	122		82.157	16.971 17.499	0.0077	9.5E-16	0.030	0.00		7.1	60.93	1217.0	2.5
2222-01L	125		84.726		0.0056	4.0E-15	0.029	0.00		7.1 B.0	64.40	1342.2	2.5
2222-01M	128		85.293	18.508	0.0056	4.2E-15	0.028	0.00		3.0	78.90	1393.6	1.9
2222-01N	135		84.748	20.769	0.0081	1.3E-15	0.025	0.00		7.2	94.35	1424.2	1.9
2222-010	170		03.985	22.787	0.0181	1.5E-16	0.022	0.00		3.7	99.25	1422.2	2.5
			gas age	22.084	0.0830	5.3E-17	0.023	-0.00			99.81	1380	11
					=15	2.7E-14	0.060	0.120		, 	100.00	1381	32
ML-10-6 Bi	ulk An	nphibole	. JM. J=1	0.01451016								1278.5	2.3
	70	0 a	4.678	1.633		M-19	:	2.8 mg					
2223-01B	80		2.125	0.752	0.0734	1.4E-15	0.312	0.045	3 74	.4	404	44-4	
2223-01C	90	0 5	4.921	1.333	0.0114 0.0185	1.1E-15	0.678	0.015			4.81 8.88	1171.0	3.7
2223-01D	100	0 6	4.212	6.786		1.5E-15	0.383	0.017			14.30	1112.0	2.6
2223-01E	104		3.636	10.679	0.0055	3.6E-15	0.075	0.036		_	27.12	975.4	2.2
2223-01F	107	0 7	8.033	12.791	0.0052 0.0056	2.5E-15	0.048	0.034		_	35.90	1165.6	1.8
2223-01G	1100		5.630	13.541	0.0053	2.0E-15	0.040	0.027		_	43.07	1291.4	1.9
2223-01H	1130		8.715	10.738	0.0050	3.7E-15	0.038	0.024			56.17	1345.4	2.2
2223-011	1160	67	7.242	12.254	0.0051	1.7E-15	0.048	0.017			62.20	1438.5	2.2
2223-01J	1190		2.180	16.099	0.0031	6.9E-16	0.042	0.019			64.67	1355.9	2.2
2223-01K	1220	113	3.927	16.450	0.0071	1.4E-15	0.032	0.0278			69.61	1208.3	3.5
2223-01L	1250	112		17.192	0.0064	4.8E-15	0.031	0.0396		_ `	36.79	1507.9	2.6
2223-01M	1280		.885	19.046	0.0064	3.3E-15	0.030	0.0348			98.55	1741.4	2.1
2223-01N	1350			22.018	0.0606	3.2E-16	0.027	0.0399			9.67	1731.2	2.1
2223-010	1700	171	.604	26.263	0.1293	4.8E-17	0.023	0.0566			9.85	1745.4	5.1
		total ga	s age	n=1		4.3E-17	0.019	0.0566		_	0.00	1891	28
At -10				1	-	2.8E-14	0.100	0.1893			2.20	1943	35
ML-10-8-1 B	ulk An	nphibole	9, JM, J=	0.01442939	A) as	-10						1416.9	2.4
424-UIA	700	68	.189	0.877	0.0410			3 mg					
224-01B 224-01C	800	64.	.237	0.710	^ ^	3.3E-15	0.582	0.0173	82.2		5.24	1060 4	
004	900		710	0.839		2.2E-15 4.6E-15	0.719	0.0092	96.4		8.68	1069.1 1152.0	2.5
UID	1000	64.	218	3.870		8.2E-15	0.608	8800.0	97.5			1127.9	2.1
						13	0.132	0.0114	98.8			1172.3	1.8
													1.6

Run ID#	Temp	40Ar/ ³⁹ Ar	³⁷ Ar/ ³⁹ Ar	³⁶ Ar/ ³⁹ Ar	39.						
2224-011								% ⁴⁰ Ar*	% ³⁹ Ar	Age	± Err
2224-011			7.899	0.0033	6.6E-15	0.065	0.0110	98.6			
		74.138	10.648	0.0038	6.5E-15	0.048					1.9
2224-010	G 1100	72.176	7.000	0.0024	8.1E-15	0.073					2.0
2224-01	H 1130	67.015	3.105	0.0013			0.0050			1278.3	1.6
2224-011	1160	68.319	3.060		6.6E-15	0.164	0.0030		73.74	1215.1	2.0
2224-01.				0.0014	4.3E-15	0.167	0.0030	99.4	80.69		1.7
2224-01		75.162	8.816	0.0042	1.8E-15	0.058	0.0071	98.3			
		95.510	19.838	0.0064	2.4E-15	0.026	0.0206	98.0			2.2
2224-011		93.243	17.718	0.0056	3.8E-15	0.029					2.2
2224-01	M 1280	96.754	21.310	0.0076			0.0140	98.2	93.29	1519.2	2.1
2224-011		126.384			3.2E-15	0.024	0.0170	97.7	98.35	1551.8	2.3
2224-010			41.245	0.0183	8.9E-16	0.012	0.0358	95.7	99.77		
#224-01C		120.002	33.723	0.0560	1.5E-16	0.015	0.0229	86.2	100.00		3.5
		total gas age	r	า=15	6.3E-14	0.173	0.2422	30.2	100.00	1648	11
						27.70	0.2422			1277.6	2.0
ML-10-2	Biotite, Ji	M, J=0.0143	9601	i	NM-19		4.0				
2225-014	600	34.986	0.143	0.0239			1.2 mg				
2225-01E		63.891			1.9E-15	3.578	0.0028	79.7	76.83	609.1	2.2
2225-010			0.047	0.0182	1.4E-15	10.957	0.0007	91.6	132.41	1102.3	
	-	72.951	0.020	0.0093	3.5E-15	25.510	0.0000	96.2			2.5
2225-01		76.680	0.007	0.0022	1.3E-14	70.059			272.15	1259.7	2.0
2225-01E	800	75.710	0.006	0.0005			0.0000	99.1	802.90	1333.5	1.7
2225-01F	870	76.197	0.003		6.2E-15	90.802	0.0000	99.8	1051.44	1327.7	2.0
2225-01G				0.0004	2.4E-14	152.651	0.0001	99.8	2024.58	1334.3	1.7
2225-01H		76.724	0.003	0.0003	3.1E-14	146.705	0.0000	99.8	3269.09	1340.9	
		79.823	0.002	0.0002	6.9E-14	208.842		99.9	6031.10		1.9
2225-011	1050	80.813	0.004	0.0002	3.1E-14	128.577				1379.3	3.0
2225-01J	1100	80.982	0.003	0.0001			0.0000	99.9	7259.82	1391.4	1.9
2225-01K		80.594	0.002		3.5E-14	171.312	0.0000	99.9	8655.82	1393.5	2.1
2225-01L				0.0001	6.4E-14	284.048	0.0000	99.9	11223.83	1389.0	2.5
4420-012		81.461	0.003	0.0002	2.7E-14	170.285	0.0000		12304.12	1399.0	
		total gas age	n	=12	2.5E-15	0.010	0.0095		12004.12		2.0
						0.0.0	0.0035			1599	11
ML-6-10 E	Biotite, J	M, J=0.01455	5661	N	IM-19						
2226-01A	600	18.366	0.026	0.0064			2.0 mg				
2226-01B	650	53.155			2.2E-14	19.955	0.0041	89.6	432.30	387.6	1.0
2226-01C			0.009	0.0057	3.5E-15	57.022	0.0014	96.8	500.42	1008.5	1.6
	700	71.659	0.006	0.0035	1.2E-14	86.877	0.0008	98.5	735.96		
2226-01D	750	79.816	0.004	0.0013	1.7E-14	145.496	0.0004			1275.3	1.9
2226-01E	800	78.863	0.002	0.0006	1.9E-14			99.5	1061.45	1386.1	2.0
2226-01F	870	78.122	0.002	0.0005		211.151	0.0002	99.7	1436.83	1376.8	2.0
2226-01G	920	79.494			2.7E-14	223.305	0.0001	99.8	1970.93	1368.2	1.8
2226-01H	1000		0.005	0.0004	2.3E-14	94.187	0.0002	99.8	2418.11	1385.4	1.9
		81.068	0.026	0.0004	2.7E-14	19.380	0.0002	99.8	2948.97		
2226-011	1050	84.263	0.044	0.0003	3.4E-14	11.547	0.0005	99.9		1404.5	2.1
2226-01J	1100	83.068	0.016	0.0002	4.3E-14				3622.48	1442.7	1.9
2226-01K	1200	80.589	0.018	0.0002		31.982	0.0005	99.9	4453.05	1429.0	2.1
2226-01L	1700	80.044			7.3E-14	28.770	0.0003	99.9	5881.77	1399.3	2.5
			0.029	0.0003	4.5E-14	17.367	0.0002	99.9	6763.21	1392.5	2.0
		otal gas age	n=	:10	5.1E-15	0.030	0.0912			1334.0	9.9
III 40 C A	-4!111									1004.0	9.9
ML-10-6 A		JM, J=0.009	526213	N	M-29	0	.175 mg				
2496-01A	700	142.173	1.547	0.3054	2.4E-15	0.330					
2496-01B	800	89.215	0.972	0.0411			0.2238	36.6	#REF!	726.8	7.1
2496-01C	900	80.615			2.4E-15	0.525	0.0366	86.4	#REF!	994.3	2.2
2496-01D	1000		1.343	0.0234	3.5E-15	0.380	0.0295	91.5	#REF!	961.0	2.2
		91.517	6.365	0.0067	6.5E-15	0.080	0.0435	98.3	#REF!		
2496-01E	1040	107.001	9.430	0.0085	3.6E-15	0.054	0.0386			1120.7	2.2
2496-01F	1070	114.836	10.099	0.0072	4.9E-15	0.051		98.3	#REF!	1258.4	2.1
2496-01G	1100	89.367	6.322	0.0071			0.0300	98.8	#REF!	1328.7	2.4
2496-01H	1150	97.918			2.7E-15	0.081	0.0215	98.2	#REF!	1099.5	2.4
2496-011	1200		9.563	0.0084	2.5E-15	0.053	1.1475	98.2	#REF!	1178.8	2.3
		140.211	14.356	0.0123	3.2E-15	0.036	0.1558	98.2	#REF!		
2496-01J	1250	153.487	14.837	0.0099	7.6E-15	0.034	0.0396	98.8	#REF!	1521.9	2.3
2496-01K	1400	129.308	13.016	0.0196	1.0E-15	0.039				1624.1	2.4
2496-01L	1700	124.195	12.397	0.0402	6.1E-16		0.0347	96.3	#REF!	1419.8	3.5
	to	tal gas age	n=			0.041	0.0418	91.2	#REF!	1328.4	6.3
		J3-		12	4.1E-14	0.125	0.1689			1250.7	2.7
ML-13-9b H	fornhiend	e, JM, J≃0.0	00500754								
2497-01A	700				1-29	02	.08 mg				
		51.020	1.169	0.0522	2.9E-15	0.437	0.2639	69.9	10.45	E29.2	0.5
2497-01B	800	57.470	0.536	0.0093	2.4E-15	0.952	0.0230			528.2	2.5
2497-01C	900	57.895	1.152	0.0073	3.0E-15	0.443		95.2	19.13	757.5	1.8
2497-01D	1000	78.482	6.238	0.0073			0.0155	96.4	29.86	769.8	1.7
2497-01E	1040	98.503			4.6E-15	0.082	0.0254	98.4	46.21	998.5	1.7
2497-01F	1070		9.223	0.0068	2.9E-15	0.055	0.0314	98.7	56.40	1188.1	1.8
		101.205	8.681	0.0059	3.5E-15	0.059	0.0261	98.9	68.75		
2497-01G	1100	84.686	7.235	0.0065	1.8E-15	0.071	0.0305			1213.9	2.1
2497-01H	1150	94.529	10.708	0.0084	1.9E-15	0.048		98.4	75.30	1058.2	2.2
2497-011	1200	122.649	14.459	0.0123			1.1666	98.2	81.96	1149.8	2.2
2497-01J	1250	135.867	14.629		2.1E-15	0.035	0.3089	97.9	89.46	1386.0	2.3
2497-01K	1400			0.0104	2.5E-15	0.035	0.0553	98.6	98.24	1494.0	2.2
2497-01L		143.669	16.244	0.0326	3.5E-16	0.031	0.1156	94.1	99.48	1505.4	
2731-UIL	1700	159.663	15.096	0.1282	1.5E-16	0.034		77.0			6.2
	tot	al gas age	n=1	2	2.8E-14	0.216		, ,	100.00	1409	14
		_			17	0.210	0.2835			1044.8	2.1
ML-10-8 B	iotite, JM.	J=0.009510	19	NM	-20						
2694-01A	650	27.190	0.618	.			mg				
			. 0.010	0.0450	8.2E-16	0.825	1.1374	51.2	13.82	224.4	3.4

Run ID# 2694-01	Temp			³⁶ Ar/ ³⁹ Aı	r ³⁹ Ar moles	K/Ca	CI/K	% ⁴⁰ Ar	• 0/39•		
2694-01			0.217	0.0125	4.4E-16	2.346			· · · · · · · · · · · · · · · · · · ·	7190	± Err
	-		0.098	0.0069	7.6E-16	5.213					4.5
2694-01			0.117	0.0050						1149.3	3.3
2694-01		98.740	0.139	0.0048		4.349			42.56	1152.1	3.4
2694-01	F 1075	102.204	0.169			3.680		5 98.6	53.06		3.2
2694-01	G 1120		0.149	0.0162		3.026	0.135	7 95.3			
2694-01				0.0049		3.417	0.181				3.5
2694-01			0.035	0.0042	1.0E-15	14.678	0.096				2.7
2694-01		110.446	0.130	0.0210	2.5E-16	3.917					2.8
	-	118.180	2.617	0.1312	1.4E-17	0.195					7.1
2694-011		142.132	18.976	0.3753						1018	110
2694-01	L 1700	215.801	6.989	0.6773	4.9E-18	0.027		0 23.0	98.71		478
		total gas age			7.7E-17	0.073		7.5	100.00		71
		gar ago		n=12	5.9E-15	5.198	3.957	2		1042	5
ML-14-25		ite, JM, J=0	0.009518372		****						3
2695-01	700	81.615	1.127		NM-29		10.97 mg				
2695-01E		105.449		0.1416	7.7E-15	0.453	0.1879	48.8	4.85	580.3	• •
2695-010			0.596	0.0407	4.9E-15	0.856	0.0292		7.90		3.8
2695-01		98.849	0.599	0.0183	1.0E-14	0.852	0.0160			1148.2	2.5
2093-012		95.799	2.168	0.0037	1.7E-14	0.235			14.15	1148.5	1.9
2695-01E	_	107.558	5.923	0.0038	1.5E-14		0.0159	-	24.99	1161.9	1.6
2695-01F	1070	118.217	10.507	0.0065		0.086	0.0185		34.55	1269.7	1.8
2695-01G	1100	115.279	8.335		1.6E-14	0.049	0.0160	99.0	44.56	1357.8	2.0
2695-01H		110.003		0.0034	1.9E-14	0.061	0.0102	99.7	56.61		
2695-011	1200		4.182	0.0017	3.3E-14	0.122	0.0066			1338.5	1.8
2695-01J		108.410	10.660	0.0054	4.3E-15	0.048	0.2419		77.15	1293.4	2.0
	1250	124.606	13.274	0.0054	7.1E-15	0.038			79.83	1279.1	2.1
2695-01K		125.005	15.300	0.0060	2.5E-14		0.0284		84.31	1415.3	2.2
2695-01L	1750	132.289	11.341			0.033	0.0153	99.5	99.97	1419.8	2.0
	t	otal gas age		0.1947	4.0E-17	0.045	0.0351	57.2	100.00	984	
	•	gas age	n	=12	1.6E-13	0.181	0.3109			1265.5	50
ML-9-3b A	Actinolite	JM, J=0.00	0540074							1205.5	2.1
2698-01A	700				IM-29	1	1.58 mg				
2698-01B		29.747	10.117	0.0624	1.4E-15	0.050	24.1362	40.6	40.0-		
	800	58.128	13.688	0.0198	5.2E-16	0.037		40.6	12.95	197.9	3.7
2698-01C	900	73.490	4.837	0.0033	1.2E-15		19.2022	91.7	17.88	747.7	5.3
2698-01D	1000	76.039	6.619	0.0030		0.105	0.0761	99.1	28.94	954.6	2.7
2698-01E	1040	80.592	10.630		9.6E-16	0.077	0.0589	99.5	38.02	983.7	3.1
2698-01F	1070	90.373		0.0045	4.8E-16	0.048	0.0771	99.3	42.55	1029.6	
2698-01G	1100		15.910	0.0035	4.9E-16	0.032	0.0688	100.2	47.17		5.3
2698-01H		104.000	22.613	0.0103	4.7E-16	0.023	0.0603	98.7		1132.7	4.9
	1150	115.831	29.435	0.0137	2.0E-15	0.017			51.63	1246.4	3.9
2698-011	1200	91.510	5.387	0.0056	7.3E-16		0.0344	98.4	70.66	1347.8	2.6
2698-01J	1250	90.586	11.732	0.0068		0.095	1.5878	98.6	77.58	1123.7	3.3
2698-01K	1400	116.897	28.892		8.9E-16	0.043	0.6886	98.8	85.98	1120.1	2.8
2698-01L	1700	150.978		0.0141	1.4E-15	0.018	0.0379	98.3	99.34	1354.9	
_		tal gas age	19.148	0.2086	6.9E-17	0.027	0.0326	60.1	100.00		3.3
	•	nai yas age	n=	:12	1.1E-14	0.048	0.0297	••••	100.00	1137	30
ML-9-35 H	ornhlanda	215 1 2 2 2								1028.8	3.6
2699-01A	ombiende,	JM, J≃0.00	9543184	N!	M-29	3	.65 mg				
	700	21.469	3.575	0.0742	1.8E-15						
2699-01B	800	15.629	16.474	0.0471	5.8E-16	0.143	11.3050	-0.9	18.18	-3.4	3.6
2699-01C	900	40.399	64.990			0.031	41.3690	18.9	23.94	50.6	5.4
2699-01D	1000	70.169		0.0547	8.8E-16	0.008	7.7156	72.3	32.74		
2699-01E	1040		46.557	0.0260	5.8E-16	0.011	0.1675	94.1	38.56	462.1	4.7
2699-01F		64.297	60.550	0.0251	9.8E-17	0.008	0.2006			905.1	4.7
	1070	110.267	53.223	0.0280	3.1E-16	0.010		95.7	39.53	862	20
699-01G	1100	119.116	56.595	0.0271	7.1E-16		0.1031	96,2	42.63	1296.2	5.5
2699-01H	1150	121.295	45.028	0.0222		0.009	0.0658	96.9	49.71	1378.6	4.2
699-011	1200	100.091	78.445		1.4E-15	0.011	0.0511	97.4	63.45	1393.1	
699-01J	1250			0.0663	1.1E-16	0.007	13.6127	86.4	64.60		3.0
2699-01K	1400	118.535	56.231	0.0280	1.7E-15	0.009	0.1678	96.6		1133	14
2699-01L		126.383	65.025	0.0330	1.8E-15	0.008	0.0561		81.25	1370.8	3.0
099-01L	1700	147.024	55.577	0.2300	9.4E-17	0.009		96.2	99.06	1435.4	2.5
	tot	al gas age	n≖1				0.0549	56.7	100.00	1088	26
		-			1.0E-14	0.035	0.0386			940.3	4.1
IL-9-6 Acti	nolite, JM	l, J=0.00954	7693								
700-01A	700	36.685	2.792		-29		9 mg				
700-01B	800	35.386		0.0981	4.9E-16	0.183	4.8841	21.5	11.31	131.4	7.
700-01C			7.923		9.0E-17	0.064	3.2136	54.2			7.1
	900	20.975	1.887		3.0E-16	0.270			13.41	305	22
700-01D	1000	16.558	1.642		1.1E-15	0.270	0.1897	72.9	20.27	246.3	6.3
700-01E	1040	38.946	7.870				0.0405	93.1	45.73	248.1	1.7
700-01F	1070	56.988	11.237		3.4E-16	0.065	0.1118	95.5	53.75	550.9	4.6
700-01G	1100	80.481			2.6E-16	0.045	0.1481	97.0	59.89		
			17.239		2.4E-16	0.030	0.1270	98.0		769.8	5.3
		104.834	22.980		6.5E-16	0.022	0.0986			1021.8	6.5
700 041	1200	43.113	19.155		1.7E-16			98.6		1253.0	3.5
	1250	96.476	45.008			0.027	9.6192	86.1	84.27	553.4	9.1
'00-01J		121.708	62.118		4.2E-16	0.011	0.2262	95.7		1167.7	4.2
700-01J 700-01K	1400			0.0524	1.6E-16	0.008		91.2			7.4
700-01J 700-01K									97.77	1245 1	0.0
700-01J 700-01K	1700	148.641	20.455	0.1408	9.6E-17	0.025				1345.1	8.8
700-01J 700-01K	1700			0.1408	9.6E-17		0.0593		97.77 100.00	1297	19
700-01J 700-01K 700-01L	1700 tota	148.641 Igas age	20.455 n=12	0.1408	9.6E-17	0.025 0.136					
700-01J 700-01K 700-01L L-9-6 Horni	1700 tota blende, Ji	148.641 gas age M, J=0.00954	20.455 n=12	0.1408 2	9.6E-17 4.3E-15	0.136	0.0593 0.1055			1297	19
'00-01J '00-01K '00-01L -9-6 Hornl	1700 tota	148.641 gas age M, J=0.00954	20.455 n=12	0.1408 2 NM-	9.6E-17 4.3E-15	0.136 3.0 6	0.0593 0.1055 mg	73.1		1297	19
700-01J 700-01K 700-01L	1700 tota blende, Ji	148.641 gas age M, J=0.00954	20.455 n=12	0.1408 2 NM-	9.6E-17 4.3E-15	0.136 3.0 6	0.0593 0.1055 mg			1297	19

	n ID#	_	_														
	n ID# 01-01	В	remp 800	40Ar/39A	_		Ar/ ³⁹ Ar	39Ar me	nles	K							
27	01-01	С	900	60.63	_	854 (0.1101	2.3E	4.0	K/Ca	CI/I		%40Ar*	%	³⁹ Ar	A :	
27	01-01	_	000	64.02	,	072 c	0.0709	3.0E	4.4	0.087		300	47.0		.20	Age	<u>± Err</u>
	01-01	_	040	83.61			.0290	5.8E	4.0	0.025	0.2	669	69.7		4 -	436	13
	01-01	_		97.406	- 7.0	613 0	.0305	4.8E-		0.017	0.1	459	92.5	23.	_	47.7	8.2
	01-01		070	113.041	43.1		.0354			0.015	0.1	696	93.4			14.7	4.6
270	1-011		100	114.883	3 26.9		.0133	8.7E-		0.012	0.1	134	93.7	28.		49.3	5.1
270	11-011		150	89.353	27.3		.0183	1.2E-		.019		734	98.4	36.		88.5	4.5
270	11-011		200	96.801	30.0			4.0E-		.019	0.0			48.			3.5
270	1-01J	. 12	250	124.585	32.4	••	.0256	4.4E-	16 0	.017	3.7		96.3	52.			4.5
270	1-01K	14		133.384			0215	1.5E-	15 o	.016			94.5	56.	77 11:		4.4
270	1-01L	. 17	00	131.120			0218	2.6E-		.013	0.07		96.9	71.2		NE 0	
			tota	l gas ag	46.1	•.	0523	3.2E-1		011	0.07		97.4	96.8	36 148	14.0	3.5
						n≈12		1.0E-1			0.07		90.9	100.0	~~	0.0	2.5
ML-6	5-7 A	tinol	ito IN	J=0.00					٠٠ .	019	0.02	80				- -	7.1
2702	2-01A	7	AA	3=0.00	9554171		N:	M-29							114	7.0 E	5.7
2702	2-01B		00	54.953	9.50	04 0	1333			1.	11 mg						
2702	-01C		00	62.295	9.69		0439	6.7E-1		054	2.83		29.6	0.0			
			00	73.987	1.54			1.4E-1	6 0.	053	0.43			9.2		2.3 6	.7
2702		100	00	87.240	5.70	_	238	3.8E-1	6 0.;	331	0.15		30.3	11.2	0 7		0
2702		104		09.337			189	4.6E-16		089			0.6	16.4	7 894		
2702		107		15.428	11.40	_	096	4.4E-16		45	0.113		4.1	22.80	1047	· ·	
2702	-01G	110		17.746	11.95		100	3.8E-16			0.105		8.2	28.85		٠,	
2702	-01H	115		22 145	13.42	6 0.0	088	3.9E-16		43	0.114	_	8.2	34.16			
2702-		120		23.145	14.40	3 0.0	063	2.7E-15			0.088		8.6	39.49			
2702-		125	•	13.019	11.142			3.0E-16			0.018	3 9	9.4	76.11			
2702-	011			23.106	15.950		^ -		***	1	6.296		4.9				2
2702	011	140	U 12	24.821	17.450			9.3E-16		32	0.141	_	3.8	80.26			5
2702-	VIL.	170	0 16	32.123	17.279			3.7E-16		29	0.051	_		93.03		.0 g _. .	
			total	gas age				1.4E-16	0.0		0.032		7.5	98.13	1403.		
						n=12		7.2E-15	0.0				2.7 1	00.00	123		
ML-6-7	, Hori	blen	de, JM	J=0.00	05566				2.0.		0.084	I			1214.	_ ′-	
		700) · · · · · ·	9.134			NM-	29							· · · ·	0 4.0	
2703-0		800			2.343	0.07	15	I.0E-15	^ ^		mg						
2703-0)1C	900		9.938	1.112	0.02		5.0E-16	0.21		1.8463	46	.4	6.80	200	•	
2703-0	110		-	5.904	1.577	0.03			0.45		0.1803		_	10.17	289.		
2703-0	1.E	1000		6.077	6.158	0.01:	~ ~	1.9E-16	0.32	4 (0.0886				861.9		
2703-0		1040		0.309	11.702	0.009		.0E-15	0.08		0.0559	96.	_	6.14	783.9	3.0	
2700 -		1070	107	7.861	13.488		`	.0E-16	0.04		0.0646			3.05	1055.2	2.9	
2703-0		1100		3.187	17,774	0.009		.2E-16	0.03		0.0522	98.		7.72	1203.4	3.1	
2703-0	1H	1150		3.452		0.009		.8E-16	0.02			98.	-	2.53	1271.0	3.3	
2703-0	11	1200		.471	13.495	0.006	31 4.	1E-15	0.03		.0301	98.	-	9.10	1323.1	0.0	
2703-0	1J -	1250		.292	14.889	0.012		1E-15		.!	.0104	99.:	36	6.76	1368.6	0.0	
2703-01		1400			16.542	0.008		2E-15	0.034		.6507	97.7	-	4.39			. ບຸ
2703-01		700	124	.800	17.829	0.015		1E-15	0.031		.0345	99.0		9.31	1337.5	2.7	
	-		123	.372	18.864	0.030			0.029		.0196	97.5		5.59	1406.5	2.9	
			total ga	s age	n	=12		1E-16	0.027	0.	0247	93.8			1404.1	3.2	
881 44 4							1.3	5E-14	0.082		1435		100	0.00	1356.0	4.4	
2704.04	0-1 A	ctinol	ite, JM	, J=0.0	09558301										1211.6	2.8	
			43.	759	22.573	0.000	NM-29			1.99 n	na						
2704-01	R .	800		137	6.493	0.0836		E-15	0.023		79 5475	47 -					
2704-010	0	900	55.			0.0379		E-16	0.079			47.4		.20	331.4	3.8	
2704-011	D 1	000	88.		7.403	0.0196	1.6	E-16	0.069		9178	74.0	16	.53	462		
2704-01E	= 10	040			14.634	0.0151	2.7	E-16			4391	90.6		.37	709	13	
2704-01F		70	110.2		15.922	0.0183		E-16	0.035		2043	96.2	21.		1079.2	11	
704-010		00	118.6	089	16.146	0.0103		E-16	0.032	0.1	158	96.2	25.	17	1070.2	7.0	
704-01			123.6	521	16.026	0.0081			0.032		559	98.5	32.		1273.2	5.9	
704-011		50	112.0	181	14.876	0.0499		E-15	0.032		209	99.0			1364.0	4.5	
704-01J		00	79.7	30	14.665			E-16	0.034	0.1	289	87.8	64.		1408.8	2.2	
704-U]J		50	112.4	7.0	15.879	0.0188		-16	0.035		327		66.		1206	10	
704-01K			167.3	• •	19.082	0.0125	2.28		0.032		381	94.4	72.		985.9	3.8	
704-01L	17	00	200.0			0.0879	1.6E	-16	0.027			97.8	97.		1306.5	2.7	
			tal gas	200	16.549	0.3271	6.1E		0.027	0.2		85.3	99.2	29	1567		
					n=	12	8.6E			0.16		52.3	100.0		1262	11	
L-14-18-	1 Hor	nhle-	do	_			2.00		0.033	0.01	171		• •	-		33	
05-01A	ur	ייהופע	ue, JM,	J=0.00	09560309	1	NM-29								1151.9	4.0	
05-01A	• • •	, ,	39.77	/U 2	2.474	0.0920			1.	.75 mg							
OE 01-	80		56.62		7.568		8.6E		0.023	5.02		36.0	_	_			
05-01C	90	0	72.07		8.425	0.0665	2.3E		0.067	0.42			9.7		234.7	4.8	
05-01D	100	0	110.75			0.0358	2.4E-	16	0.028		`	56.3	12.3		556	11	
05-01E	104		122.67		8.371	0.0177	1.2E-	4-	0.028	0.21	. .	37.3	15.0	2	858.3		
05-01F	107	_		•	1.808	0.0218	8.5E-	• •		0.06		6.5	28.1		282.0	7.8	
05-01G	110	_	118.85		7.237	0.0160	8.0E-	4.0	0.023	0.05		6.1	37.8			3.3	
05-01H		_	111.95		5.408	0.0117			0.030	0.044		7.1	46.90		376.3	3.5	
05-011	115		101.56	7 17	7.301	0.0211	7.3E-		0.031	0.044		8.0			353.0	3.3	
05-01J	120		119.402	_			2.9E-		0.029	3.690		5.2	55.13		304.9	3.6	
22-017	125)	123.224	4	407	0.0127	8.1E-	16 (0.027	0.422	:		58.41		191.1	5.9	
5-01K	1400		118.256		656	0.0114	2.2E-1	-	0.028			8.0	67.57	13	367.5	3.5	
5-01L	1700		139.389			0.0211	5.3E-1	`	0.025	0.032		8.4	92.52		01.0		
			gas ag	17		0.1064	1.3E-1	_ `		0.035		6.0	98.57		40.1	2.6	
					n=12		8.8E-1		0.029	0.024		8.4	100.00			4.4	
	Anti-	01:4-	10-				0=-1	- 0	.028	0.011			00		1302	14	
14-18-2	ALL LIN	onte,	JM, J	=0.01070	469	B.2 B	A-42							12	05.7	4.0	
14-18-2 7-014	700					141	71-42		0.5	1 mg							
14-18-2 7-01A	700	8	43.093	4.	546	7020	F	_	0.5	ma							
14-18-2 7-01A	700	8	43.093	4.	546 2	2.7029	5.0E-1	7 0	.112	mg 0.3414	4 =	.3	7.54		707		

D. 10.											
Run ID#				r ³⁶ Ar/ ³⁹ /	r ³⁹ Ar mole	s K/C	`a 01/14	a. 40 .			
5897-0								% ⁴⁰ Ar			<u> </u>
5897-0			2.072					-	0 12.95	381	
5897-0		0 441.414							3 18.16		
5897-0		0 273.586							8 22.00		
5897-0	1F 104							13 16.1			
5897-0	1G 108							4 28.0			
5897-0	1H 112		15.181			*.*-	9 0.273				
5897-0						0.03	4 0.129			,	
5897-0			16.268		,	0.03					
5897-0			22.615	0.1607	7 1.3E-16	0.02	3 6.431				14
5897-0			19.696	0.0273	5.2E-17						16
5897-01			19.012	0.1709	5.8E-18					1459	21
3097-01	IM 1650		11.034	1.6074						1252	170
		total gas age)	n=13	6.7E-16				100.00	1680	900
•••					J 2 10	0.066	0.098	6		1106	59
ML 14-1	8-2 Horni	blende, JM,	J=0.01073805	5	NM-42						
2098-01	A 700	789.680	2.426	2.5600			1.14 mg				
5898-01			2.398	0.5035		0.210		2 4.2	1.54	553	340
5898-01	C 900	-46.877	2.845			0.213		22.9	3.74	700	
5898-01		,		0.4139		0.179	0.4169		5.23		62
5898-01			6.687	0.3417	4.0E-17	0.076				0	678
5898-01			11.438	0.1239	5.3E-17	0.045			7.12	811	51
5898-01			16.092	0.0867	5.5E-17	0.032			9.64	1146	23
			19.008	0.0568	1.0E-16	0.027			12.24	1274	21
5898-01			17.452	0.0153	4.1E-16	0.027			16.99	1318	12
5898-011		111.081	16.602	0.0131	6.0E-16	-			36.20	1400.8	5.6
5898-01,		120.104	16.698	0.0516		0.031			64.56	1404.6	4.1
5898-011		114.625	17.485		2.0E-16	0.031		88.4	74.06	1383.4	
5898-011	1300	114.689	17.057	0.0149	4.3E-16	0.029	0.2152	97.3	94.28	1432.8	7.4
5898-01N	M 1650	132.655		0.0176	8.8E-17	0.030	0.0417		98.46		4.4
		total gas age	16.298	0.0641	3.3E-17	0.031			100.00	1426	11
		total gas age	r	1=13	2.1E-15	0.040	0.0738	-0.0	100.00	1462	27
ML 14-18	-2 Bull A	mphibole, JM					**********			1335	24
5899-01	-2 DUIK A			225	NM-42		1.04 mg				
		116.042	1.880	0.3409	1.6E-16	0.271	•				
5899-018		49.496	2.191	0.0821	1.0E-16	0.233	0.1007	13.3	5.39	277	28
5899-010		63.137	5.774	0.0938	5.5E-17		0.1164	51.3	8.84	435	18
5899-010		62.269	12.311	0.0969		0.088	0.1670	56.8	10.71	589	26
5899-01E	1000	55.609	10.940		7.3E-17	0.041	0.2530	55.5	13.20	573	20
5899-01F	1040	70.979	17.278	0.0425	1.1E-16	0.047	0.1770	78.9	16.97	701	
5899-01G		96.805		0.0409	1.5E-16	0.030	0.1303	84.8	22.23	907.8	11
5899-01H		102.632	18.579	0.0204	3.2E-16	0.027	0.0746	95.2	33.12		9.3
5899-011	1160		16.323	0.0081	5.3E-16	0.03[0.0414	98.9		1252.6	5.4
5899-01J		91.321	15.008	0.0116	3.2E-16	0.034	0.1047		51.19	1339.7	4.3
	1200	99.966	17.774	0.0171	3.6E-16	0.029		97.5	62.18	1219.2	5.0
5899-01K		109.701	22.142	0.0040	2.5E-16	0.023	2.3456	96.3	74.52	1291.6	5.3
5899-01L	1300	117.435	23.322	0.0043	3.4E-16		0.0861	100.5	82.98	1423.6	5.6
5899-01M	1650	109.286	21.683	0.0276		0.022	0.0467	100.4	94.50	1492.2	4.5
	t	otal gas age		:13	1.6E-16	0.024	0.0370	94.0	100.00	1355.7	8.7
					2.9E-15	0.050	0.0834			1168.1	8.1
ML 10-8 E	Biotite, JN	I, J=0.010725	67								0.1
5913-01A	600	46.903	0.345		M-42	0	.68 mg				
5913-01B	670	28.295		0.1400	1.7E-16	1.477	0.1022	11.8	1.57	104	
5913-01C	730		0.207	0.0241	2.0E-16	2.470	0.1124	74.8	3.43	104	15
5913-01D	780	31.931	0.105	0.0936	1.4E-17	4.864	0.0216	13.3		369.1	6.5
5913-01E		80.665	0.310	0.0142	4.2E-16	1.646	0.0417		3.56	80	87
5913-01F	830	93.741	0.041	0.0072	5.8E-16	12.404		94.8	7.54	1080.8	3.9
	900	96.488	0.025	0.0020	1.0E-15	20.810	0.0216	97.7	12.99	1234.6	3.5
5913-01G	975	97.731	0.048	0.0030	8.4E-16		0.0134	99.3	22.75	1275.7	2.9
5913-01H	1050	96.378	0.101	0.0031	1.2E-15	10.644	0.0216	99.1	30.63	1284.9	2.9
5913-011	1150	96.304	0.182	0.0012		5.046	0.0204	99.0	41.51	1271.8	3.0
5913-01J	1250	98.008	0.141		3.7E-15	2.809	0.0089	99.6		1276.5	2.4
5913-01K	1600	94.150	0.274	0.0017	2.2E-15	3.624	0.4043	99.5		1291.4	
	to	tal gas age		0.0105	3.2E-16	1.862	0.0634	96.7		1229.3	2.6
		···· guo age	n=1	11	1.1E-14	6.020	6.0719				5.5
ML 10-8 A	ctinolite .	JM, J=0.0107	1001							1231.3	3.2
5914-01A	700	006 407			A-42	0.	16 mg				
5914-01B		906.107	2.511	1.6978	5.5E-17	0.203	0.4965	44.7			
5914-01B	800	108.302	3.749	0.1818	5.9E-17	0.136		44.7	7.04	3024	53
	900	113.951	0.412	0.1440	7.1E-17	1.239	0.2041	50.6	14.57	836	30
5914-01D	950	109.655	0.566	0.1253	6.2E-17		0.2525	62.7	23.72	1026	24
5914-01E	1000	105.462	0.786	0.1318		0.902	0.3395	66.2	31.71	1039	26
5914-01F	1040	107.064	1.377	0.0969	5.6E-17	0.650	0.3234	63.1	38.84	972	28
5914-01G	1080	102.349	2.784	0.0548	4.7E-17	0.370	0.3194	73.3	44.89	1102	28
5914-01H	1120	100.212	3.805		6.0E-17	0.183	0.2587	84.4	52.56	1184	
5914-011	1160	95.948		0.0301	9.5E-17	0.134	0.1480	91.4	64.72		24
5914-01J	1200	102.406	1.292	0.0131	1.6E-16	0.395	0.1505	96.0		1236	13
5914-01K	1250		6.693	0.0695	7.1E-17	0.076	12.8278	80.4		240.3	8.1
5914-01L	1300	82.821	15.311	0.0092	3.3E-17	0.033	0.6776	98.1	94.54	1146	17
5914-01M		514.857		-0.1500	2.4E-18	0.019			98.71	1139	32
2017-01W	1650	790.157	16.000	0.2030	7.7E-18	0.032		09.0	99.02	3543	286
	tota	al gas age	n=10		7.8E-16	0.400		92.6 1	00.00	3950	120
			•		0	V.700	0.3778			1288	24

	#	Temp	⁴⁰ Ar/ ³⁹ Ar	³⁷ Ar/ ³⁹ A	r ³⁶ Ar/ ³⁹ /	Ar mol	es K/	Ca CI	/K 0.	6 ⁴⁰ Αr*	30		
ML 10-	-8 Hor	nblend	e, JM, J=0	.01073843						• Ar	% ³⁹	Ar Age	<u>±1</u>
5916-0)1A	700	299.733	1.966		NM-42		0.60	ng				
5916-0)1B	800	129.470						8792	6.0	4.6	32 040	
5916-0	1C	900	103.968	0.663					3027	50.6			
5916-0		950	80.150	0.866			6 0.5		1754	72.7	8.9		
5916-0		000		1.471					1843	80.8	13.4	- · · · -	
5916-0		040	79.267	3.037			7 0.10				17.5		1
5916-0			91.236	6.654		5 1.9E-1				91.3	20.7		1
5916-0		080	104.210	12.123	0.0113					93.9	27.0	7 1180.7	7.
		120	101.921	8.478	0.0060	4.3E-1				97.7	42.1	0 1340.5	4.
5916-0		160	96.116	6.048	0.0094					98.9	56.8		4.
5916-0	1J 1	200	108.762	12.483	0.0193					97.6	65.5		
5916-0	1K 1	250	102.001	12.134	0.0051				290	95.6	74.9		6.
5916-0		300	83.257	9.690				2_ 0.0	510	99.4	88.9		6.
5916-01	1M 1	650	73.382		0.0063		0.05	3 0.0		98.6	98.6		5.
			tal gas age	3.676	0.0357	4.0E-17	0.13			36.0			5.
		.0	tai yas age		n=13	2.9E-15		_ '	329	30.0	100.00		28
ML 10-8	-2 But	k Amn	hibole, JM,					- V.L	025			1197.5	9.8
5917-01	ΙΔ -	700	704 CO	J=0.010772	295	NM-42		0.42 m	-				
5917-01	_		761.223	3.034	2.5152	3.6E-17	0.16						
	_ '	300	122.082	1.126	0.2275	5.3E-17				2.4	2.61	324	502
5917-01		900	145.314	1.847	0.2097			_		5.0	6.43	839	45
5917-01	_	50	167.947	3.098	0.2956	4.9E-17				7.4	9.96		36
5917-01		000	133.332	5.103		5.6E-17			335 4	8.1	14.02		
5917-01		40	123.110	9.924	0.1688	6.8E-17	0.100	0.35		2.9	18.92		40
5917-010		80	111.420		0.1030	1.1E-16	0.051			5.9		1164	25
5917-011		20		15.208	0.0316	3.1E-16	0.034			2.6	27.19	1262	14
5917-011		60	102.533	7.705	0.0198	3.4E-16	0.066				49.63	1359.1	6.1
5917-01.			118.889	11.244	0.0704	1.6E-16	0.045			4.8	73.85	1297.9	5.2
	-		110.547	17.433	0.0027	1.0E-16				3.2	85.62	1316	11
5917-01			119.649	22.435	0.0035	2.9E-17	0.029				92.99	1432	13
5917-01			110.106	21.754	0.0248		0.023		70 100	0.6	95.05	1516	35
5917-01N	√1 16	50	149.673	23.523		3.2E-17	0.023		91 94	1.8	97.38	1375	
		tota	l gas age		0.0663	3.6E-17	0.022	0.06	50 88	3.1	100.00		33
			3-0 090	n	=13	1.4E-15	0.080	0.12			.00.00	1612	28
P-1-2b:a	Musc	ovite	JM, J=0.01						- •			1274	27
919-01A	4 6/	00			1	VM-42		0.99 mg					
919-01B			48.736	0.097	0.1145	9.0E-16	5.249	_		_			
			78.456	0.071	0.0203	8.9E-16		0.01			1.76	266.4	6.5
919-01C		0 .	101.548	0.035	0.0146	2.2E-16	7.219	0.006		.3	3.50	1034.0	3.1
919-01D		0 :	107.213	0.022	0.0084		14.587	0.005	8 95	.7	3.94	1284.9	
919-01E		0 1	06.445	0.010		2.0E-15	23.539	0.004	0 97	.7	7.78	1354.6	6.2
919-01F	82	_	04.617	0.007	0.0044	3.4E-15	52.224	0.002			14.45		2.9
919-01G		_	03.968		0.0032	3.7E-15	73.796	0.002			21.60	1358.3	2.8
919-01H	87	_	03.827	0.006	0.0032	3.8E-15	82.545	0.002				1344.8	2.6
919-011	90	_		0.005	0.0028	3.9E-15	101.149	0.002			29.10	1338.9	2.6
919-01J			04.239	0.004	0.0027	3.7E-15	115.313				36.77	1338.8	2.6
919-01K	92		04.742	0.004	0.0024	2.8E-15	117.528	0.002			43.92	1342.8	2.7
	95		05.554	0.006	0.0028	3.6E-15		0.002			49.37	1348.3	2.7
919-01L	97		05.303	0.005	0.0033	2.3E-15	92.270	0.003		2	56.34	1354.5	2.7
919-01M	100		04.983	0.001	0.0042		107.783	0.004	5 99.	1	60.77	1350.9	
919-01N	104		05.411	0.004		1.7E-15	470.076	0.005	6 98.	_	64.11	1345.4	2.8
19-010	108		05.722	0.006	0.0041	2.2E-15	113.839	0.005	98.		68.45		2.9
19-01P	1120		06.017		0.0027	2.9E-15	86.053	0.005		_	74.14	1349.6	2.8
19-01Q	1180			0.008	0.0019	3.6E-15	64.141	0.005				1356.4	2.8
19-01R	1220		06.872	0.020	0.0029	7.6E-15	25.846	0.1326			81.18	1361.1	2.5
19-015			06.602	0.076	0.0007	1.5E-15	6.671				96.03	1366.1	2.7
19-015	1400	, 11	3.210	0.833	0.0020	4.3E-16	0.612	0.0118			99.01	1369.6	3.1
19-011	1600	15	5.553	0.280	0.0140	7.0E-17		0.0077			99.86	1426.1	5.1
		total	gas age	n=2			1.820	0.0157	97.9	9 10	00.00	2009	21
						5.1E-14	83.821	101.9388			-	1329.0	
-1-2b:b I	Musco	/ite, JA	/i, J=0.010	68425								.020.0	2.8
20-01A	600	1	5.370	0.138		1-42	1.	2 mg					
20-01B	650		3.933		0.0388	1.7E-16	3.690	0.0142	25.3		0.36	- .	
20-01C	700	-		0.083	0.0050	1.3E-15	6.132	0.0075				74	10
20-01D	750		8.939	0.074	0.0039	2.7E-16	6.904	0.0075			3.03	536.7	1.9
20-01E			1.401	0.065	0.0011	2.0E-15	7.811		98.5		3.59	1091.3	5.0
	800		6.118	0.011	0.0008	2.6E-15		0.0038	99.7		7.77	1321.0	2.6
20-01F	825		4.929	0.009	0.0005		45.224	0.0013	99.8	1	3.29	1365.0	2.5
0-01G	850	10	4.894	0.008		2.8E-15	54.568	0.0008	99.8		9.22	1355.0	
10-01H	875		4.115	0.004		3.6E-15	62.172	0.0009	99.8		6.72		2.4
0-011	900		4.324			3.7E-15 1	14.842	0.0006	99.9			1354.4	2.6
0-01J	925			0.004			20.351	0.0013				1348.4	2.5
0-01K			4.100	0.006	0.0001		89.879		99.9			1349.8	2.5
0-01K	950		1.904	0.007				0.0015	100.0	4 (6.61	1348.5	2.5
	975		5.321	0.007			71.924	0.0016	99.9	53		1355.3	2.4
0-01M	1000	105	5.319				71.127	0.0024	100.0			1359.6	
0-01N	1040						78.183	0.0048	99.9		2.31	1359	2.3
0-010	1080						57.431	0.0045	99.9				15
.	1120				0.0005		56.329	0.0037				1354.8	2.5
	1180			0.006	0.0001	- . -	86.860		99.8			1358.7	2.8
J-0771	1100	106	.680		-			0.0030	99.9		3.68	1367.9	2.3
_	1000												
_	1220	106	.702		`		56.673 12.232	0.0863 0.0058	99.9 99.9	95	.92	1371.1	2.5

Run ID		emp 40Ar/39		r ³⁶ Ar/ ³⁹ Aı	³⁹ Ar mole	esK/C	`	40			
5920-0 5920-0		00 109.72			2.8E-1	6 0.57					± Err
0320-0	,,, 16	119.73	9 0.694								5.8
		total gas a	ge	n=20	4.8E-1				7 100.0		13
JM-2 N	/uscovite	M Long					. 00.43	700		1329.9	3.0
5921-0					NM-42		0.51 mg				
5921-0				0.5468	1.0E-16	2.63				_	
5921-0				0.0239	2.9E-16						49
5921-0				0.0006	7.4E-17						5.7
5921-0	15 0	50 69.052		0.0077	8.5E-16						18
		00 84.414		0.0051	1.1E-15					8 973.1	2.9
5921-0		25 92.190	0.014	0.0034	1.3E-15					1146.1	3.2
5921-01		50 95.485	0.008	0.0024	1.3E-15		•		20.83	1228.8	2.8
5921-0		75 97.821	0.006	0.0012		+4.000			28.08		2.9
5921-0		00 98.257		0.0025	1.3E-15				35.14		3.0
5921-01		25 98.910		0.0025	1.2E-15		0.000	55 99.2	41.64	1289.6	3.2
5921-01		0 99.260		0.0035	8.7E-16			99.2			
5921-01		5 100.156	0.032	0.0035	1.5E-15			76 98.9			3.2
5921-01	M 100				1.1E-15	16.094	0.009	99.0	61.15		3.0
5921-01	N 104			0.0045	1.1E-15	26.847			67.42		3.6
5921-01	0 108			0.0053	1.4E-15	33.331	0.010		75.13		3.3
5921-01			0.011	0.0079	1.1E-15	44.849			81.39		2.8
5921-01			0.012	0.0070	1.1E-15	41.863	0.018		87.34		3.4
5921-01			0.037	0.0205	8.1E-16	13.901	1.167			1331.3	3.2
5921-01			0.030	0.0338	3.7E-16	17.007	0.021		91.89	1320.9	3.4
5921-01			0.067	0.0045	9.3E-16	7.598	0.005		93.98	1317.0	5.6
	. 160		0.032	0.0238	1.4E-16	15.875	0.005		99.22	1302.3	3.1
		total gas age	n	=20	1.8E-14	36.338	22.877	3 93.4	100.00	1312.4	9.6
.IM-A.a.	luna	•••			- • •	-0.000	-4.0//	1		1251.8	3.6
5000 A	nuscovite	, JM, J=0.010	0612	N	M-42		1 50				-
5923-01			0.022	0.1121	1.5E-15	23.068	1.52 mg	_			
5923-018	8 650		0.016	0.0114	1.7E-15	31.202	0.0082		2.94	549.7	5.0
5923-010			0.010	0.0068	1.7E-15		0.0028		6.11	987.5	2.5
5923-01		89.981	0.007	0.0039		50.697	0.0022	•	9.42	1126.9	2.7
5923-01E		95.485	0.007	0.0039	3.9E-15	71.054	0.0015	98.7	16.92	1197.8	
5923-01F			0.005		5.7E-15	76.642	0.0010	99.0	27.79	1253.2	2.4
5923-016	850			0.0032	4.9E-15	94.071	0.0013		37.28		2.5
5923-01⊦	H 875		0.005	0.0054	5.3E-15	111.978	0.0016		47.44	1265.4	2.5
5923-011			0.004	0.0071	4.3E-15	123.744	0.0019		55.79	1270.0	2.5
5923-01J			0.006	0.0085	3.6E-15	85.678	0.0026			1271.3	2.5
5923-01K			0.006	0.0095	2.7E-15	88.523	0.0039		62.66	1280.7	2.5
5923-01L			0.006	0.0127	2.3E-15	82.451	0.0055		67.92	1284.3	3.0
			0.006	0.0119	2.4E-15	91.232			72.32	1296.9	3.0
5923-01M		102.464	0.007	0.0096	2.2E-15		0.0061	96.5	76.94	1293.7	2.9
5923-01N		102.796	0.008	0.0083	2.4E-15	70.432	0.0065	97.2	81.19	1301.2	2.7
5923-010		103.784	0.006	0.0108	2.4E-15	61.943	0.0071	97.6	85.82	1307.9	2.8
5923-01P		103.626	0.007	0.0057		82.585	0.0072	96.9	90.43	1310.0	2.9
5923-01Q	1180	108.777	0.011		2.9E-15	76.759	0.0095	98.3	95.98	1322.6	
5923-01R	1220	105.635	0.021	0.0181	1.8E-15	45.566	0.5933	95.1	99.44	1336.2	2.6
5923-01S	1400	117.401		0.0060	1.3E-16	24.039	0.0645	98.3	99.69		2.8
5923-01T	1600	118.972	0.022	0.0295	6.4E-17	22.945	0.0382	92.5	99.82	1340	10
		total gas age	0.049	0.0054	9.6E-17	10.332	0.0275	98.6	100.00	1383	20
		cotta gas age	n=:	20	5.2E-14	81.639	31.6409	00.0	100.00	1459	15
JM-4:b Mir	scovita	JM, J=0.0106								1242.1	2.8
5924-01A	600	vm, u=0.0106		NM	-42	۵.:	8 mg				
5924-01B	650	44.445	0.091	0.0468	3.8E-16	5.591	0.0302	60 0	o o-		
5924-01C	700	88.394	0.000	0.0000	6.8E-17	0.000	0.0045	68.8	2.37	507.2	4.8
5924-01D		90.311	0.023	0.0047	2.6E-16	22.482	0.0133	100.0	2.79	1193	17
5924-01E	750	94.229	0.016		1.1E-15	32.293		98.4	4.43	1198.8	5.1
5924-01E	800	96.280	0.010		1.3E-15	51.854	0.0062	99.3	11.53	1244.4	2.6
	825	95.237	0.008		1.4E-15	61.098	0.0039	99.5	19.71	1265.5	2.8
5924-01G	850	94.790	0.006		1.5E-15		0.0031	99.7	28.53	1257.6	2.5
5924-01H	875	95.200	0.013		1.1E-15	87.279	0.0027	99.7	38.19	1252.8	2.5
5924-011	900	95.418	0.008			39.549	0.0039	99.7	45.20	1256.9	2.7
5924-01J	925	95.780	0.010		1.0E-15	66.897	0.0041	99.7	51.46	1259.1	2.7
5924-01K	950	95.881	0.006	.		49.757	0.0059	99.5	57.93	1261.1	
5924-01L	975	96.782	0.006			92.648	0.0077	99.7	61.58	1263.5	2.9
5924-01M	1000	95.875	0.005			78.922	0.0099	99.6		1271.5	3.8
5924-01N	1040	96.557				08.532	0.0093	99.5			3.1
5924-010	1080	97.914	0.008		5.2E-16	61.936	0.0097	99.8		1261.5	3.1
5924-01P	1120	99.725	0.004		I.1E-15 1	33.662	0.0115	99.7		1270.4	3.2
5924-01Q	1180				4	39.790	0.0102	99.5		1282.6	3.0
5924-01R	1220	102.626				30.834	0.4736			1297.7	2.7
5924-01S		112.124			.3E-16	8.016	0.4736	99.2		1321.3	2.7
5924-015	1400	274.399			.9E-18	0.249		99.4		1408.1	9.8
2027-011	1600	329.102			.3E-18	0.414		106.3	99.95	2546	147
	to	tal gas age	n=20				0.1073	71.2 1	00.00	2255	176
IM.E. Marie		_		•	((,,,oso 3	88.2551		•	1252.8	3.2
JM-5 Musco				NM-4	12						–
5925-01A	600	45.754	0.021	-			mg				
				1.	.uc-10 2	24.429	0.0085	75.0	3.89	564.5	3.3
										- · · -	2.0

Run I		_Temp	40Ar/39A	r ³⁷ Ar/ ³¹). ne											
5925-		650	72.00				Ar mole	s k	∕/Ca	CI/K	% ⁴⁰ A	r* n/	39 .			
5925-		700	76.87			034	1.5E-1	5 27.	234	0.0047	98.		³⁹ Ar	Age	±	Err
5925-	01D	750	84.59			016	1.4E-1		695	0.0034				1019.9		2.5
5925-		800	96.642		4.0		4.2E-1			0.0017				1078.4		2.6
5925-		825	102.239			310	4.6E-15		986	0.0009				1159.4	:	2.5
5925-(01 G	850	103.739				3.6E-15		289	0.0006				1278.6	- 2	2.3
5925-0	91H	875	104.657	0.00	4.00		2.6E-15		371	0.0013			_ ·	330.4	2	2.6
5925-(900	106.477				2.0E-15			0.0015				343.7		2.6
5925-0		925	107.872	0.00			1.6E-15		73	0.0024	98.9			349.4		.7
5925-0)1K	950	110.089	0.00			1.1E-15			0.0048	98.7		'	362.5		.8
5925-0		975	110.441	0.00			5.4E-16			0.0073	99.0			373.0 395.4		.0
5925-0		1000	111.080	0.00			5.4E-16	98.0	18 (0.0096	98.1			389.8		.4
5925-0 5925 - 0	10	1040	111.660	0.00			4E 40	6363.1		0.0145	99.0			404.4		.6
5925-0		1080	112.603	0.00			4E-16	0.0		0.0200	98.4			403.0		.3
5925-0		1120	111.547	0.01	5 0.01:		.1E-16 .7E-16	293.1		0.0438	95.9			386.6		.3
5925-0		1180	112.836	0.00			.7E-16	34.5		.0694	96.6			384.1	7.	
5925-0		1220	143.642	0.045	5 0.110		.6E-17	56.11		.3650	96.1	99.7		390.0	9.	
5925-0		400	102.917	0.006			.0E-17 .2E-17	11.38		.0465	77.2			1413	6.	
5525-0	, , ,	600	403.413	0.111		_	.8E-18	90.50	_	.0523	106.3	99.9		1399	3	
		to	tal gas age		n=20	•	6E-14	4.59	3 0	.1412	61.6	100.0		2340	5	
JM-7-0	Mussa					۷.	02-14	203.90	5 1406	.2543			-	46.2	36	
JM-7:a 5927-01						NM-4	2						, 2		3.	U
5927-01	_	600 650	42.468	0.026	0.073		4E-15	10.00	1.05	-						
5927-01		650 700	85.891	0.017		•	2E-15	19.38		0055	48.7	8.4	9 3	61.0	2.6	
5927-01	_	700 750	99.931	0.008	0.008		3E-15	30.77	_	0018	95.4	13.9		37.2	3.5 2.7	
5927-01	_	750 300	103.409	0.007	0.006	_	2E-15	66.50 72.99		0017	97.5	19.80		39.5	2.7	
5927-01	_	325	104.241	0.005	0.004		9E-15	93.363		0015	98.3	30.25		28.7	2.7	
5927-01		350	104.132	0.005	0.0046		E-15	99.991		0013	98.6	42.34		10.0	2.5	
5927-011		375	105.054	0.005	0.0093		E-15	100.572		0013	98.7	51.90		9.1	2.6	
5927-011	_	000	108.355	0.004	0.0192		E-15	120.774		0017	97.4	60.63	133		2.6	
5927-01.			109.366	0.007	0.0229		E-15	68.578		0027	94.7	68.54	133		2.8	
5927-011			109.792	0.004	0.0239		E-15	126.536		0042	93.8	74.59	133		2.8	
5927-011			111.539	0.007	0.0277		E-15	73.442		052	93.6	79.28	133		3.0	
5927-01N			111.454	0.005	0.0250		E-15	96.735		073	92.6	82.69	134	4.4	3.1	
5927-01N	1 10		110.138	0.005	0.0219		_	106.421		083	93.4	86.65	135	1.1	3.3	
5927-01C			109.641 108.804	0.006	0.0185		E-15	78.959		094	94.1	90.12	134	7.3	3.4	
5927-01P				0.009	0.0172		E-15	55.435			95.0	93.64	1352	2.0	3.2	
5927-01Q			110.380 129.805	0.013	0.0212		≅-16	40.646			95.3	96.79	1348	3.0	3.2	
5927-01R			109.176	0.032	0.0890	3.78		15.711			94.3	98.96	1351		3.3	
5927-01S	140		75.026	0.000	0.0029	2.45	-17	0.000	2.39 0.22		79.7	99.88	1345	5.9	7.4	
5927-01T	160		10.155	0.000	0.0179	1.95	-17	0.000	0.16		99.2	99.94	13:	90	39	
			gas age	0.077	0.4816	4.2E	-18	6.662	0.20		97.0	99.99	181	69	53	
					=20	4.0E	-14	79.685	40.68		37.2	100.00	438	36	305	
JM-7:b ML	scovit	e, JM.	J=0.01071	1340					40.00	,00			1243	.2	3.0	
3320-01A	60	0	28.745	0.142		NM-42			1.1 mg							
5928-01B	65		68.359		0.0546	1.5E	-16	3.590	0.05	56 4	3.8					
5928-01C	70		87.149	0.056	0.0183	1.4E	-16	9.103	0.08		2.1	3.99	22		12	
5928-01D	75		97.683	0.103	0.0092	8.6E	-17	4.960	0.05		6.9	7.69	929.		8.3	
5928-01E	80		99.750	0.021 0.017	0.0085	3.3E	16	24.026	0.02	_	7.4	10.01	116		12	
5928-01F	82		99.782	0.000	0.0040	4.3E-	16	30.017	0.01		7. 4 8.8	18.86	1267.		5.0	
5928-01G	850		1.062	0.010	0.0013	4.3E-		0.000	0.01		9.6	30.44	1300.		3.7	
5928-01H	875	5 10	0.304	0.004	0.0018	3.8E-		52.939	0.01	-	9.4	42.17	1308.		3.6	
5928-011	900	10	1.700	0.024	-0.0002	3.2E•		44.909	0.017			52.39	1318.		4.3	
5928-01J	925		0.838	0.007	0.0012	2.6E-		21.192	0.025		9.6	60.91	1316.	7	4.4	
5928-01K	950		2.952	0.008	0.0062	2.3E-		9.297	0.035		3.1	67.82	1326.0		4.7	
5928-01L	975	10	2.902	0.000	0.0008	1.4E-		1.990	0.040		.8	74.01	1304.		5.2	
5928-01M	1000	10	2.245	0.024	-0.0025	1.7E-		0.000	0.047			77.92	1338.7		7.9	
5928-01N	1040		2.658	0.000	-0.0032	1.5E-1	_	0.966	0.047		_	82.46 86.46	1347.2	?	7.5	
5928-010	1080		4.225	0.023	0.0034	1.7E-1		0.000	0.052		_	91.08	1342.9		7.5	
5928-01P	1120	10:	2.763	0.024	0.0007	1.5E-1		2.074	0.064				1328.8		7.9	
5928-01Q	1180	11:	3.820	0.292	0.0051	1.1E-1		1.047	0.115		_	95.24 98.28	1350.6		8.1	
5928-01R	1220		2.697	0.000	0.0525	3.5E-1	7	1.749	20.287		_		1325		11	
5928-01S	1400	22	1.693	0.000	-0.2023 0.0015	2.7E-1		0.000	2.975			99.23 99.30	1298		29	
5928-01T	1650	152	2.914	1.144		2.8E-1		0.000	0.602		_	99.30 99.38	1747	_	40	
		total g	as age	n=2	0.0750	2.3E-1	_ '	0.446	0.0142		_ '	99.38	2192	_	35	
.114 0					•	3.7E-1	5 34	1.553	35.5000		- "	-0.00	1581		43	
JM-8:a Musc	ovite,	JM, J	=0.0106228	7									1256.1	•	6.7	
0023-01A	600	47	.326	0.050	0.1302	1-42	_		5 mg							
5929-01B 5929-01C	650		.402	0.068	0.1302	1.1E-15		.295	0.0088	18.	6	5.94	101 0			
	700		.157	0.129	0.0250	5.6E-16		.475	0.0035			8.92	161.6		5.9	
5929-01D 5929-01E	750		.235	0.017	0.0250	1.8E-16		.945	0.0016				897.0		1.4	
5929-01E 5929-01F	800	105	488		0.0063	1.4E-15		.684	0.0015		_	_	1178.3		5.9	
5929-01F	825		737	•		2.3E-15	. •	.798	0.0006	98.2		A	1283.9	2	.9	
032010	850	104.				2.3E-15		.897	0.0005	98.5	_		1339.0 1335.2		.9	
			ě		0	2.2E-15	110	.130	0.0009	98.€			1335.2		.7	
												-		3	.0	

Run ID#		D 40Ar/39Ar	27 20								
5929-0	1H 87	5 104.988	7.117		39Ar mole	<u>s </u>					
5929-0 ⁻	11 90	- 107.300	0.006		2.1E-1	5 91.43	5/11	% ⁴⁰ A	^ ^ A	rAge	
5929-01	ال ا		0.003	0.0153	1.8E-15	- 01,70			- 07.01	1328.6	<u>± Err</u>
5929-01	K 95		0.008	0.0265	1.4E-15					1332.6	2.8
5929-01	L 97		0.000	0.0355	5.5E-16				81.13	1332.2	2.9
5929-01	M 1000		0.003	0.0447	8.2E-16	0.000			84.06		3.5
5929-01	N 1040		0.002	0.0469	5.6E-16		,				4.3
5929-01	0 1080	122.000	0.004	0.0517	5.4E-16				91.38	1357.3	4.0
5929-01	P 1120	. ~ ~ / 0	0.000	0.0523	5.1E-16				94.26	1370.0	4.7
5929-01	Q 1180	. ~ 7 . 0 3 1	0.014	0.0630	3.8E-16	0.000	,		96.97		4.8
5929-018	7 1220	-00.437	0.036	0.2534	1.7E-16	90.004	V. 0 T 0		98.98	1373.3	5.4
5929-018		,00.255	0.000	0.0412	7.7E-18	14.279		0 62.6	99.86	1389.3	5.5
5929-017		332.604	0.295	0.1183	1.1E-17	0.000	0.530	4 93.5	99.90	1529	15
		631.277	0.000	0.2919		1.731	0.224	0 89.5	99.96	1902	122
		total gas age	n	=20	7.8E-18	0.000	0.143	2 86.3	100.00	2573	87
JM-8:b M	(1000				1.9E-14	83.448	63.835		.00.00	3455	156
5930-01A	uscovite,		2287		iM-42					1254.0	3.8
5930-01B		18.225	0.112	0.0150			1.05 mg				
5930-01B		66.305	0.091	0.0082	6.8E-16	4.549	0.0248	75.6	7.91		
5930-010		84.243	0.004	0.0070	3.3E-16	5.623	0.0378			246.5	2.3
5930-01E		93.114	0.015	0.0070	3.0E-16	125.116	0.0244		11.73	934.3	4.5
5930-01E	800	100.132	0.008	0.0037	6.9E-16	34.520	0.0143		15.21	1131.8	4.3
5030-015		101.207	0.007	0.0024	1.0E-15	66.513	0.0090		23.29	1230.0	3.1
5930-01G	850	100.869	0.000	0.0014	9.5E-16	76.977	0.0062	99.6	35.03	1300.3	2.8
5930-01H	875	100.735	0.010	0.0020	8.3E-16	0.000	0.0071	99.4	46.09	1312.8	3.0
5930-011	900	100.893	0.004	0.0010	7.2E-16	49.270	0.0078	99.4	55.77	1308.0	3.0
5930-01J	925	100.532	0.004	0.0025	5.7E-16	139.496	0.0111		64.15	1309.7	3.2
5930-01K	950	98.861	0.009	0.0023	4.8E-16	86.139	0.0163	99.2	70.75	1307.0	3.2
5930-01L	975	99.866	0.009	0.0017	3.5E-16	59.013	0.0251	99.3	76.38	1304.3	3.7
5930-01M	1000	99.952		0.0000	3.4E-16	56.835	0.0231	99.5	80.42	1290.3	4.7
5930-01N	1040	101.756	0.000	-0.0023	3.0E-16	0.000	0.0243	100.0	84.40	1304.4	4.3
5930-010	1080	103.979	0.000	-0.0025	3.5E-16	0.000		100.7	87.89	1311.5	4.8
5930-01P	1120	103.074	0.016	0.0022	3.2E-16	31.880	0.0283	100.7	91.91	1328.7	4.2
5930-01Q	1180	105.387	0.037	-0.0002	2.4E-16	13.675	0.0431	99.4	95.61	1336.3	4.3
5930-01R	1220	86.224	0.231	0.0107	7.6E-17	2.210	0.1913	100.0	98.40	1334.4	5.0
5930-018	1400	94.456	0.000		1.1E-17	0.000	10.2346	97.0	99.29	1326	
5930-01T	1650	127.759		-0.0119	2.1E-17	1.660		129.1	99.41	1408	12
	-	127.759	0.936		3.0E-17			103.7	99.65	1287	60
	.0(al gas age	n=2	•		0.545	0.2507	98.3	100.00	1530	35
				•		48.465	4.4E+01			1199.7	29
										33.7	3.8

D= 1.00510 0.00200 Ca 39/37= 0.00070 0.00005 Ca 36/37= 0.00026 0.00002 K 38/39= 0.01190 K 40/39= 0.02700 0.00200 This thesis is accepted on behalf of the faculty of the Institute by the following committee:

And to
Advisor
Cl. Develin
Mate
Date
I release this document to the New Mexico Institute of Mining and Technology.
AMM 1 127/90
Student's Signature / Date